

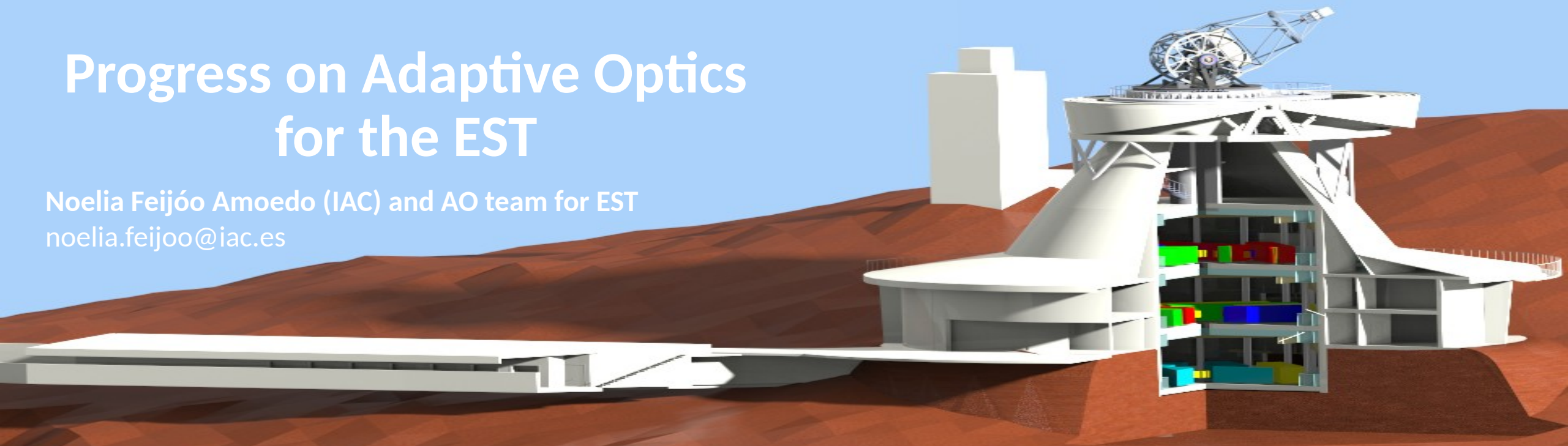


european solar telescope



# Progress on Adaptive Optics for the EST

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Solarnet -S3



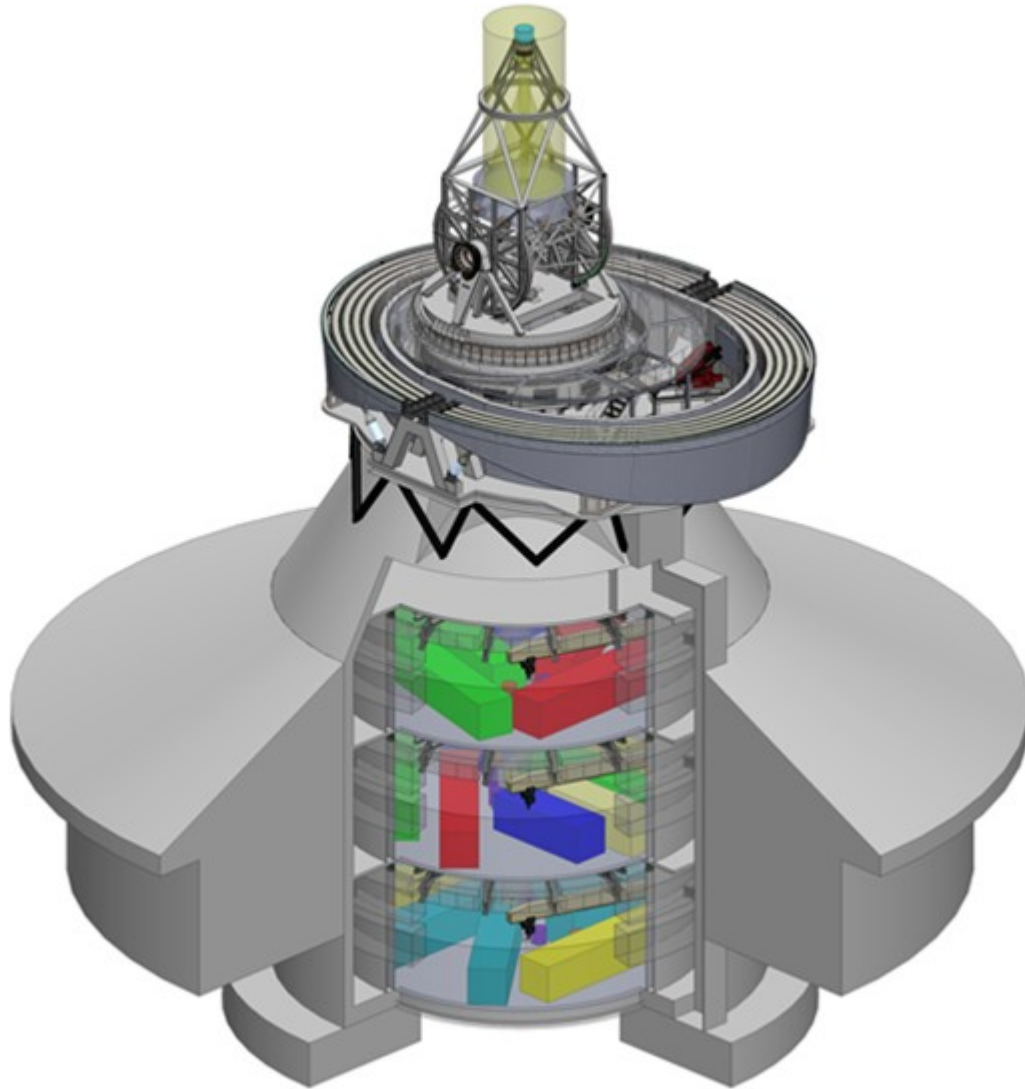
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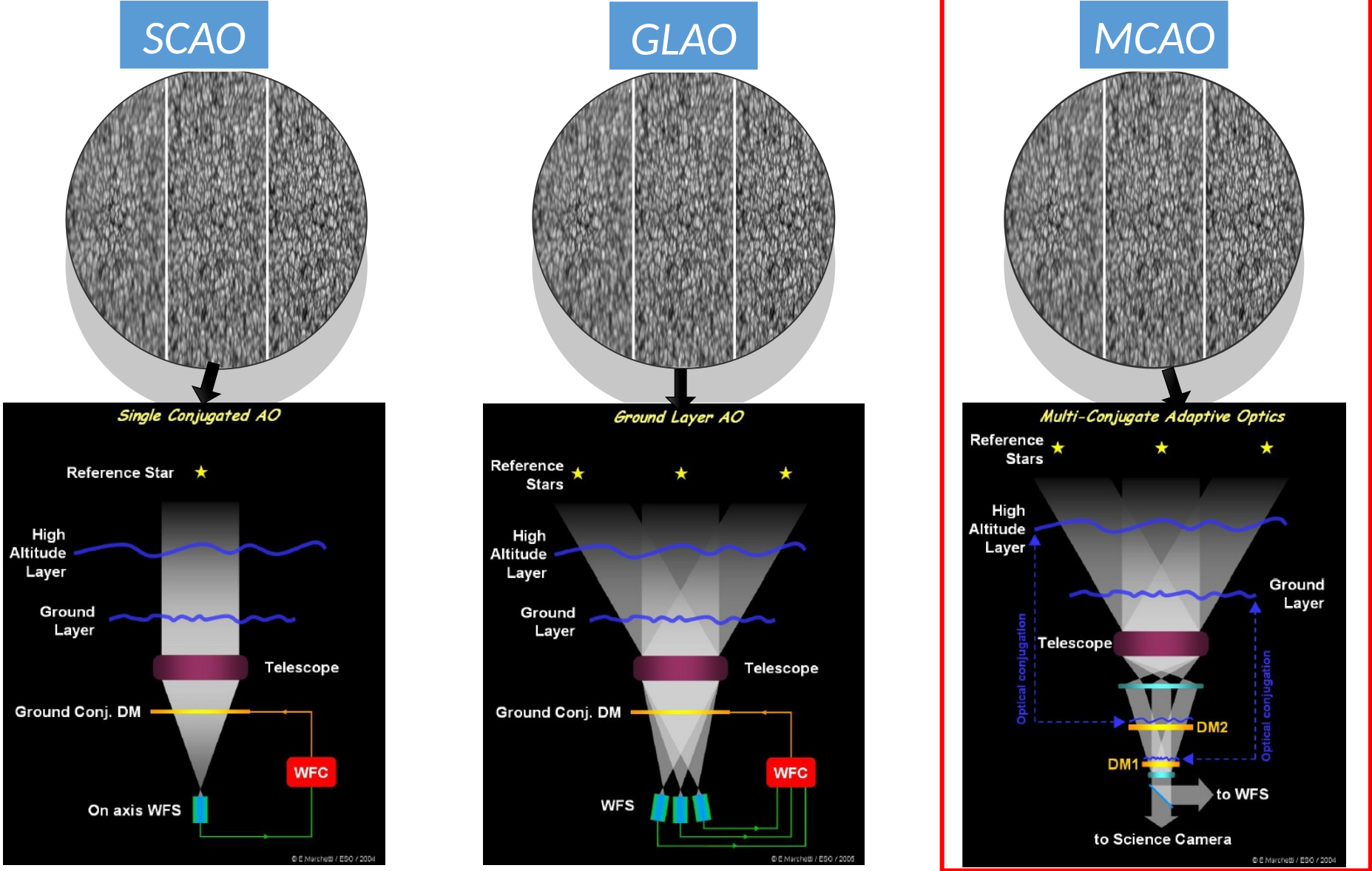
Gobierno de Canarias

Consejería de Economía, Conocimiento y Empleo  
Agencia Canaria de Investigación, Innovación y Sociedad de la Información



- 4-meter Solar Telescope
- On-axis Gregorian configuration
- Spectral range: 400 - 2500 nm
- Strong daytime turbulence
- High resolution in 60'' FoV
- AO/MCAO integrated in the optical path

# Adaptive Optics Configurations



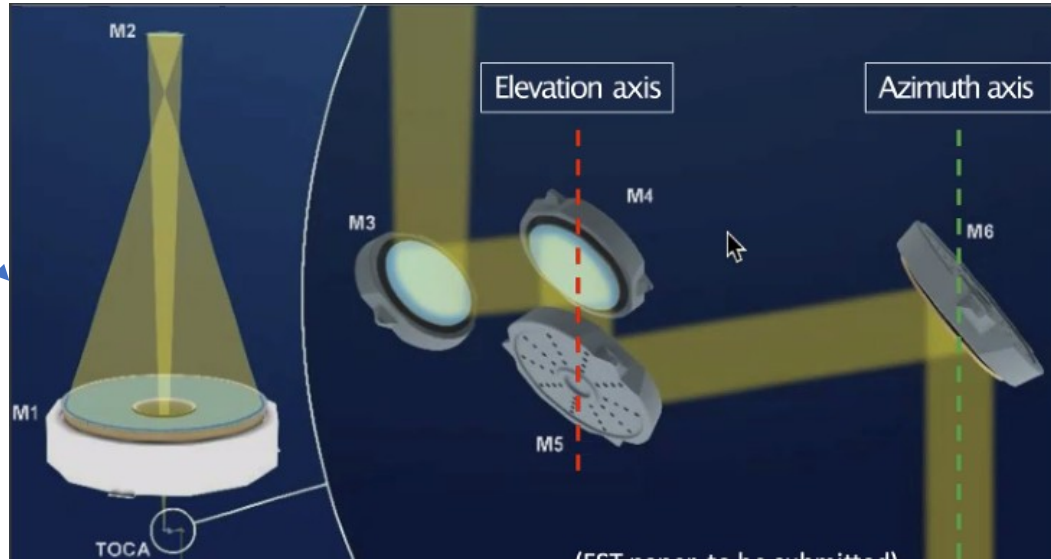
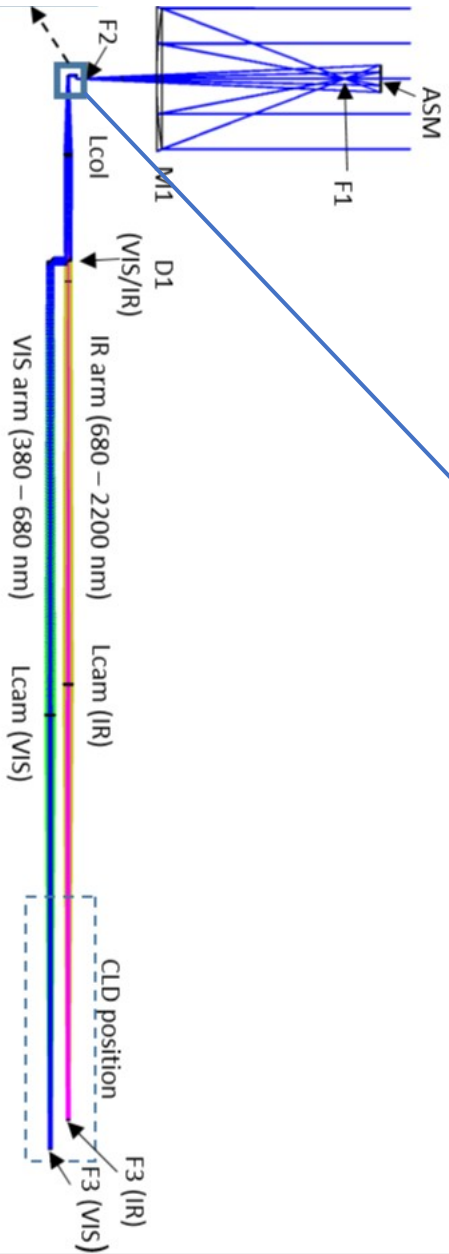
\*The CLEAR multi-conjugate adaptive optics (MCAO) on the Goode Solar Telescope compared to observations with classical AO correction. / Schmidt et al. (2017)



# EST Adaptive Optics

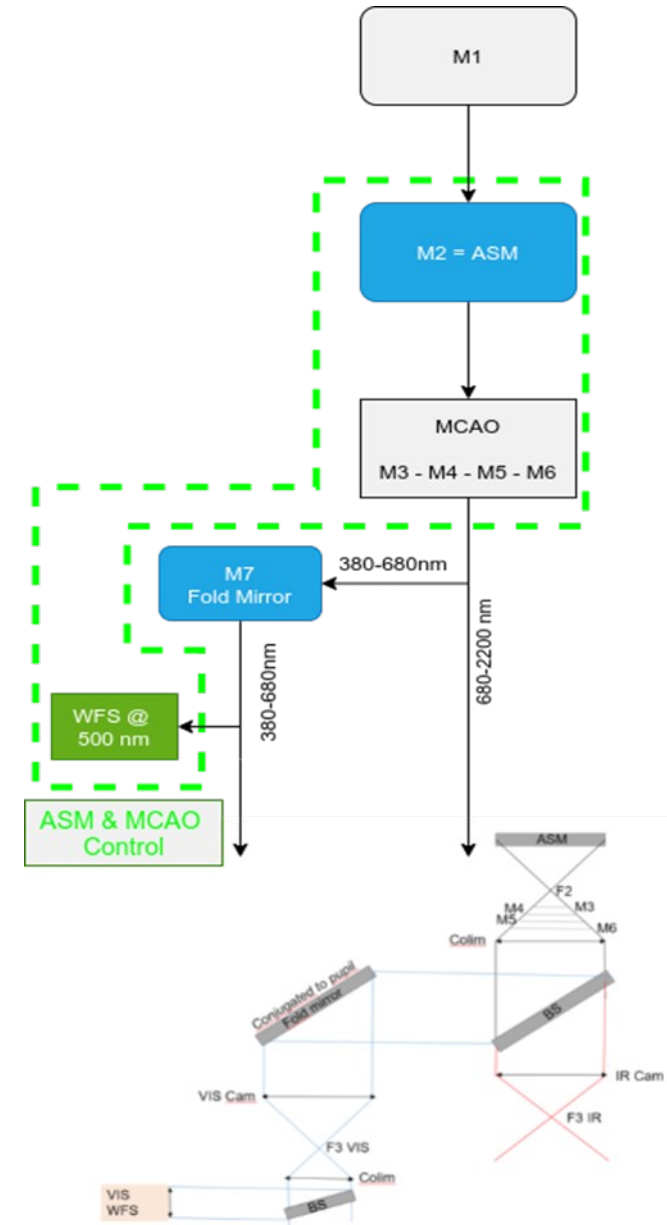


- Adaptive Secondary Mirror (ASM): AO + Tip/Tilt
- 4 Altitude DMs conjugated between 5 km and 20km
- Visible SH correlation WFS



The SR in the corrected FoV when using MCAO shall be:

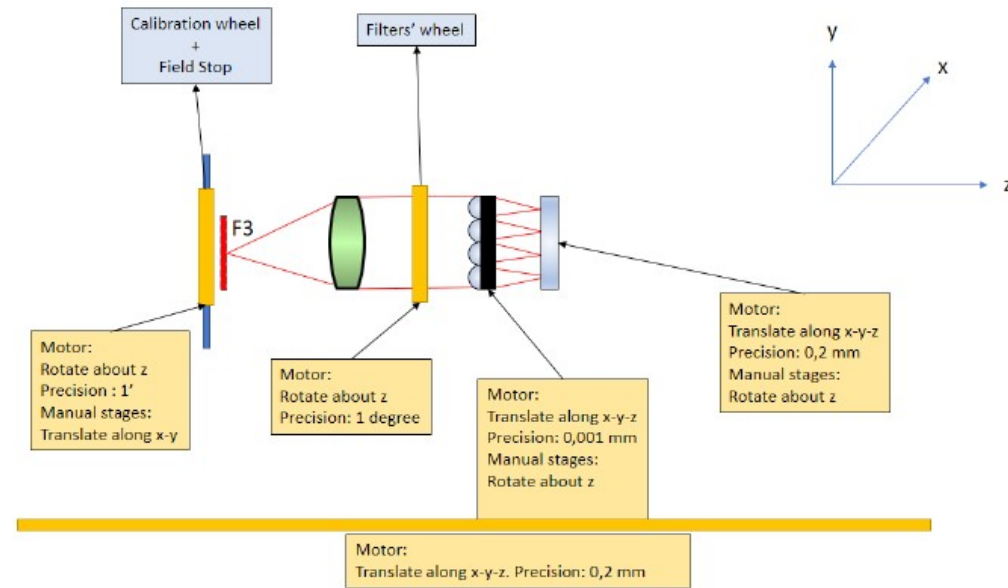
- $r_0 > 10\text{cm}$  &  $\text{FoV} \leq 30$  -->  $\text{SR} \geq 0.25$
- $r_0 > 10\text{cm}$  &  $30 \leq \text{FoV} \leq 60$  -->  $\text{SR} \geq 0.15$
- $r_0 > 20\text{cm}$  &  $\text{FoV} \leq 30$  -->  $\text{SR} \geq 0.4$
- $r_0 > 20\text{cm}$  &  $30 \leq \text{FoV} \leq 60$  -->  $\text{SR} \geq 0.3$



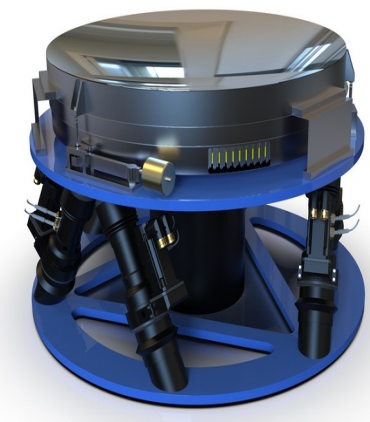
# Ongoing AO Activities

- Error Budget of AO combining analytical and CFD modelling
- Definition of WFSs: configurations, number of microlenses...
- ASM: two different prototypes
- N to N simulations: DASP, FriMM, YAO, PROPER
- Laboratory test bench

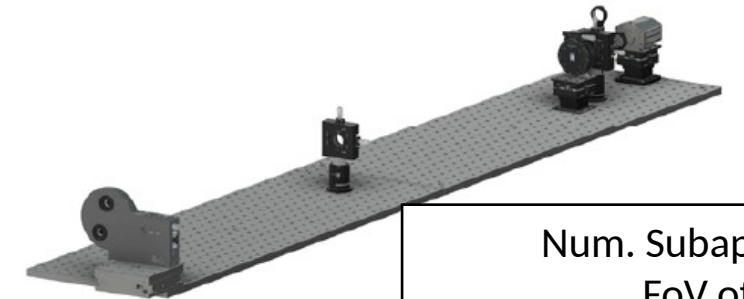
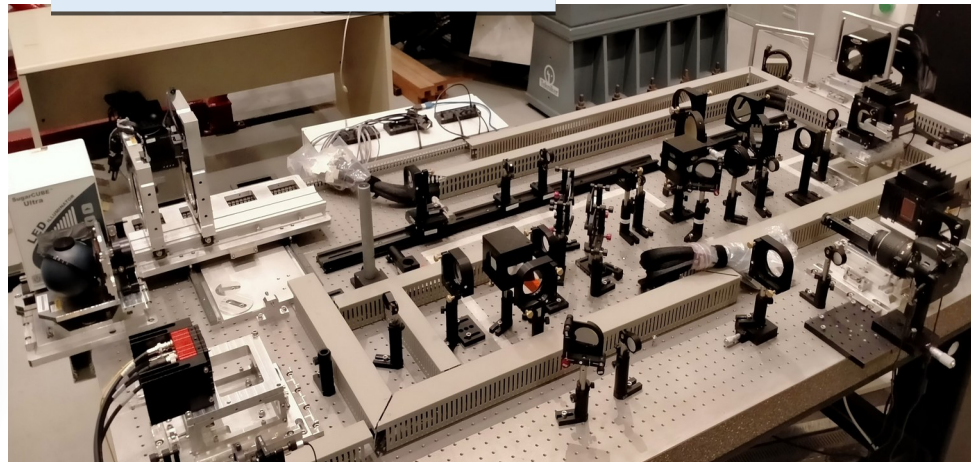
## Narrow Field HO-WFS



M2 - ASM



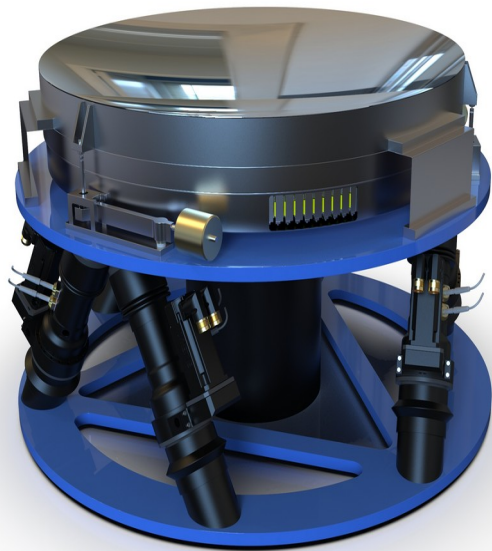
MCAO test bench



Num. Subap > 50  
FoV of 10''  
Frame rate  $\geq$  2000 fps

# Adaptive Secondary Mirror

- The ASM is a challenging technology.
- Two different prototypes are under development: TNO + AdOptica
- Aperture Stop
- Fast Tip/Tilt: Assembled on a hexapod mount with 5 degrees of freedom (piston,  $\delta x$ ,  $\delta y$  and slow tip-tilt)
- High order correction



TNO prototype

Aperture	800 mm
Shape	Concave ellipsoid
Num. Actuators	2000
Act. pitch	$\leq 16.2$ mm

# Objectives of AO tests



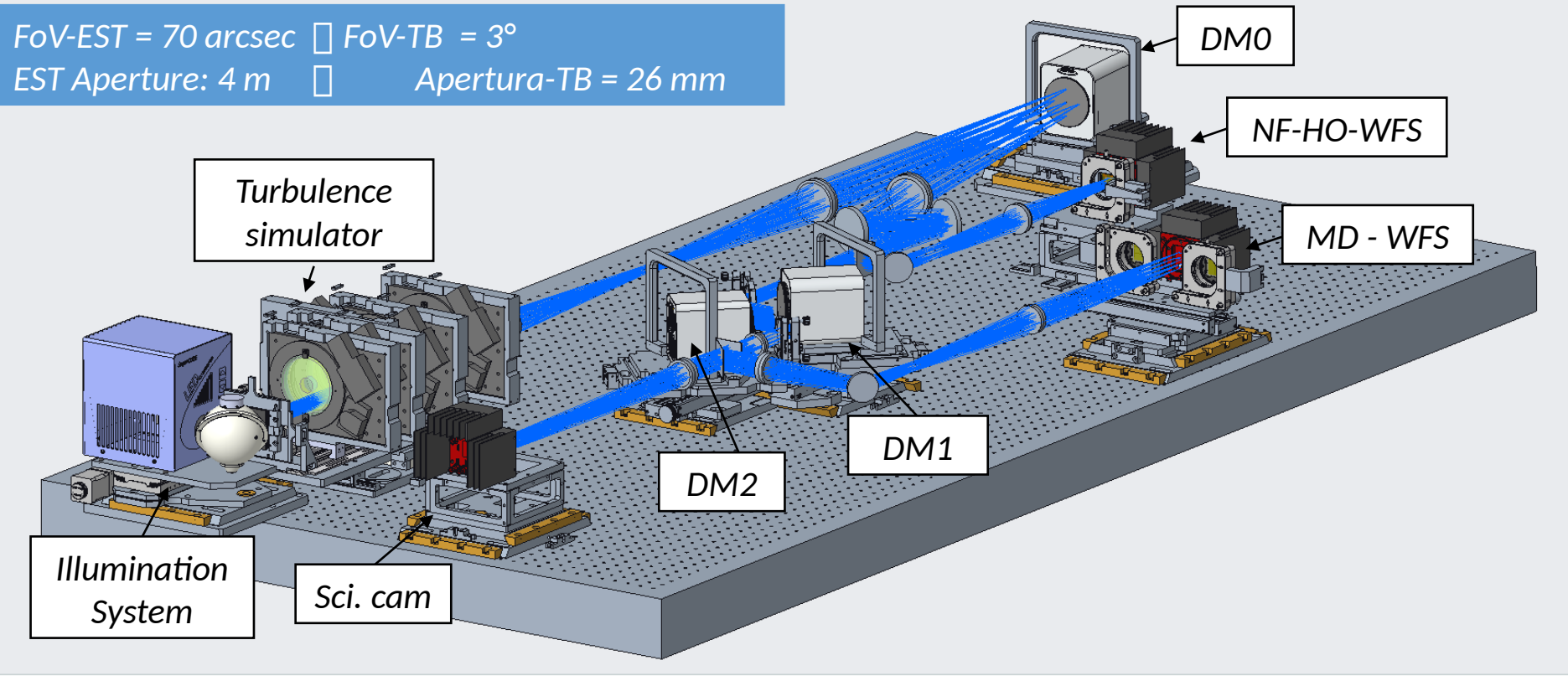
- Find **optimal configuration** for solar MCAO.
- Impact of the **order of the DMs** in the MCAO performance.
- Impact of the **conjugation of DMs at certain heights**.
- Impact of using **45° oriented DMs** on the MCAO
- Mitigate **pupil misregistration effects** due to pupil distortion caused by altitude DMs.
- Mitigate **pupil rotation and spider footprint effects** on wavefront sensing and reconstruction.
- Implement **Artificial Neural Network (ANN)** in the reconstruction and control of the MCAO.



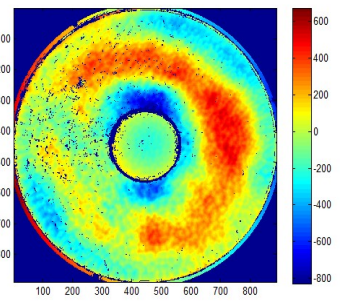
# MCAO Test Bench (TB)



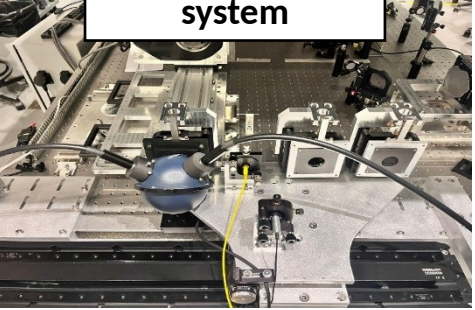
$FoV-EST = 70 \text{ arcsec}$     $FoV-TB = 3^\circ$   
 $EST \text{ Aperture: } 4 \text{ m}$     $Apertura-TB = 26 \text{ mm}$



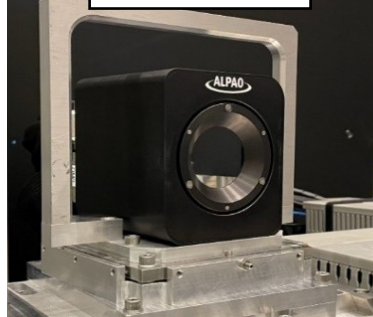
Phase Screens



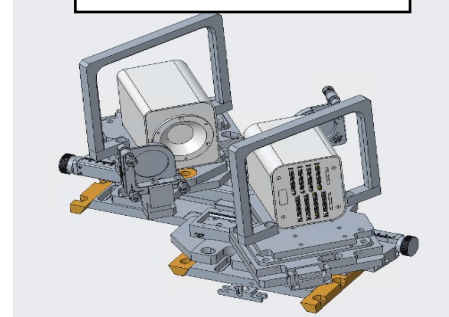
Illumination system



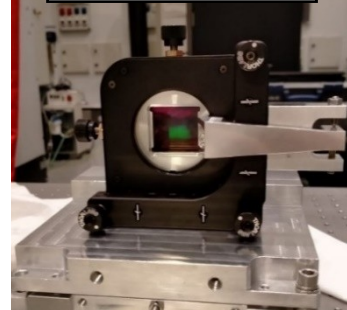
DM0@820



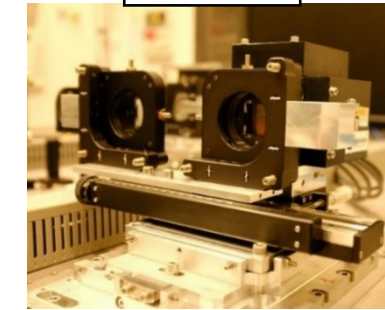
DM1 & DM2 @468



NF-HO-WFS

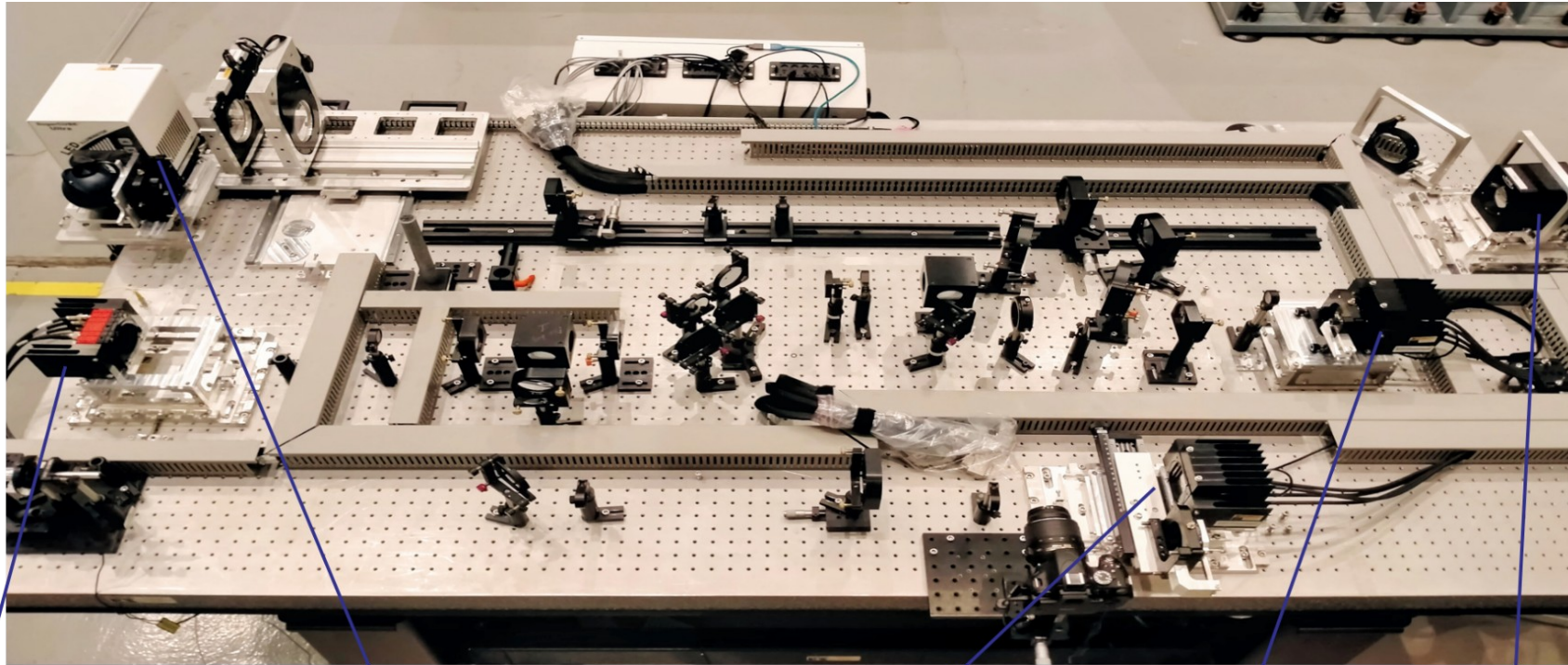


MD-WFS





# MCAO Test Bench (TB)



Sci - cam

Illumination System and  
Turbulence simulator

MD - WFS

HO - WFS

DM

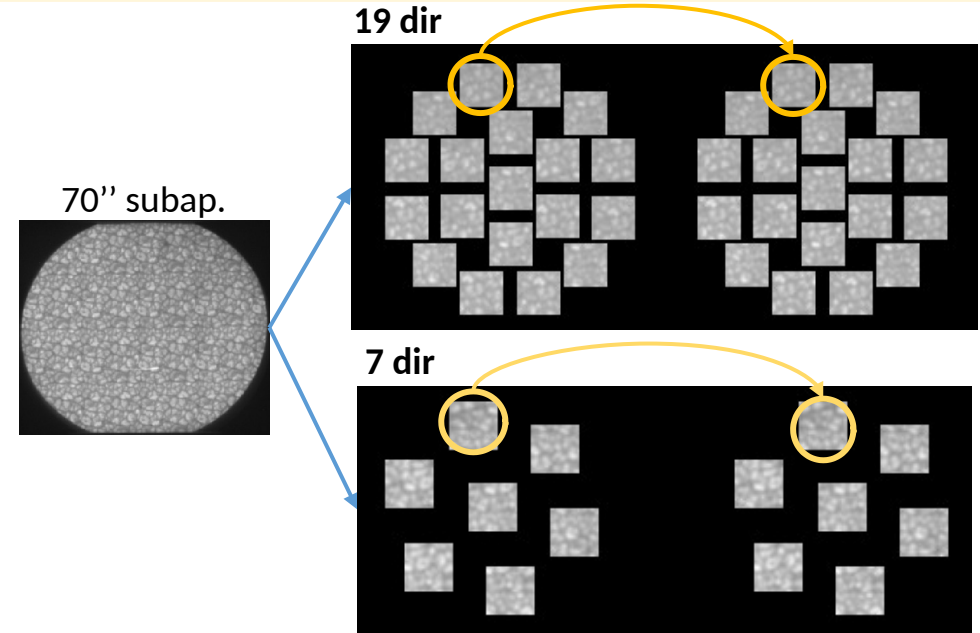
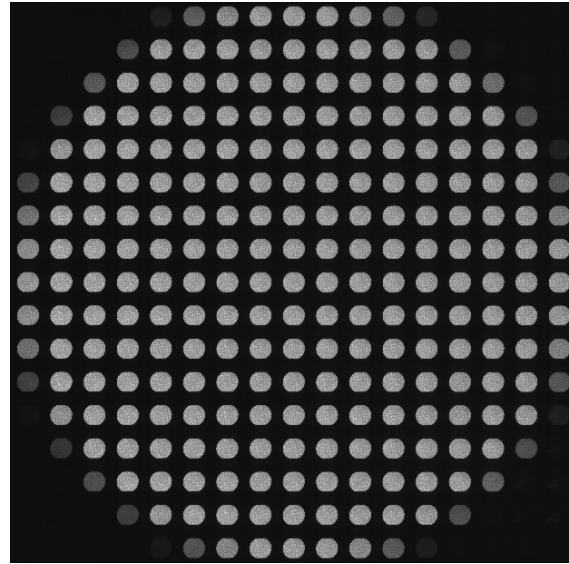
## Configurations

SCAO 1	GLAO1	GLAO 2	MCAO 1	MCAO 2	MCAO 3	MCAO 4
DM0 / HO-WFS	DM0 / HO-WFS / MD- WFS	DM0 / HO- MD- WFS	DM0 / HO-WFS / DM1/ MD-WFS	DM0 / HO-WFS / DM2 / DM1 / MD-WFS	DM0 / DM1/ HO- MD-WFS	DM0 / DM2 / DM1/ HO-MD- WFS

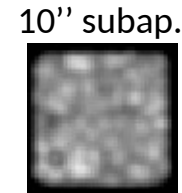
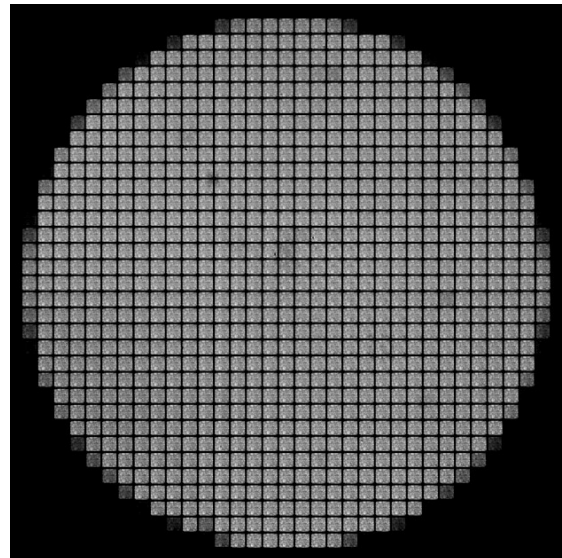
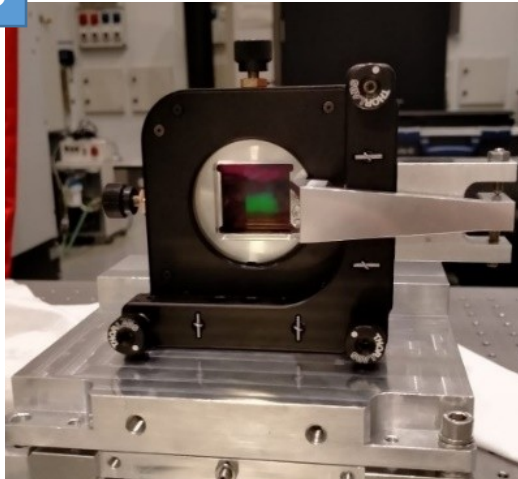
# TB: Wavefront Sensors (WFSs)



MD-WFS



HO-WFS

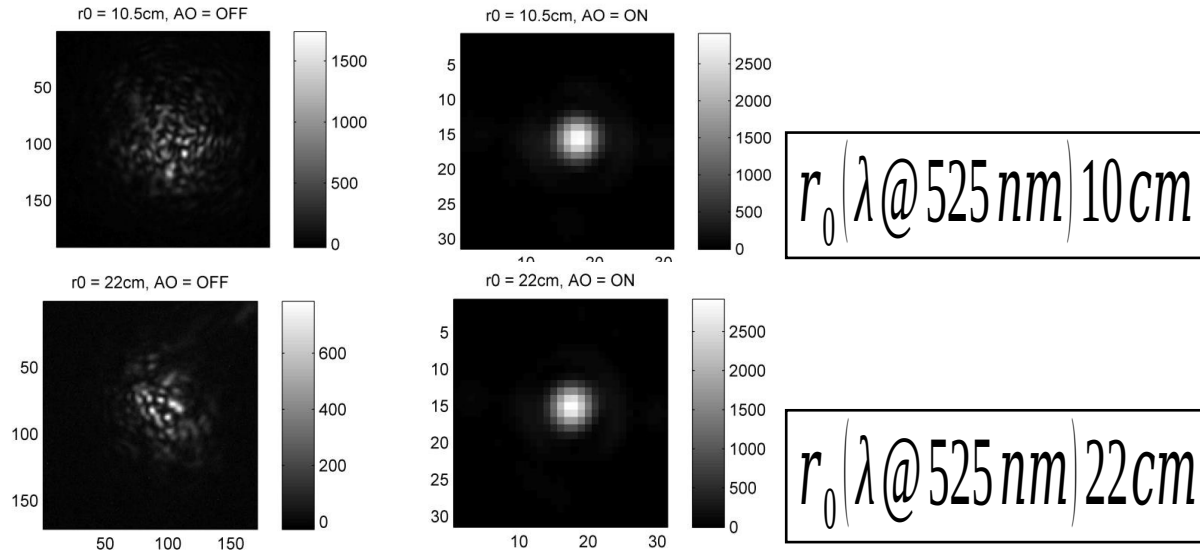


	HO-WFS	HO-MD-WFS	LO-MD-WFS
ML array (MLA)	33 x 33	33 x 33	17 x 17
Pix/subap	40 pix (binning mode 2x2)	175 pix	233 pix
FoV	10"/subap	70"/subap	70"/subap
Plate Scale	0.25 arcsec/pix	0.4 arcsec/pix	0.3 arcsec/pix

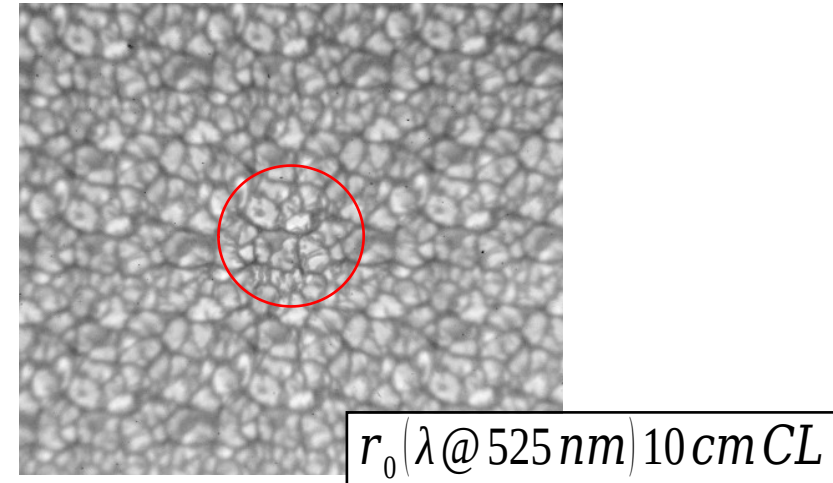
Control RTC: DARC. (Durham University)  
Frequency: 30 Hz (WFS limited)



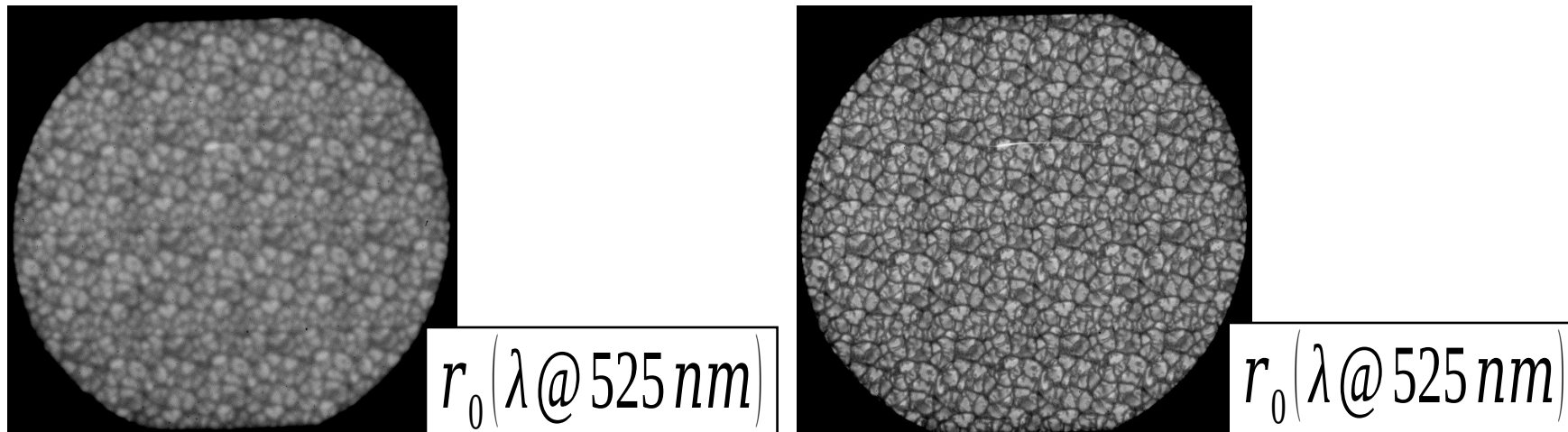
## SCAO point source



## SCAO extended object



## GLAO extended object







- Improve bench alignment to increase image quality



- Improve correlation algorithms.



- Complete GLAO configurations



- Implement MCAO with altitude DMs.



- Verify simulation results



- Future AO projects

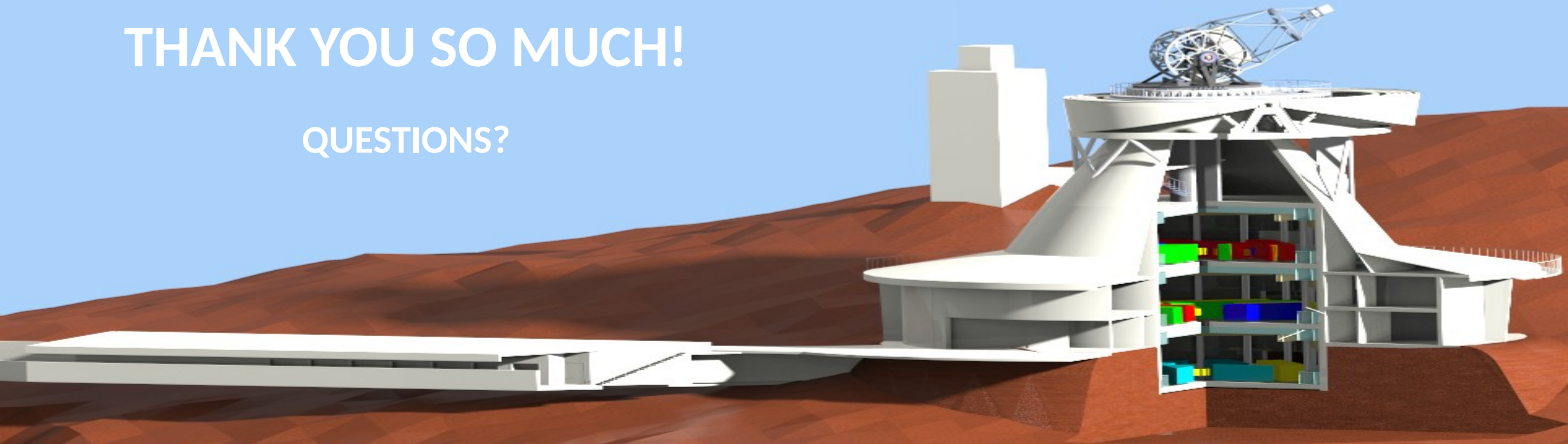


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# THANK YOU SO MUCH!

## QUESTIONS?



Noelia Feijóo, Luzma Montoya, Nicolás Rodríguez, Yolanda Martín, Sergio Bonaque, Jose Manuel González, Haresh Chulani, Francisco González, Jorge Quintero, Miguel Núñez, Ángela Hernández, EST-PO team.

Instituto de astrofísica de Canarias



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