



Leibniz-Institut für  
Astrophysik Potsdam

## *Multiwavelength study of penumbral decay using GREGOR, VTT, DST, NST, and Hinode.*

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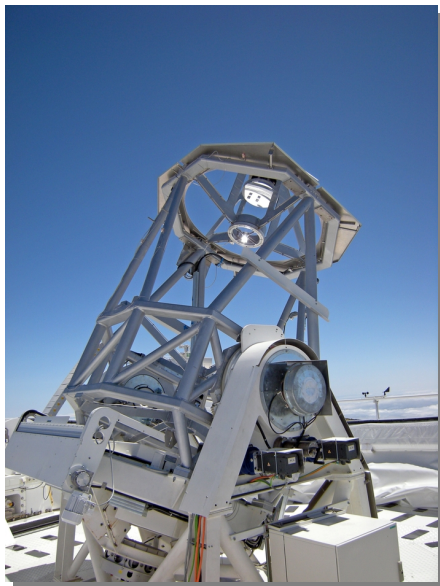
*5 National Solar Observatory (NSO), Boulder, USA*

*6 Universität Potsdam, Institut für Physik und Astronomie, Potsdam, Germany*

# Motivation

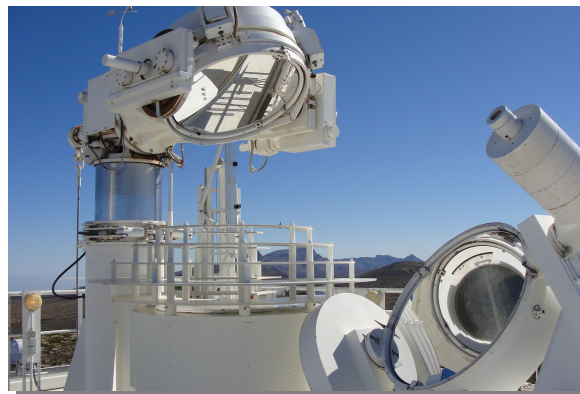
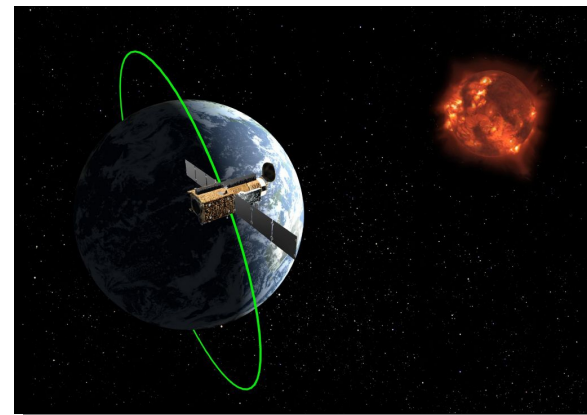
- To exploit the maximum potential of available ground-based as well as space-borne telescopes
- Multi-instrument and multi-wavelength observations
- Various instruments
  - Visible and EUV Imager
  - Fabry-Pérot Interferometer
  - Spectrograph
- Information on the propagation of changes from photosphere to chromosphere and even to transition region

GREGOR



# Motivation

Hinode



VTT

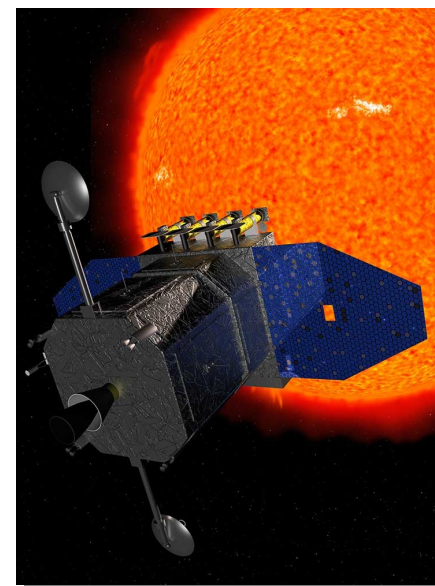


NST



DST

SDO



# Proposed Observations

Active Region Filaments: Observing Shear Flows and the Evolution of Magnetic Shear along Magnetic Neutral Lines with GREGOR, VTT, DST, NST, and Hinode

- GREGOR
  - HiFI: G-band ( $\lambda 430.7\text{nm}$ ), blue continuum ( $\lambda 450.6\text{ nm}$ )
  - GFPI: spectroscopic data - Fe I ( $\lambda 617.34\text{ nm}$ ) and Fe I ( $\lambda 543.4\text{ nm}$ )
  - GRIS: spectropolarimetric data - Si I ( $\lambda 1082.7\text{nm}$ ) He I ( $\lambda 1083.0\text{ nm}$ )
- VTT – Echelle spectrograph - spectral data H $\alpha$  and Na D<sub>2</sub>
- DST
  - IBIS: Ca II ( $\lambda 854.2\text{ nm}$ ), Na( $\lambda 589.0\text{ nm}$ ), and spectroscopic H $\alpha$
  - ROSA: Images in G-band, Ca IIK, and Hbeta
  - FIRS: Spectropolarimetric data in 1083.0 nm spectral range
- NST
  - BFI: TiO
  - NIRIS: spectroscopic data He I triplet ( $\lambda 1083.0\text{ nm}$ ) range
  - FISS: H $\alpha$  and Ca II H ( $\lambda 854.0\text{ nm}$ )
- Hinode – SP: Fast scans Fe I ( $\lambda 630.15\text{nm}$ ) & ( $\lambda 630.25\text{nm}$ )

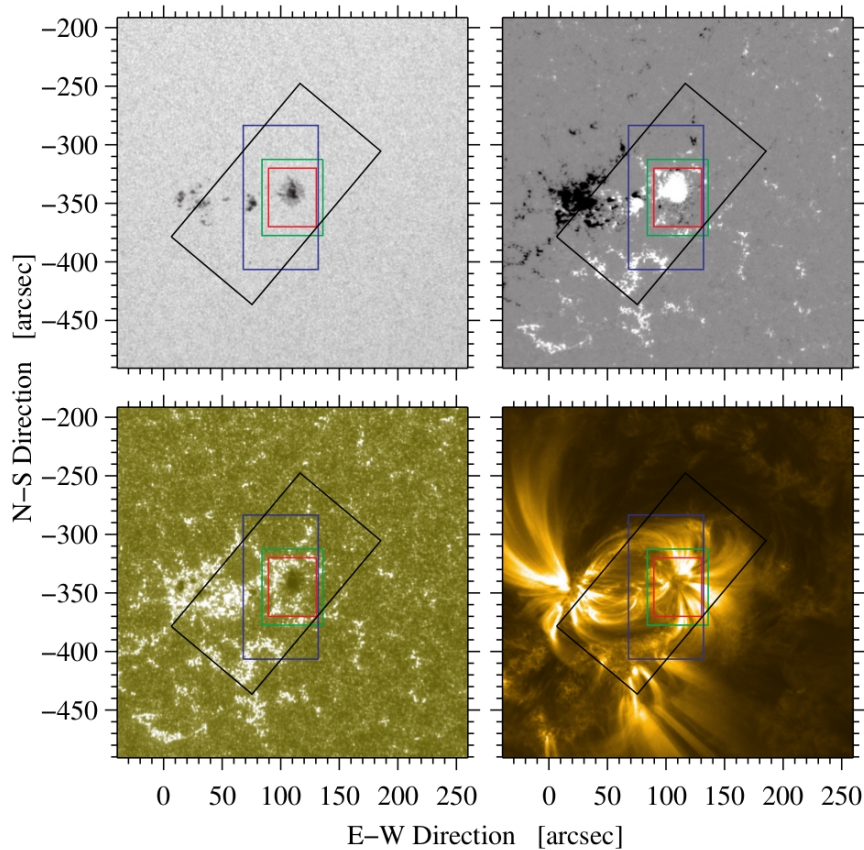
# Observations

DATE	TELESCOPES	REGION
16/09/2016 – 19/09/2016	VTT, Hinode	AR12592, AR12593
20/09/2016	VTT, DST*, Hinode	AR12594
21/09/2016 – 23/09/2016	VTT, GREGOR, DST*, NST** Hinode***	AR12593
24/09/2016	VTT, GREGOR, NST, Hinode	AR12597
25/09/2016	NST, Hinode	AR12597
26/09/2016	VTT, GREGOR, Hinode	AR12597
27/09/2016	VTT, Hinode	AR12597
28/09/2016	VTT, GREGOR, DST, Hinode	AR12597
29/09/2016	VTT, GREGOR, Hinode	AR12597, Filament in NE

# Observations

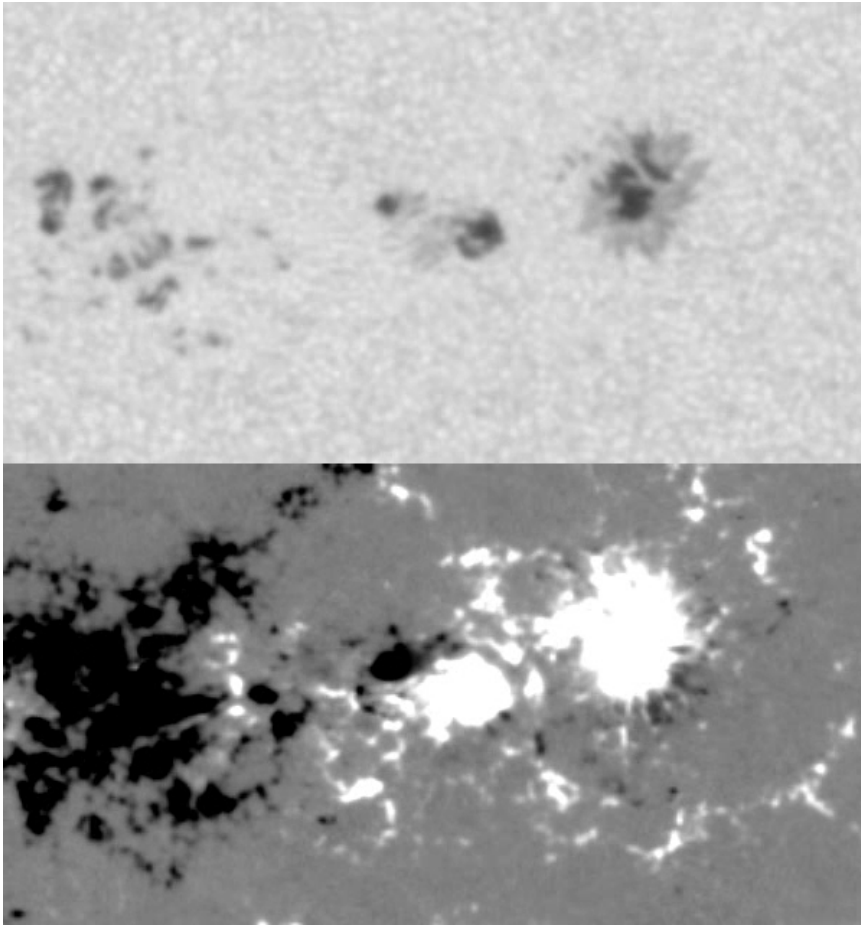
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# Leading Sunspot in NOAA 12597



- Glimpse of data and preliminary results
- 08:52 UT on 2016 September 24
- Appeared on south east near disk center on September 22
- Position on September 24 (110", -350")
- Classified as  $\beta$ - simple bipolar region but developed later to  $\beta\gamma/\beta$  region with complex neutral line
- Focused on the leading spot
- Mature sunspot with decaying penumbra
- SDO – continuum, LOS magnetogram, 1600 nm, and 171 nm
- Boxes are FOV covered by GFPI, GRIS, Hinode, and VTT

# Leading Sunspot in NOAA 12597

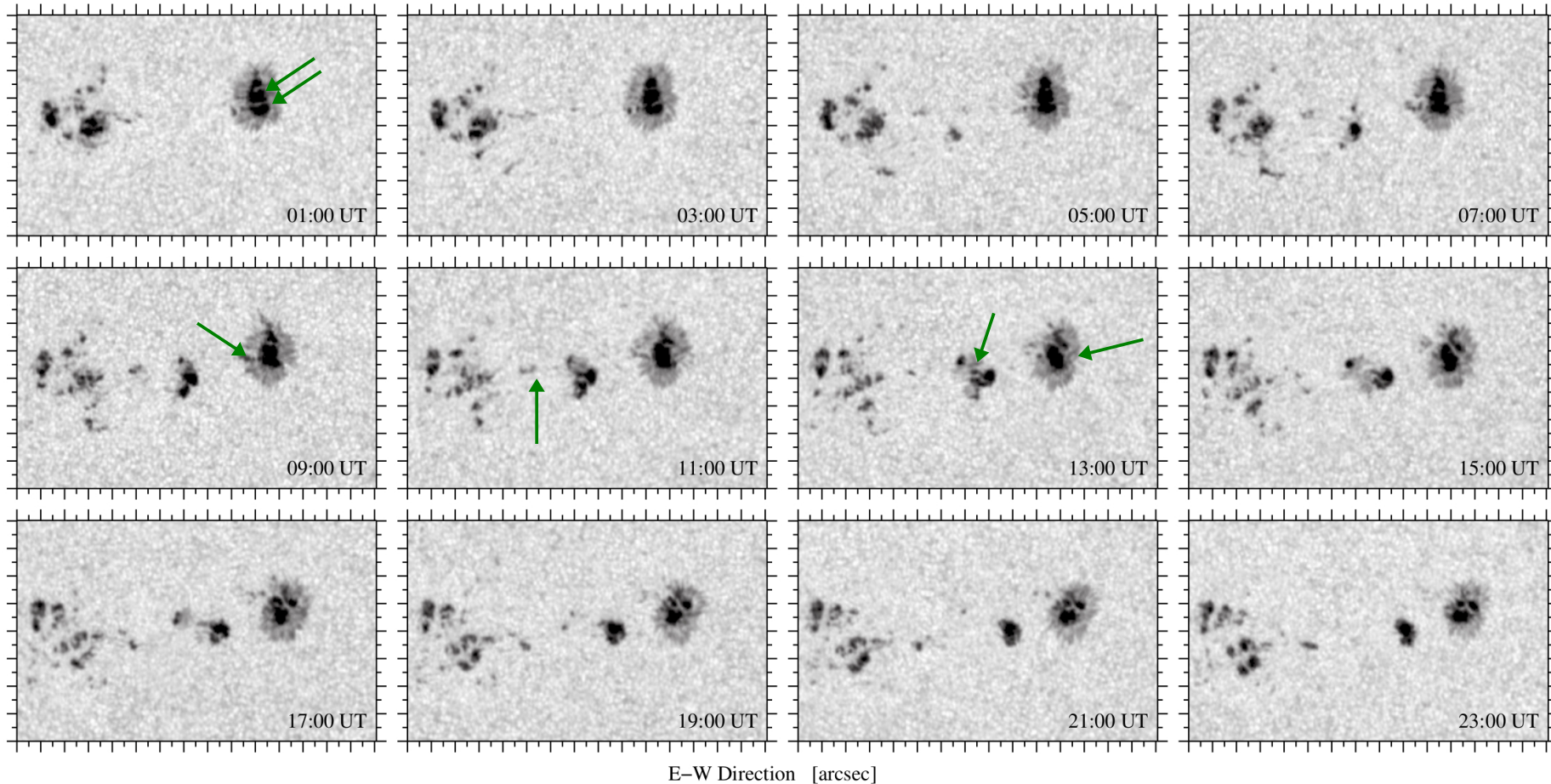


- Appeared on south east near disk center on September 24
- Position on September 24 (110", -350") and classified as  $\beta$ /-
- Focused on the leading spot
- Mature sunspot with decaying penumbra
- Two light-bridges, one disappearing
- Appearance of darkened area resembling umbral core on the edge of disappeared light bridge
- Sunspot rotation
- Flux emergence

SDO continuum and LOS magnetogram movie



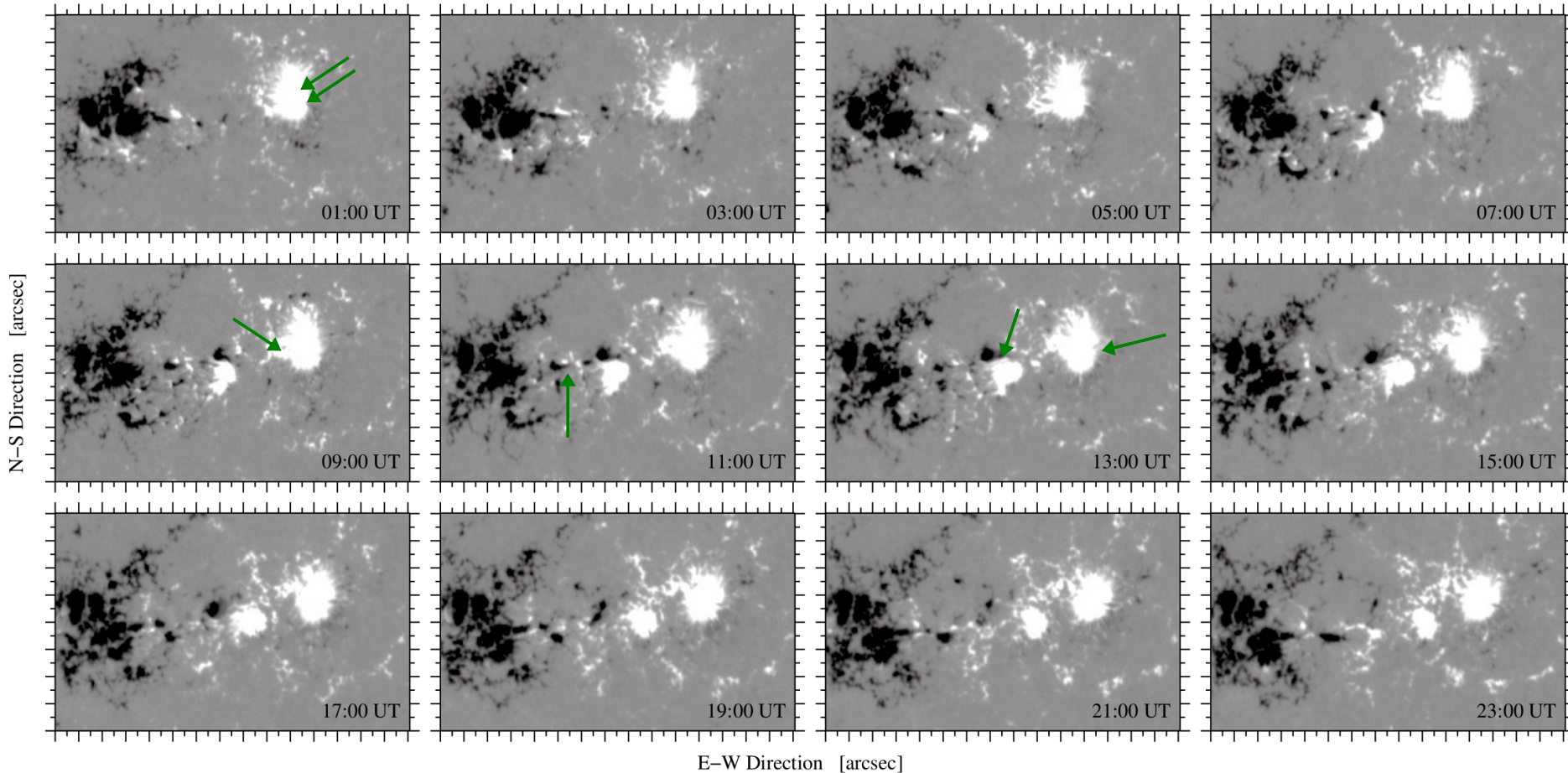
# Evolution on the day of observation



- Two light-bridges, one disappearing
- Appearance of umbral core extrusion on the edge of disappeared light bridge

- Flux emergence
- Sunspot rotation
- Pores of different polarity sliding

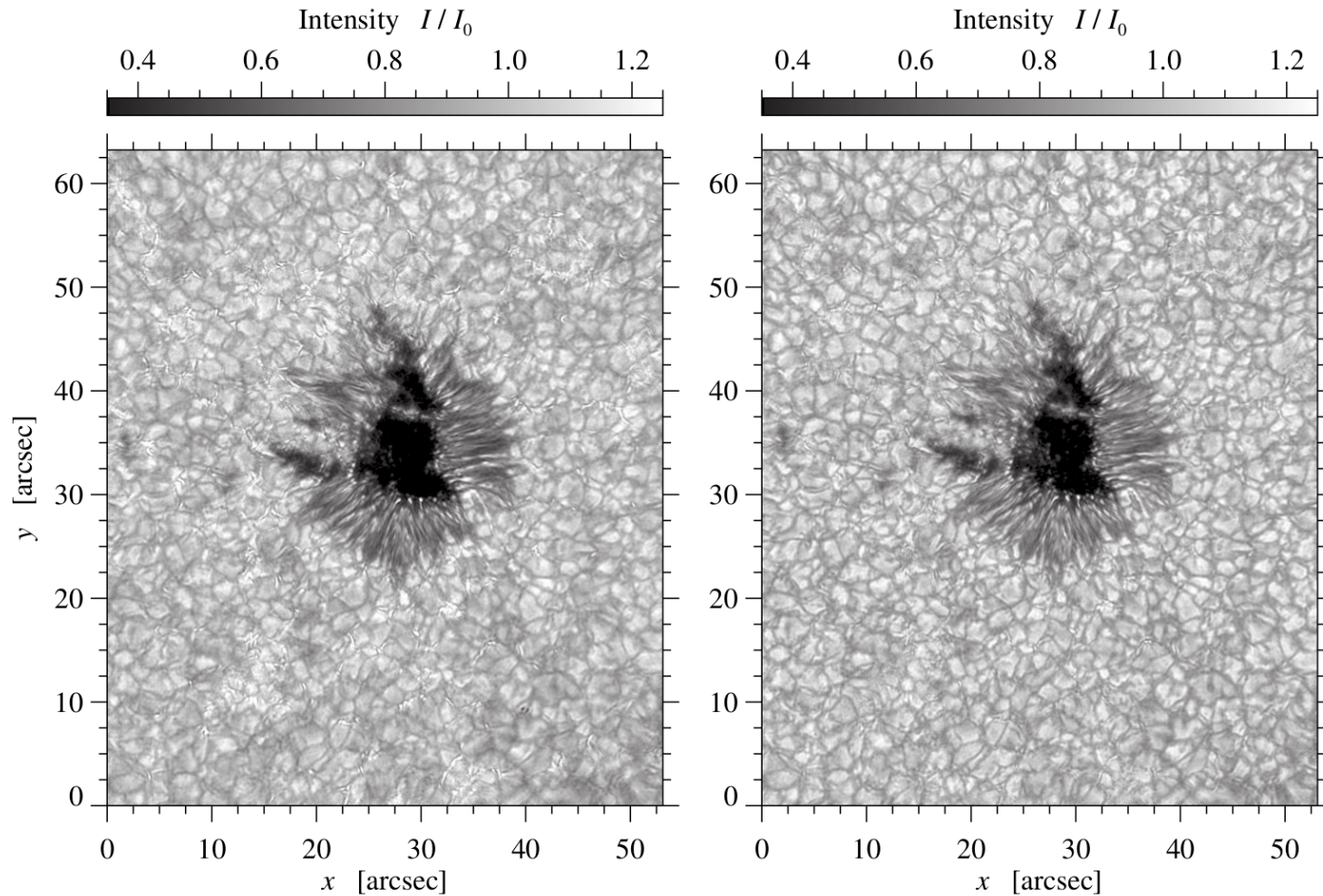
# Evolution on the day of observation



- Two light-bridges, one disappearing
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# GREGOR High-resolution Fast Imager (HiFI)



G-band and Blue  
continuum images

Time-series for about  
40 min

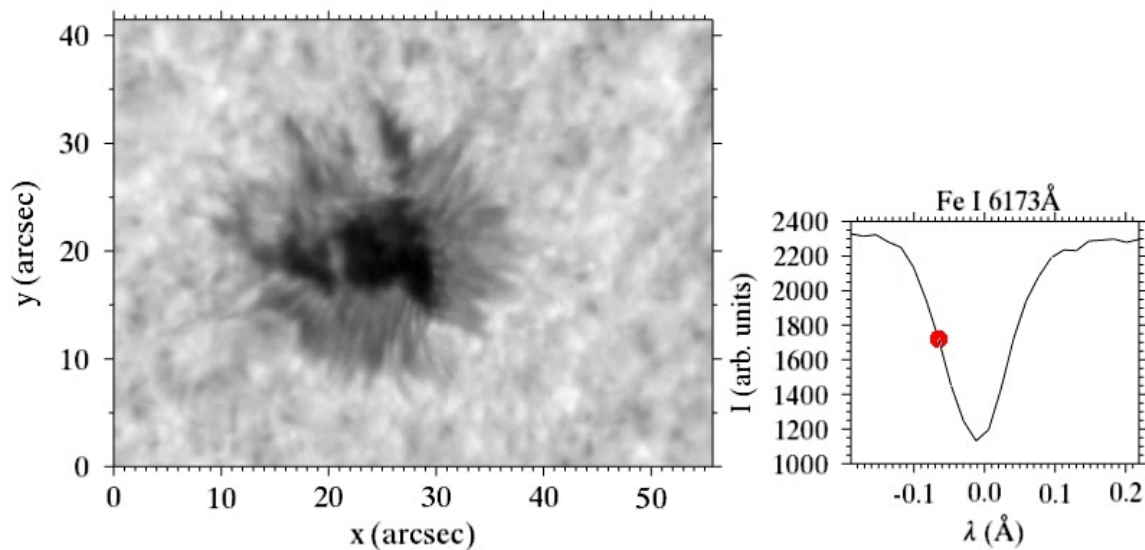
Appearance of  
darkened umbral  
core like region

Granulation in the  
penumbral gap

Evolving light-bridge

Next step – Apply LCT  
to follow horizontal  
proper motions

# GREOR Fabry-Pérot Interferometer (GFPI)

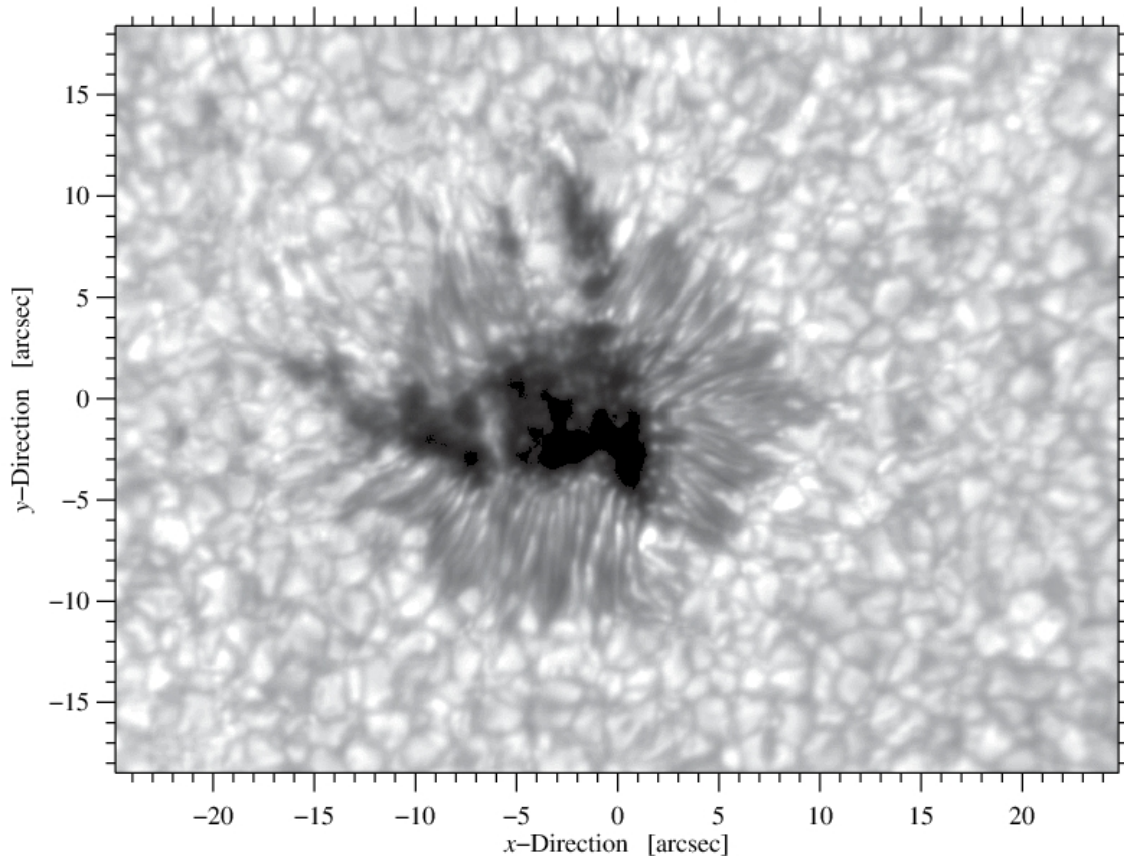


GFPI line scan movie

- Spectroscopic data in 617.3 nm Fe I line
- Exposure time ~ 10 ms
- One line scan ~ 24 s
- More than 140 scans
- Level1 & Level2 MOMFBD data
- Example of line scan

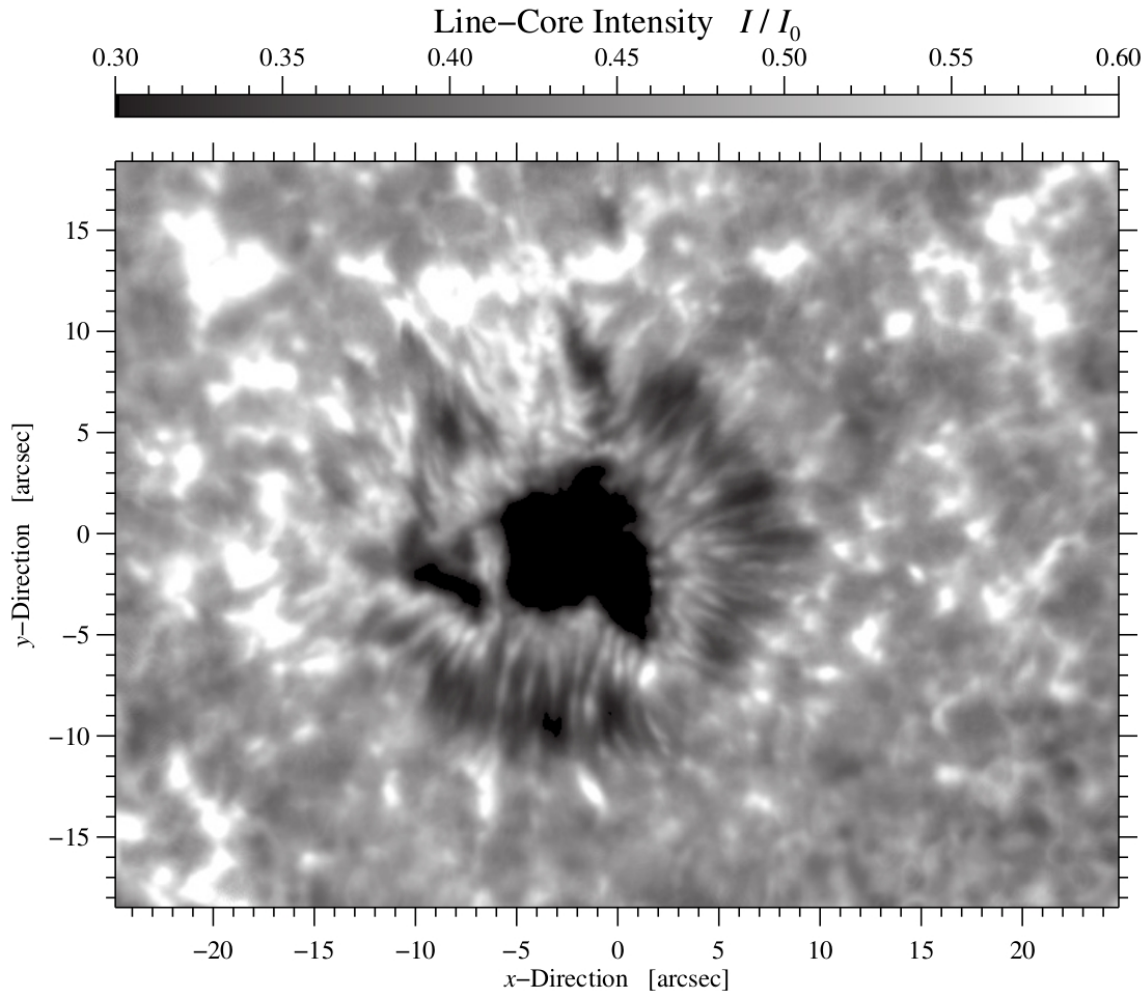
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Broad-Band Image



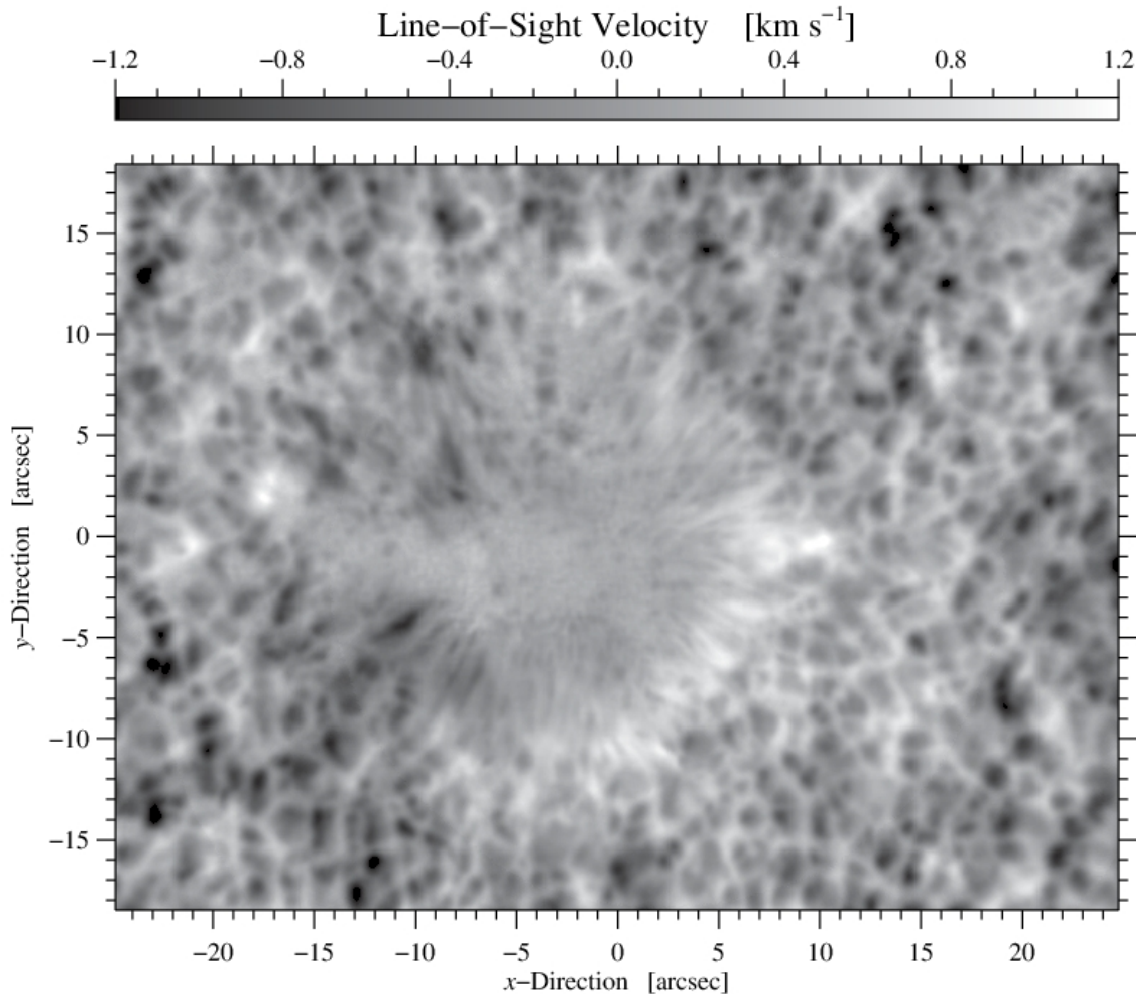
- Spectroscopic data in 617.3 nm Fe I line
- Exposure time ~ 10 ms
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- Example of broad-band image

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- Spectroscopic data in 617.3 nm Fe I line
- Exposure time ~ 10 ms
- One line scan ~ 24 s
- More than 140 scans
- Level1 & Level2 MOMFBD data
- Example of line core intensity map

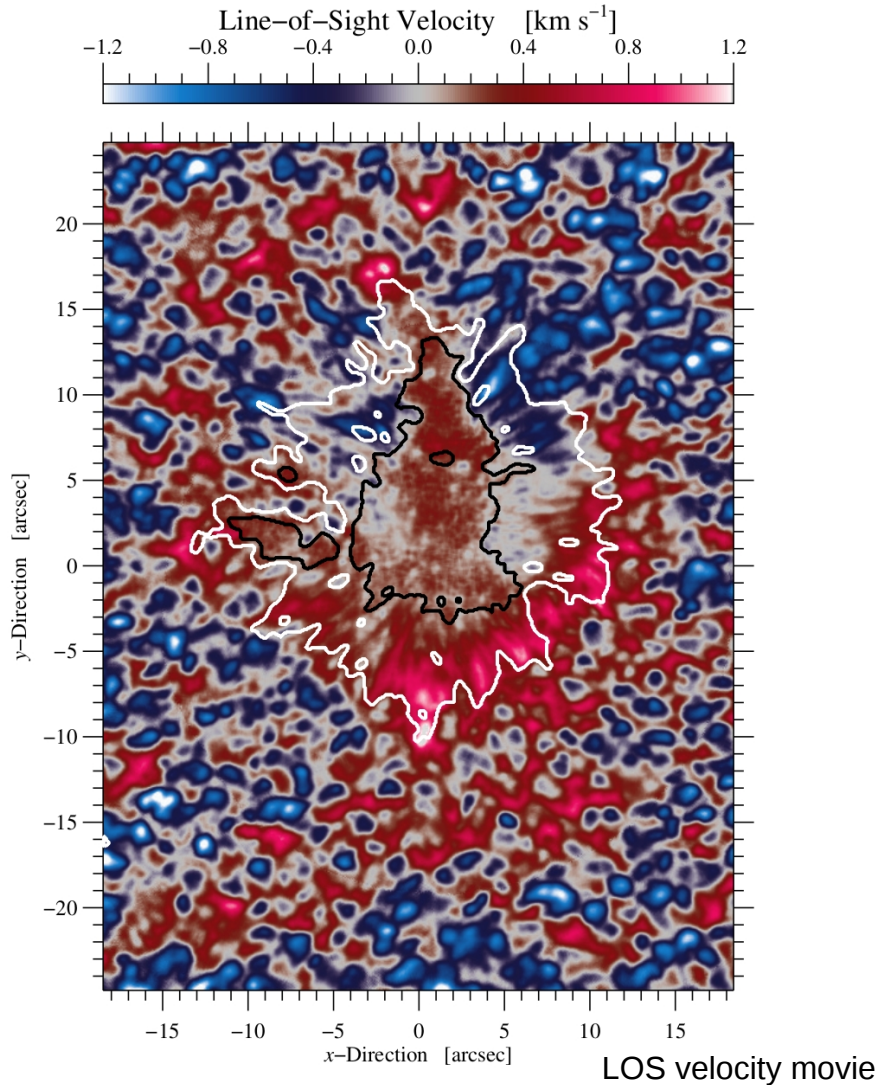
# GREOR Fabry-Pérot Interferometer (GFPI)



- Spectroscopic data in 617.3 nm Fe I line
- Exposure time  $\sim 10$  ms
- One line scan  $\sim 24$  s
- More than 140 scans
- Level1 & Level2 MOMFBD data
- Example of LOS velocity map

See poster on sTools-GFPI data pipeline

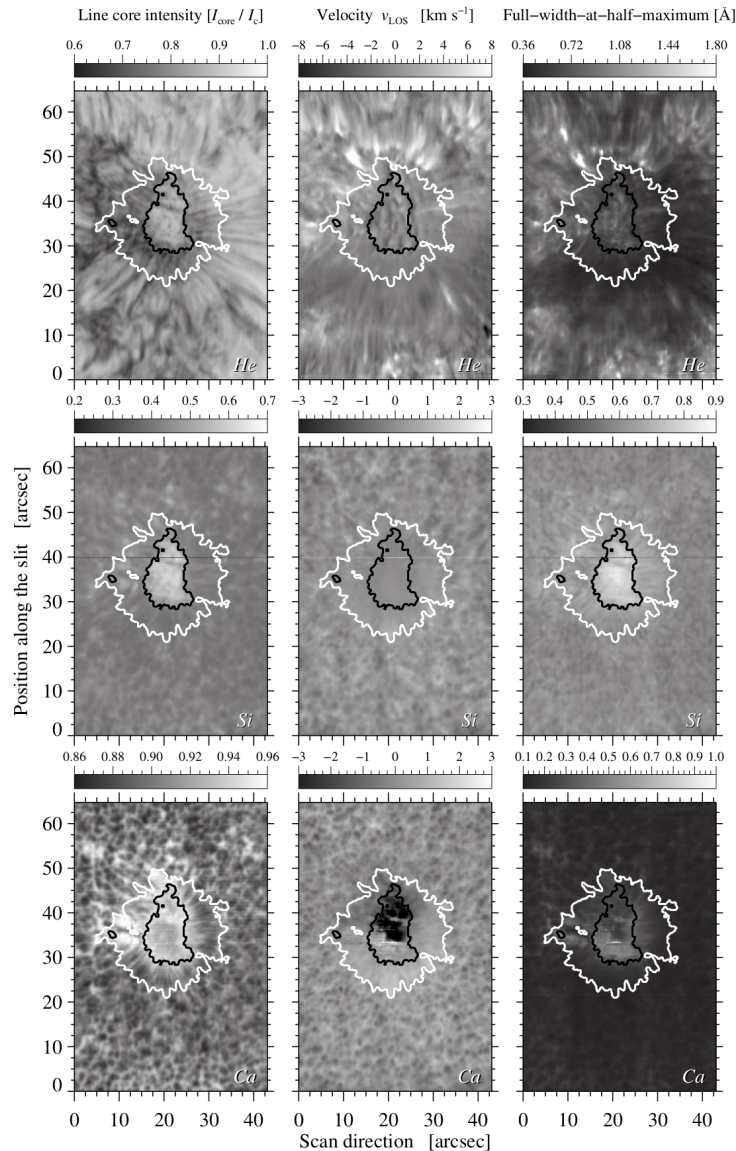
# GREOR Fabry-Pérot Interferometer (GFPI)



- LOS velocities computed using Ipff
- More than 140 maps covering time-period of 40 minutes
- Evershed effect
- Not in the sector of decaying penumbra
- The region next to umbral extrusion has granulation properties
- Indication of umbral flashes

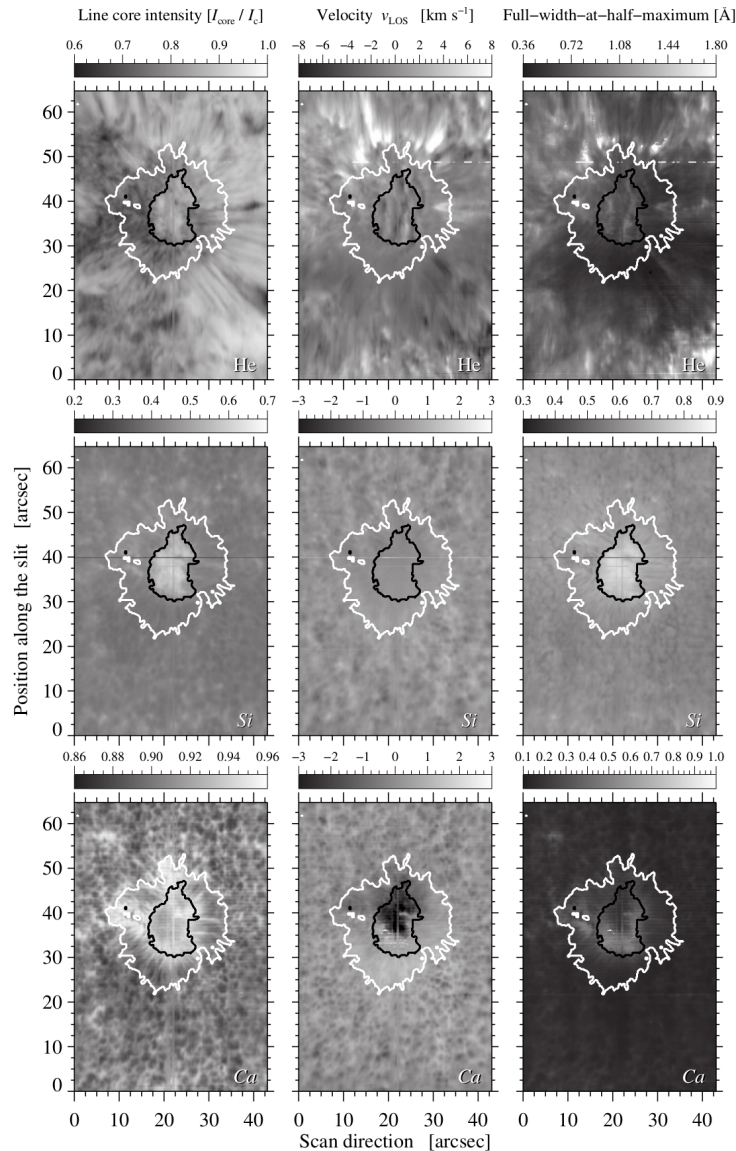


# GREGOR Infrared Spectrograph (GRIS)



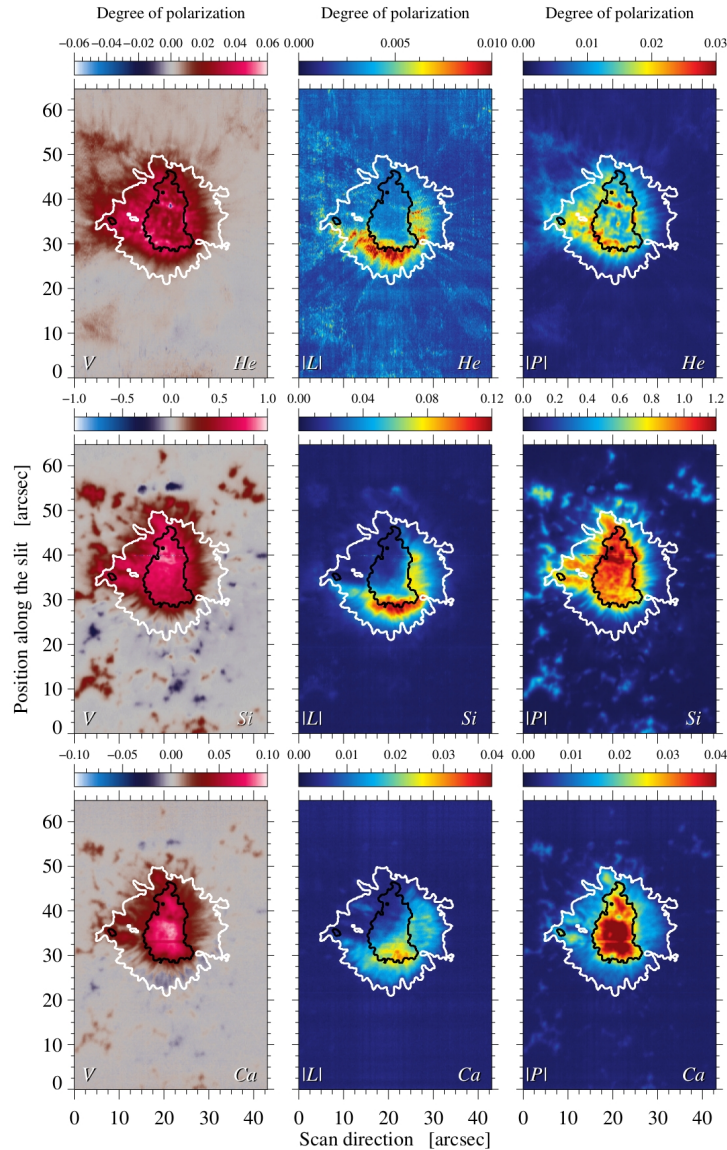
- IQUV Stokes Spectra
- Ca I ( $\lambda 1083.9$  nm, deep photosphere)
- Si I ( $\lambda 1082.7$  nm, photosphere)
- He I ( $\lambda 1083.0$  nm, chromosphere)
- Two scans - 09:02 UT & 10:30 UT
- 360/300 steps
- FOV of  $62'' \times 52'' / 62'' \times 42''$
- Infer magnetic and flow field information from photosphere to chromosphere
- Line core intensity, LOS velocity, & FWHM

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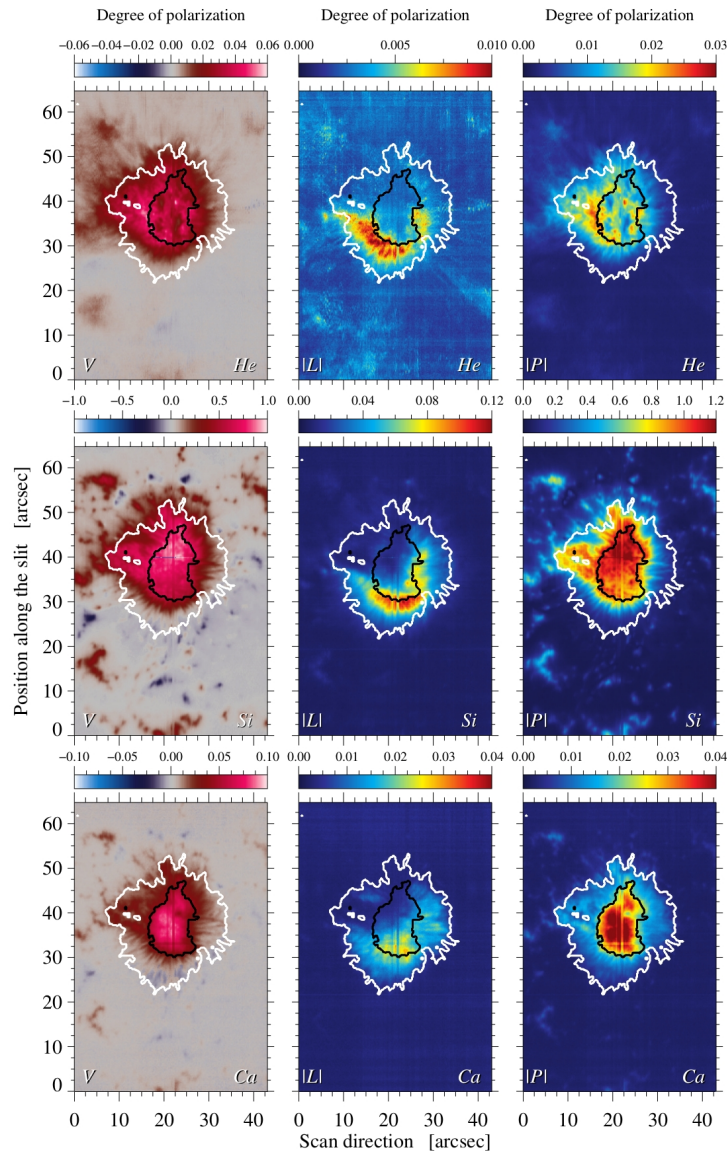
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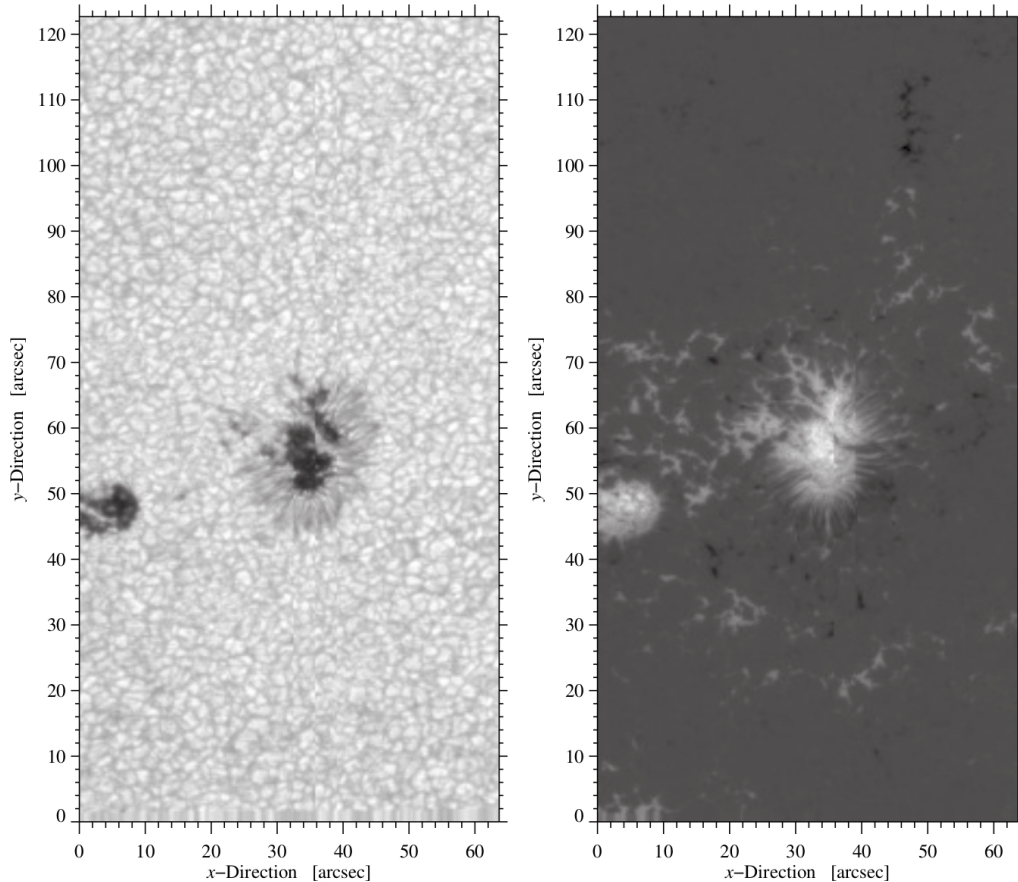
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- Signed V, linear & total degree of polarization
- Small scale magnetic features – unipolar as well as bipolar
- Decaying penumbra sector low linear degree of polarization
- In Si I strong total degree of polarization in the umbral extrusion

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# Summary & Future Work



Hinode continuum and magnetic field

- Glimpse of the potential of coordinated observing campaign
- Preliminary results of multi-wavelength study
- The penumbral sector facing region with flux emergence decays first - forms umbral core extrusion
- Flux emergence triggering the decay of penumbra
- Further steps include
  - Inversion of GRIS data
  - LCT on GFPI and HiFI data
  - Analysis of VTT, Hinode, NST data

*Thank you!*

