Submillimeter Galaxies (SMGs) with JCMT

Fuyan Bian

Stromlo Fellow

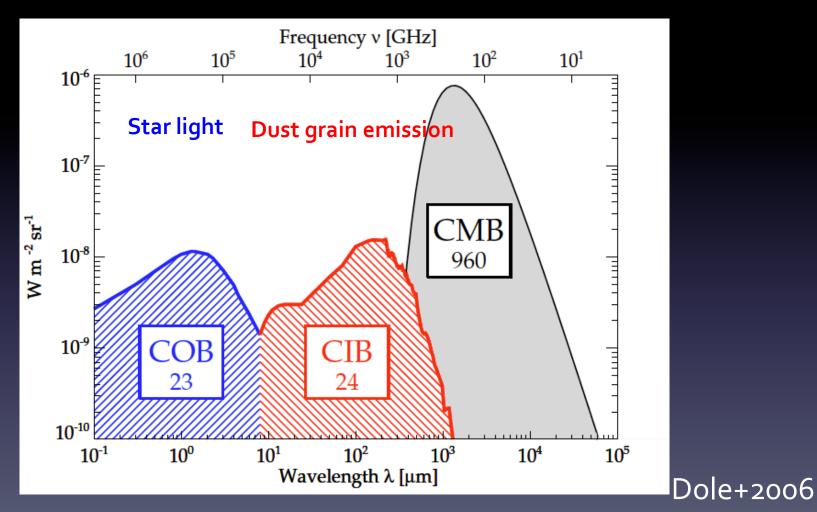
RSAA@ANU

JCMT Workshop UNSW

Outlines

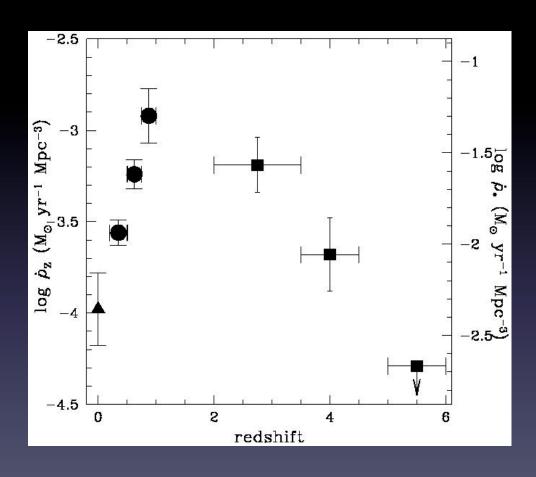
- Introduction
- Discovery of submm galaxies (SMGs)
- Characterize SMGs
- JCMT legacy survey of SMGs (ongoing)
- JCMT large programs (future)

Cosmic Infrared Background (CIB)



CIB is comparable to cosmic optical background

Cosmic Star-formation History



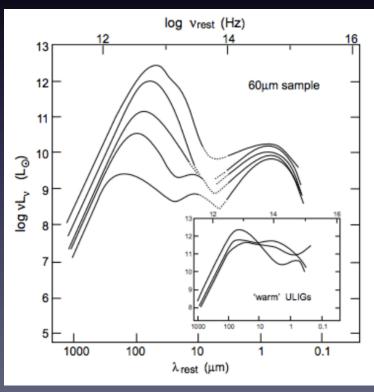
- High-redshift SFR density is based in UVselected galaxies.
- How to correct dust extinction for SFR measurements?
- Missing optical faint but infrared luminous population.

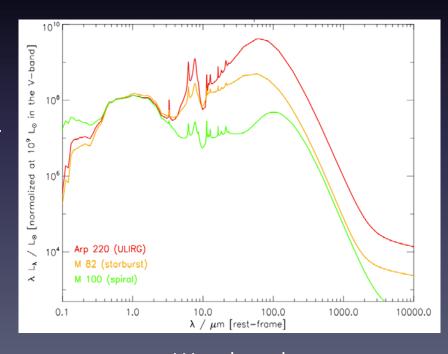
Local Dusty Star-forming Galaxies

• Luminous InfraRed Galaxies (LIRGs) L_{IR} > 10¹¹L_{sun}

-uminosity

• Ultra-Luminous InfraRed Galaxies (ULIGs) L>1012Lsun

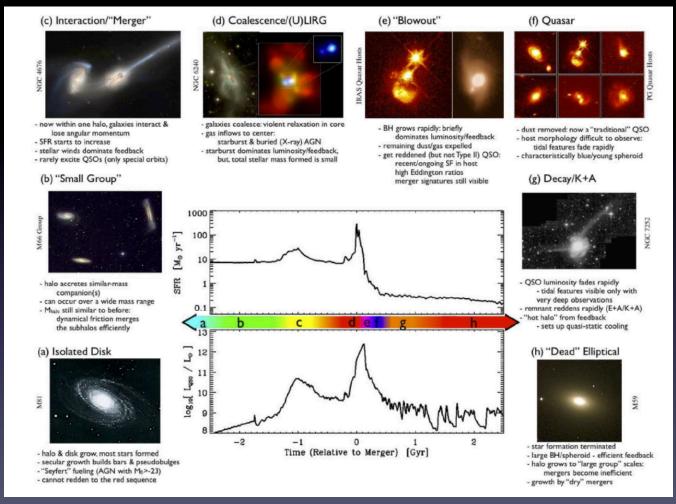




Wavelength

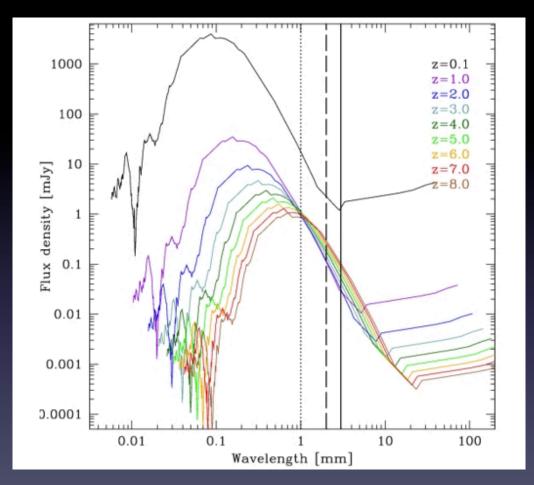
da Cunha+13

Local Dusty Star-forming Galaxies



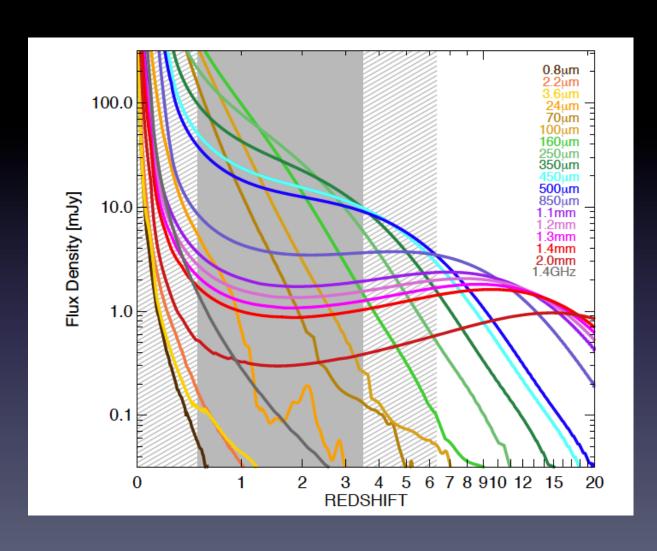
ULIRGs play an important role on massive galaxy formation and supermassive black hole growth

Negative K-correction



- Modified blackbody peaks at 100μm
- Submm wavelength locates in the Rayleigh-Jeans regime.

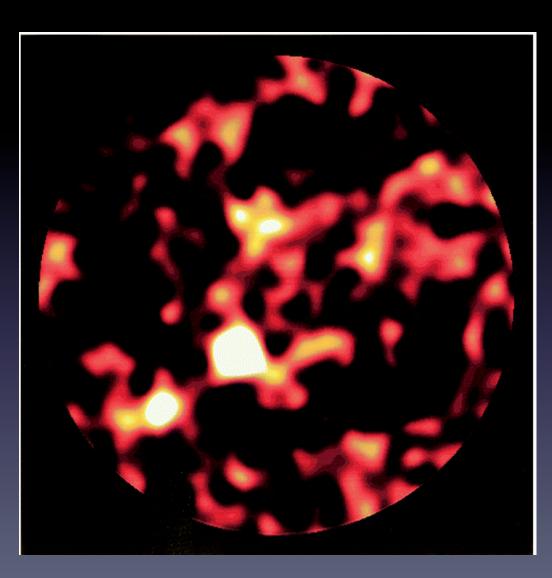
Negative K-correction



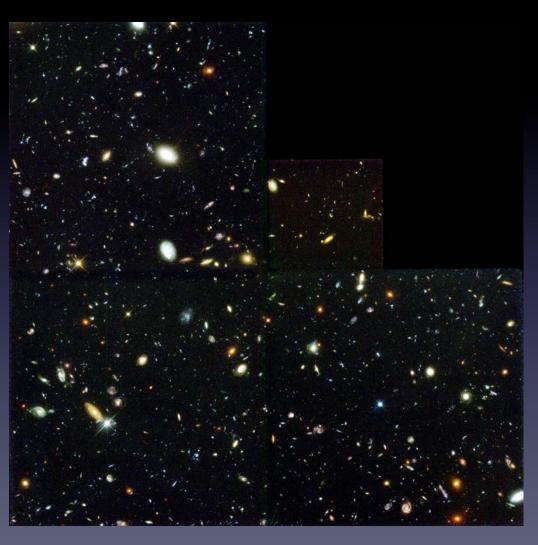
SCUBA



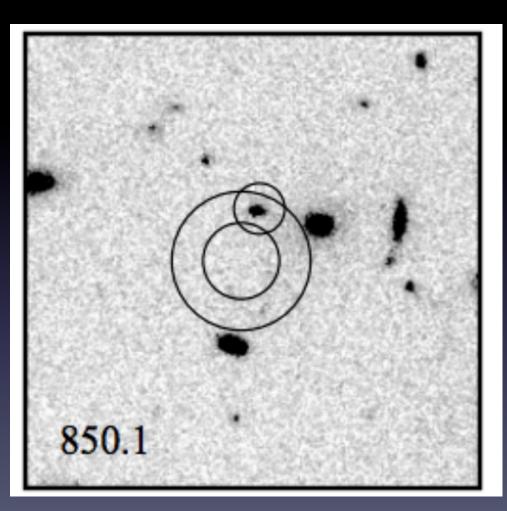
- SCUBA is the Submillimetre Common-User Bolometer Array
- Operated simultaneously at 450 and 850 μm, with 91 and 37 pixels, respectively.
- When starting operation, it was 20 times more sensotive than its nearest competitor.



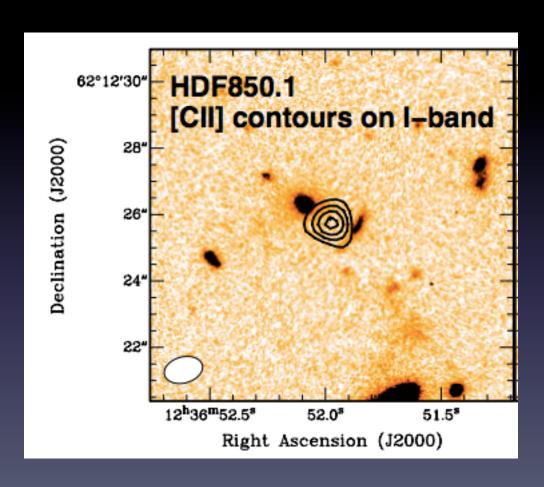
- SCUBA/JCMT 850µm map in the Hubble deep field (HDF).
- HDF850.1 is the brightest submm source in the survey.



- SCUBA/JCMT 850µm map in the Hubble deep field (HDF).
- HDF850.1 is the brightest submm source in the survey.
- The redshift of HD850.1 is 5.2 and SFR=850M_{sun}/yr (Walter+2012)



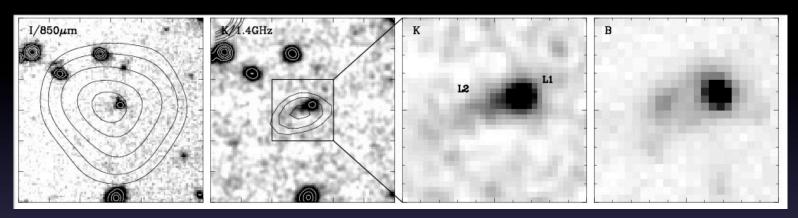
- SCUBA/JCMT 850µm map in the Hubble deep field (HDF).
- HDF850.1 is the brightest submm source in the survey.



- SCUBA/JCMT 850µm map in the Hubble deep field (HDF).
- HDF850.1 is the brightest submm source in the survey.
- The redshift of HD850.1 is
 5.2 and SFR=850M_{sun}/yr
 (Walter+2012)

Redshift Distribution

SMM J02399-0136

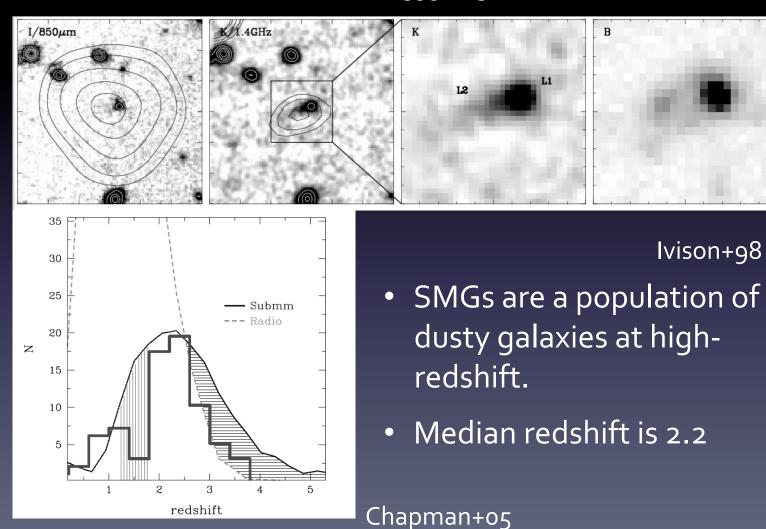


Ivison+98

- High Spatial resolution VLA 1.4GHz continuum image to determine position.
- Optical or radio line follow-up to determine redshift or photometric redshift.

Redshift Distribution

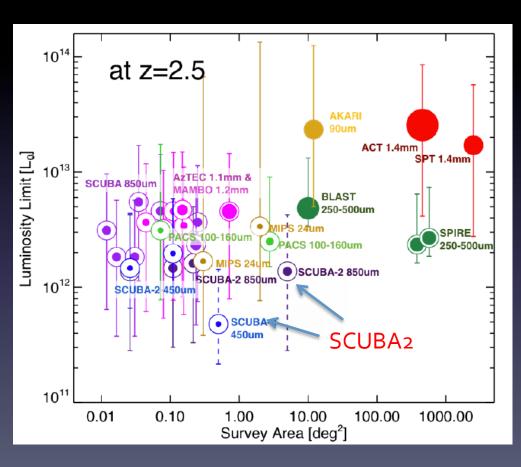
SMM J02399-0136



Properties of SMGs

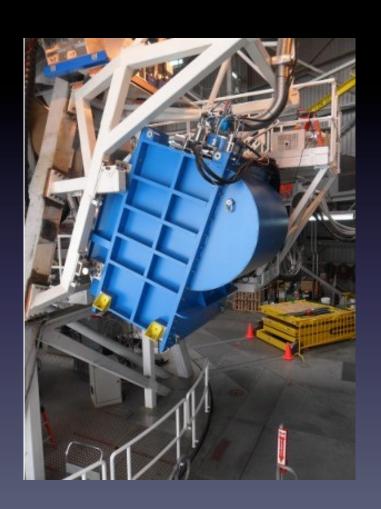
- Rare: Volume number density ~ 10⁻⁵ Mpc⁻³
- Massive: Stellar masses ~ 10¹¹M_{sun}
- Star-forming: SFR >500 M_{sun} yr⁻¹
- Dusty: dust mass 10⁸-10⁹ M_{sun}
- Halo mass: 10¹³ M_{sun}

SMG Surveys



- In the last decade, a large sample of SMGs have been established, using SCUBA, MAMBO, AzTEC, SPT, SPIRE and PACS etc.
- In particular, SCUBA2
 allows survey to achieve
 both deep and wide.

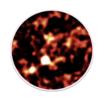
SCUBA₂



- Simultaneous imaging at 850 μm and 450 μm with large field of view (45 arcmin²)
- Mapping the submm sky
 ~1000 times faster than
 SCUBA

15 years ago... our first glimpse, with SCUBA

The HDF

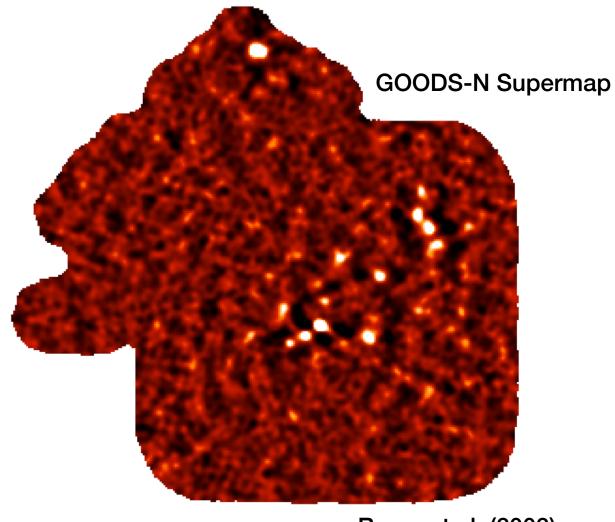


Hughes et al. (1998)

courtesy of Jim Geach

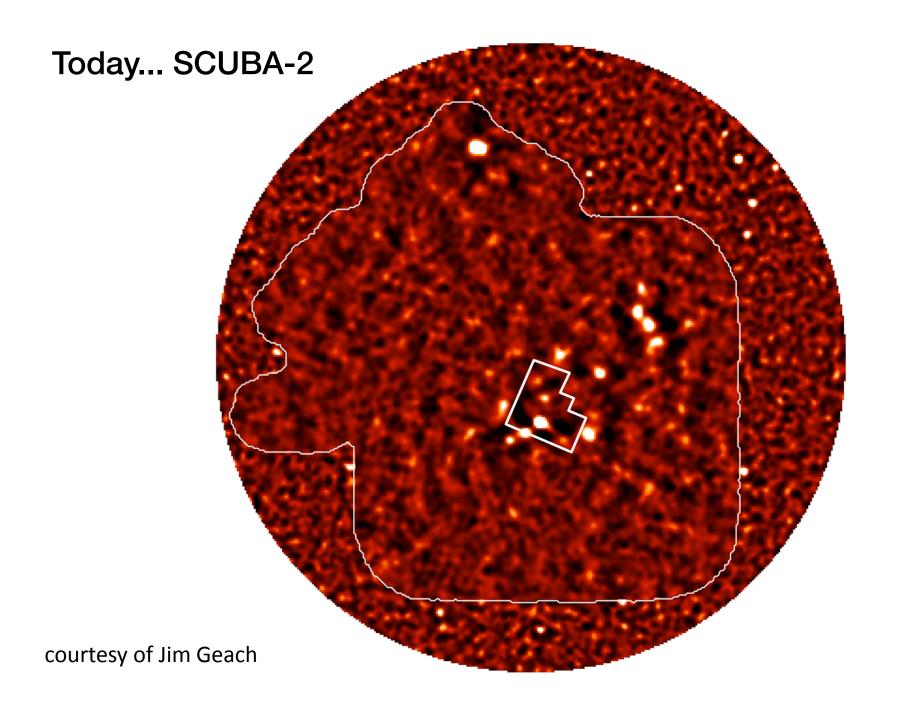
Also: SCUBA surveys of strong lensing clusters, Smail et al. (1997)

10 years ago... 'proper' surveys

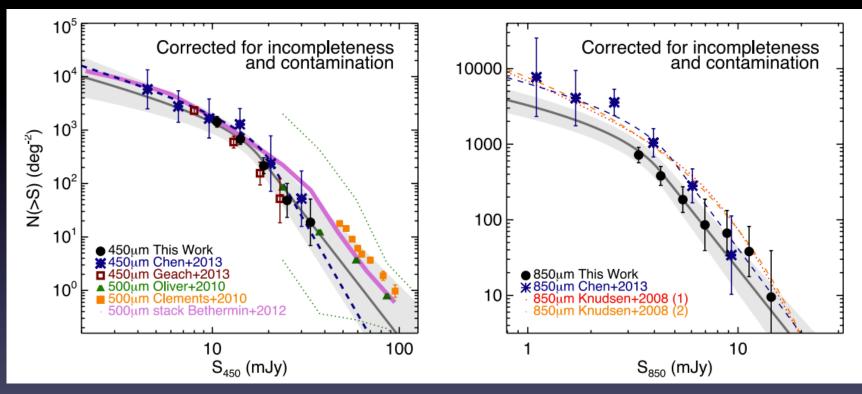


courtesy of Jim Geach

Borys et al. (2003) Pope et al. (2005)

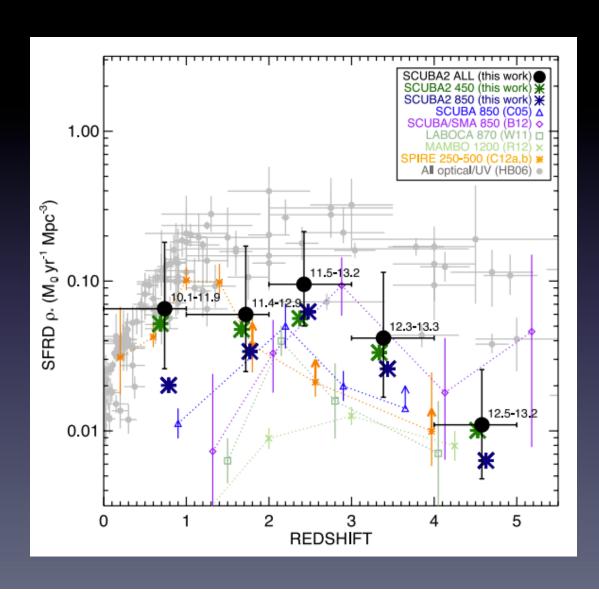


SMG Number Counts

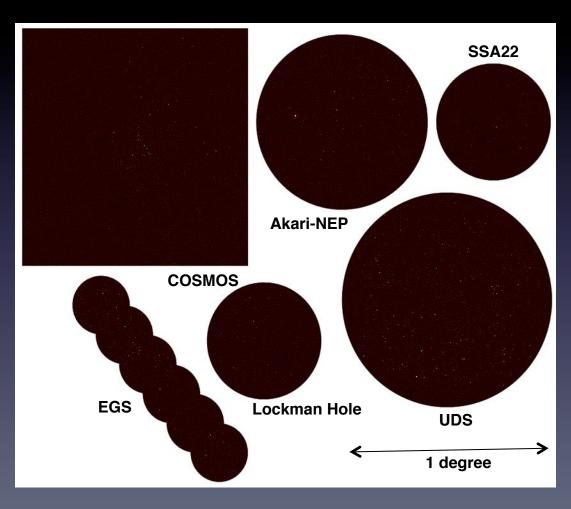


- SMG accumulate number counts in 450µm and 850µm in Casey+13
 a part of COSMOS field using SCUBA2.
- Multi-wavelength: better constraint on FIR SED reliable photometric redshift

Cosmic Star-formation History



SCUBA-2 Cosmology Legacy Survey

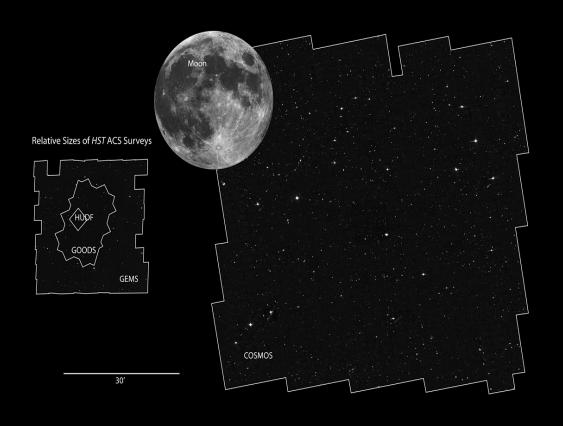


- 1778 hours JCMT legacy survey
- Two tier strategy covering well-know field
- Wide survey: 10 deg²,
 1.2mJy at 850μm
- Deep Survey: 0.25 deg²,
 1.2mJy at 450µm

Science drivers

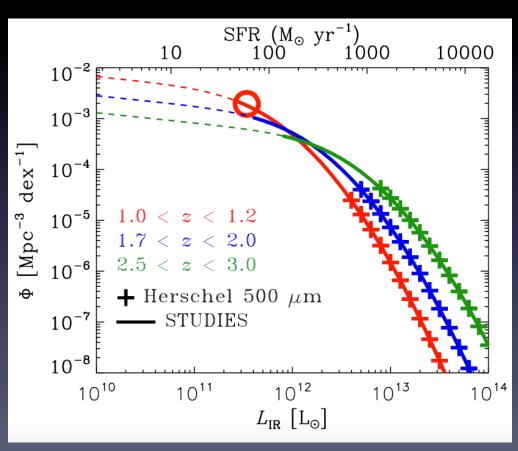
- Comprehensive survey of SMGs detected at 850um (N~10⁴) over several ~degree scale fields
 - measure clustering / halo properties
 - constrain bright-end counts
 - reliably probe properties of fainter SMGs (3-5mJy)
 - ► AGN/starburst connection, transition objects
 - 'very' high-z science (IR luminosity density at z>5, rare objects)
 - cross-correlation studies
- 'Keyhole' confusion limited survey at 450+850um
 - properties of 450um-selected SMGs below HSO confusion limit
 - resolving CIB at 450um (closer to peak)
 - probing very faint end of counts (P(D) analyses)
 - joint stellar/dust/gas morphologies (CANDELS overlap, ALMA)

S2-COSMOS



- A SCUBA-2 survey of 1,000 SMGs in the COSMOS field
- Uniform 1.2mJy rms map of the full COSMOS
- Superb supporting multiwavelength data

STUDIES: SCUBA-2 Ultra Deep Imaging EAO Survey



- Ultra deep, confusion limited 450 µm image in the COSMOS-CANDELS region
- The deepest 450 µm image ~ √2× deeper than the S2CLS
- Triple the area previously achieved by S2CLS in the COSMOS field