



Mopra Ideas for Future Operations

Turning challenge into opportunity

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In times of turmoil, great opportunity looms

- Snapshot of my private brain
 - This is **not** a CSIRO proposed/endorsed perspective!
- Opportunities for improvements in
 - Management (operations/facility)
 - Software
 - Hardware

Management aspects

- Operations
 - Observation planning
 - Maintenance planning (might fit in with CSIRO's resource planning?)
 - Execution of observations

Software

- Improving operations software
 - Queue mode
 - Web frontend
 - Data quality monitoring
- Some already available, others easily made available

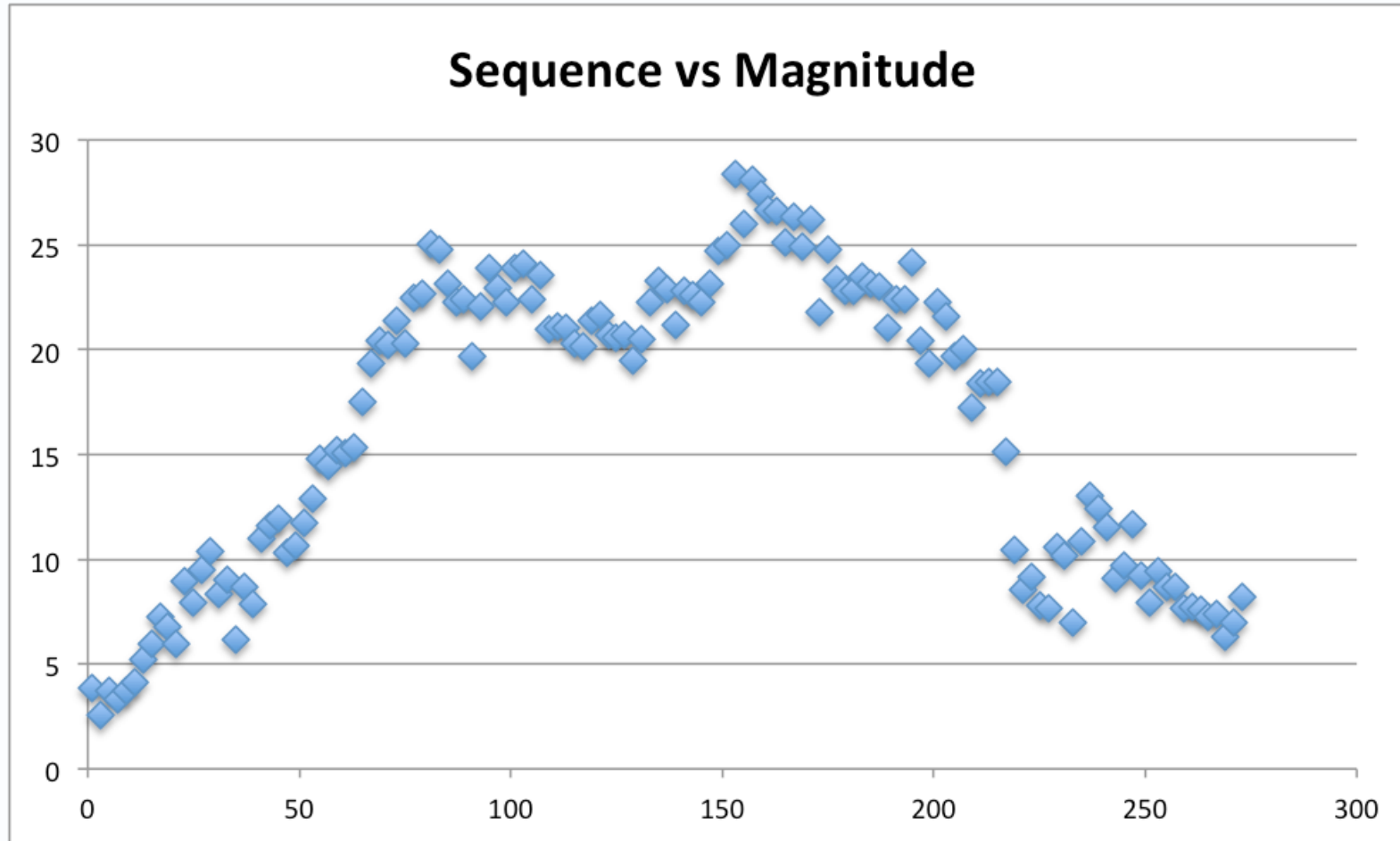
Software: Expanding Queue Mode 1

- Already able to do much more than you know:
 - Autopoint (c.f. pointing)
 - M999, VY_CMa, J2000, 07:22:58.33, -25:46:03.00, pointing.sch,PNT 6
 - Executes pointing until residuals better 6 arcsec
 - Stowing the antenna
 - M999, , , , ,NULL,STOW
 - Waiting for a number of seconds (or until a UT)
 - M999, , , , ,NULL,WAIT 2015-12-12 22:33:44
 - M999, , , , ,NULL,WAIT 60
 - Setting the focus
 - M999, , , , ,NULL,FOCUS 16.8

Interlude: On the topic of pointing...

- Followed AH Sco at 3mm and UU For at 7mm across the sky
- Results for 3mm:
 - El < 40 deg: Pointing once an hour OK
 - El >= 40 deg: Pointing optional for up to ~3.5h
- Results for 7mm:
 - Pointing may be entirely optional.
- **Must** start with uncorrected pointing mode, e.g. “newcain pparam”!

Interlude: On the topic of pointing...



Software: Expanding Queue Mode 2

Automated scheduling modes

- Long Term Predictive scheduling
- Short Term Predictive scheduling
- Reactive scheduling

Software: Expanding Queue Mode 2

Automated scheduling modes

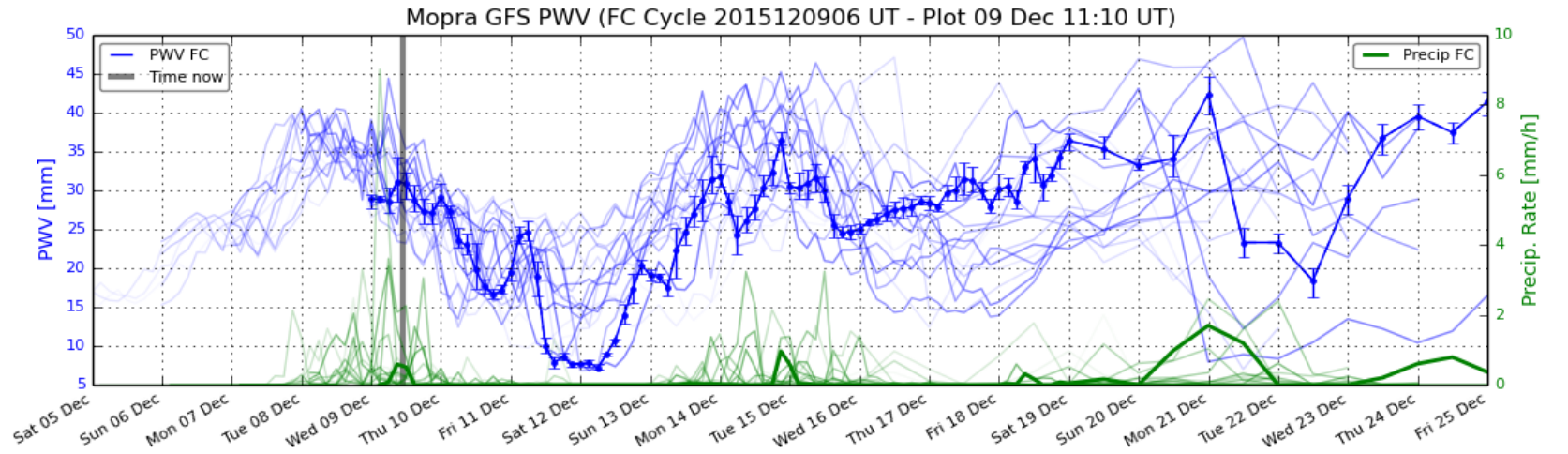
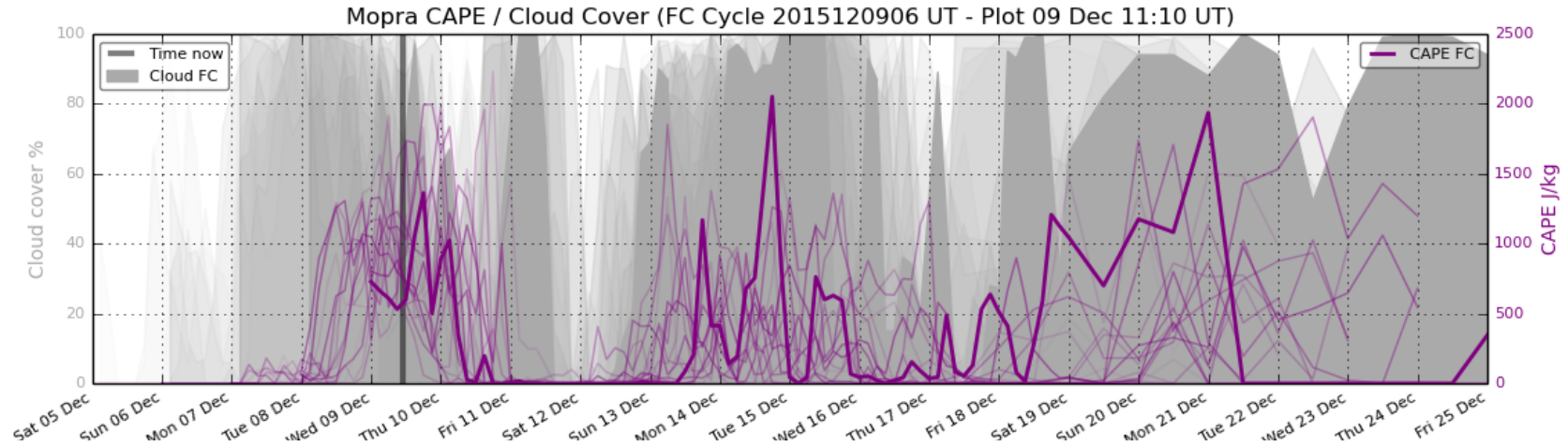
- Predictive scheduling:
 - Climate considerations
 - Past years Tsys
 - Atmospheric conditions
 - No CO scheduled with more than 20mm PWV forecast/calculated
 - (We already have GFS forecast data, and the presently calculated data)
 - Source elevation / transit times
 - Optimise maps for ideal start times resulting in as much high elevation observing time as possible
- Reactive scheduling:
 - System monitors actual conditions and aborts e.g. >1000K CO observations and substitutes a 7mm or 15mm project

Software: Expanding Queue Mode 2

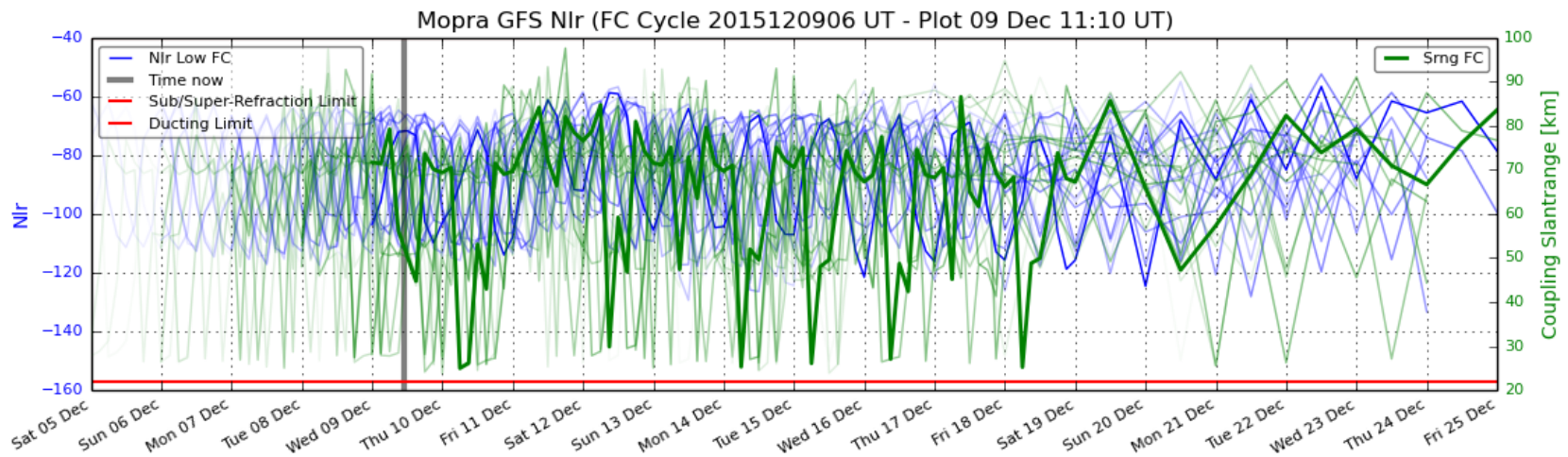
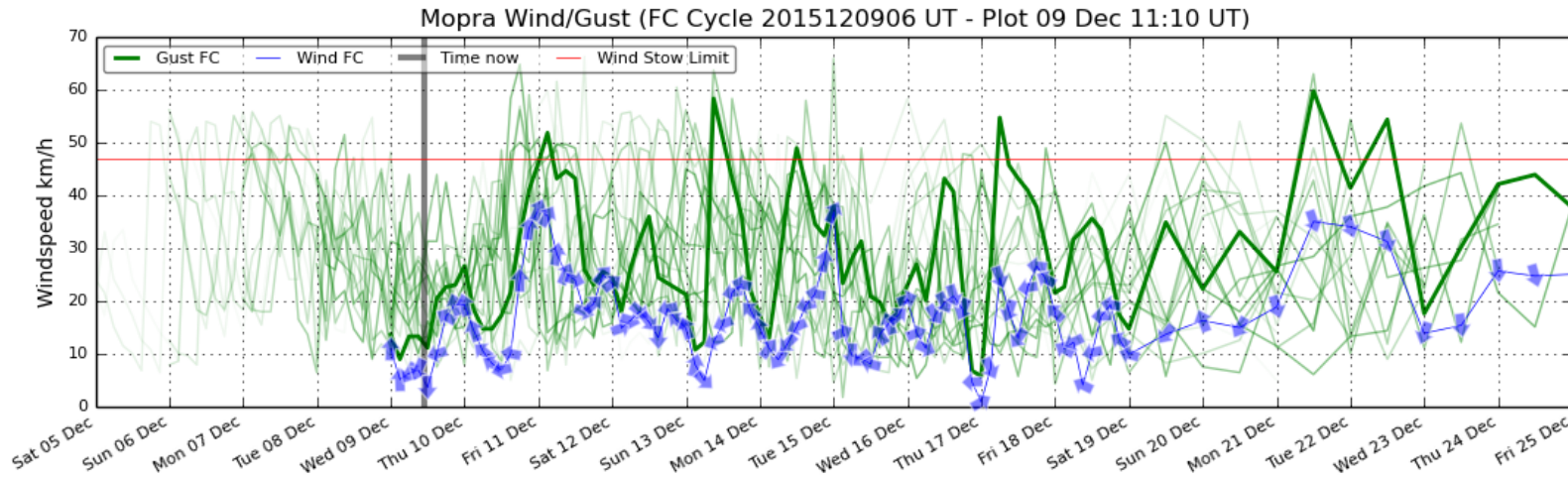
Automated scheduling modes

- Long Term Predictive scheduling
 - Manual process, based on statistical analysis
- Combined manual/automatic approach initially, fully automated after first semester
 - Short Term Predictive scheduling
 - Reactive scheduling

Predictive scheduling: Short term GFS



Expanding Queue Mode 2



<http://wxserver.inside.net>

Expanding Queue Mode 2

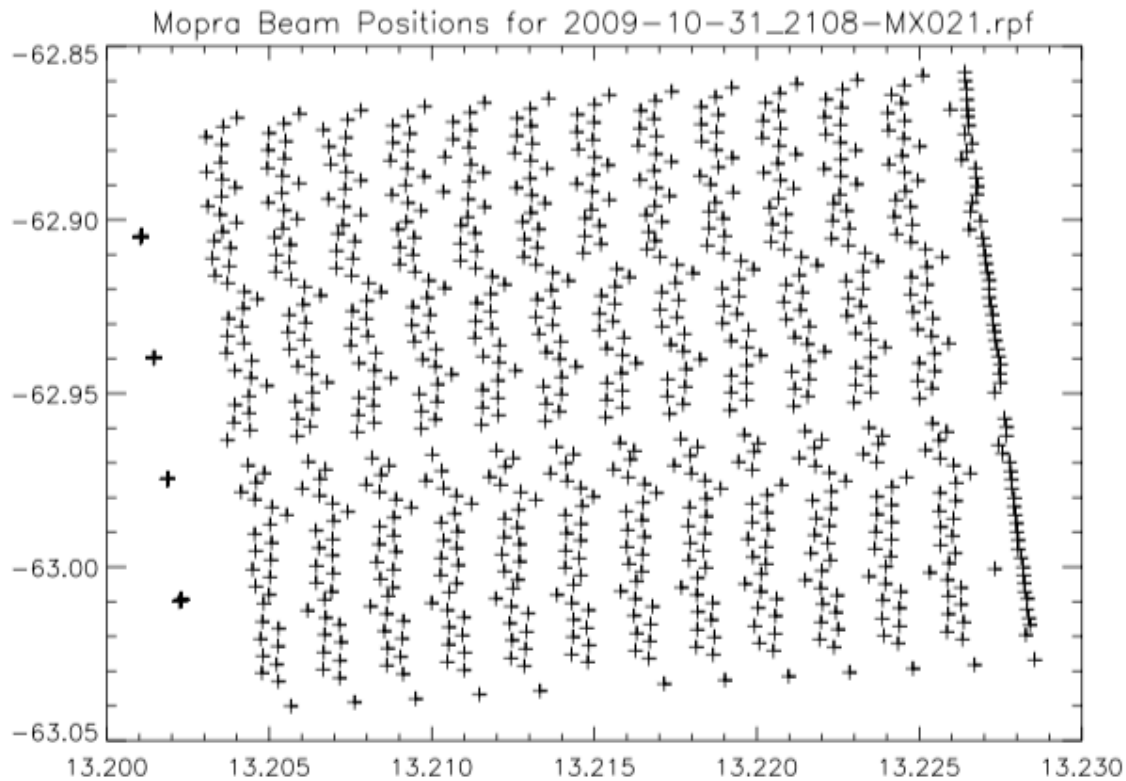
- Web interface
 - Queue files were never meant to be manually generated
 - But perhaps that's preferable and more flexible as all users really are “power users”?
- Leverage schedule file generators built by existing Mopra teams
- Queue generator considers predictive criteria when scheduling
- Queue manager considers reactive criteria while observing
 - Can abort/reschedule/switch campaign
- This mechanism largely implemented many years ago for ThRUMMS adaptive undersampling technique

Data pipeline

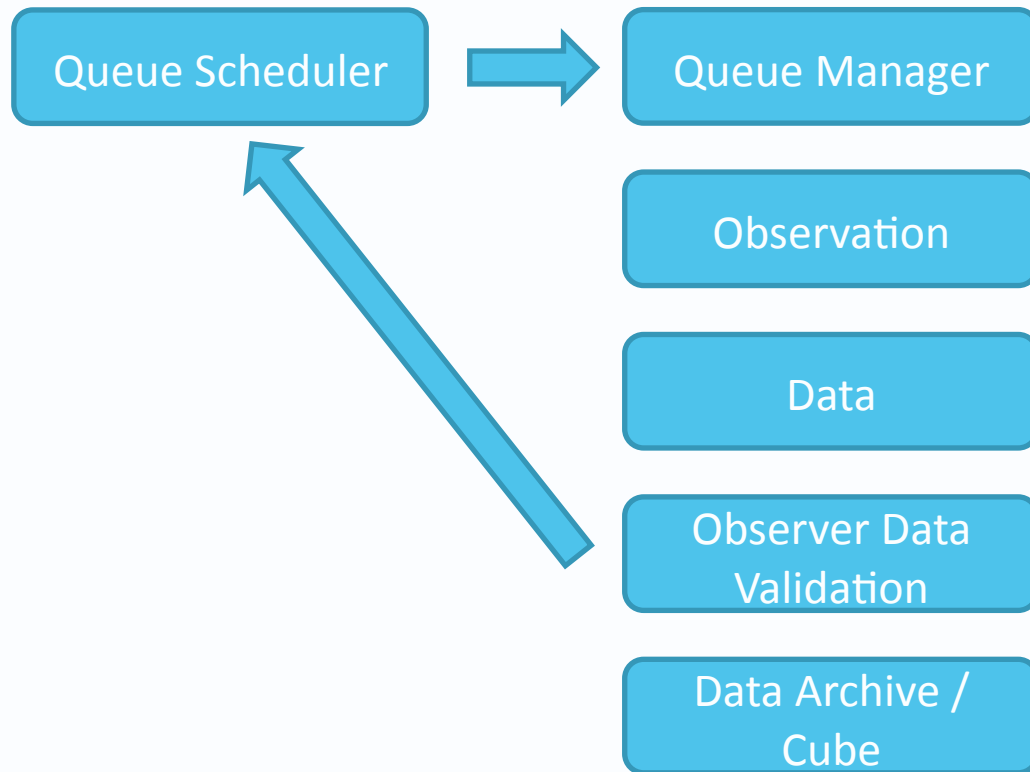
- More automation
- Observer intervention only when something goes wrong
- Data validation
- Requires survey style project
- Well defined outcome metrics
 - Otherwise, waste of time!
- Sensible and cheap approach could be to observe automatically and accept a certain percentage of waste.
- Reobserve what didn't work

Data pipeline

- No point in developing quality tools if nobody looks at them:



Observing Summary



Future perspectives

- Mopra financial impasse might be temporary
- MOPS/MMIC tremendously future proof investments
- But, now is the time to look at what can be done to
 - Retain the forefront position
 - Future proof the facility for when funding returns
- Perhaps optimistic...
- Burton 2009 upgrades paper:
 - Fast mapping (since implemented): Improved mapping speed x 8
 - Multibeam
 - UWB Correlator

Hardware 1: Multibeam

- 7 element Multibeam (Burton 2009)
- Close packed hexagonal arrangement could work
 - CSIRO Concept design study from 2008
- Increase mapping speed x 7
- Requires major improvements:
 - backend (7 x DFB processing)
 - receiver system x 7
- Expensive: On order \$ >2-3 m

Hardware 2: Backend

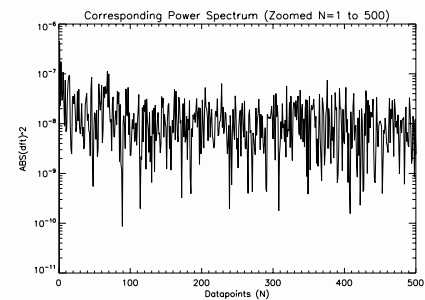
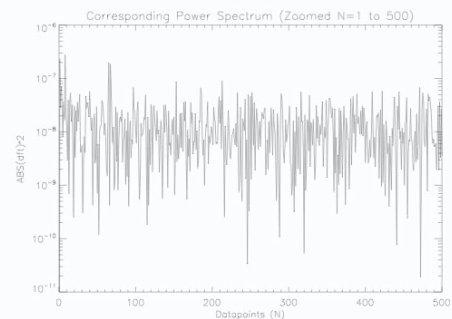
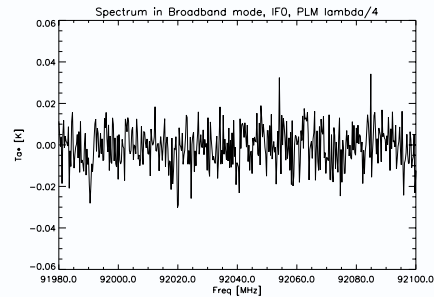
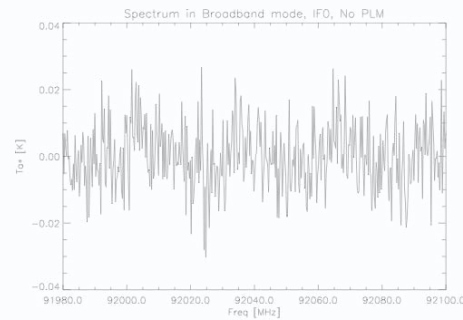
- UNSW MOPS = Mopra Spectrometer
 - DFB = digital filter bank
 - Not a correlator. A correlator correlates wave functions from separate signal sources (e.g. an interferometer). There's only one signal source in Mopra.
 - Runs a sequence of FFTs on a stream of data, ~500ks/s
 - Weighting function is applied to each segment to give the shape of the frequency response
 - Computational problem that has massively parallel solutions
 - UNSW MOPS built in early 2000s, operational since 2007, entering 9th year of operation
 - With 8 GHz BW still the world's widest instantaneous BW (corrections?)

Hardware 2: UWB DFB, MUSE

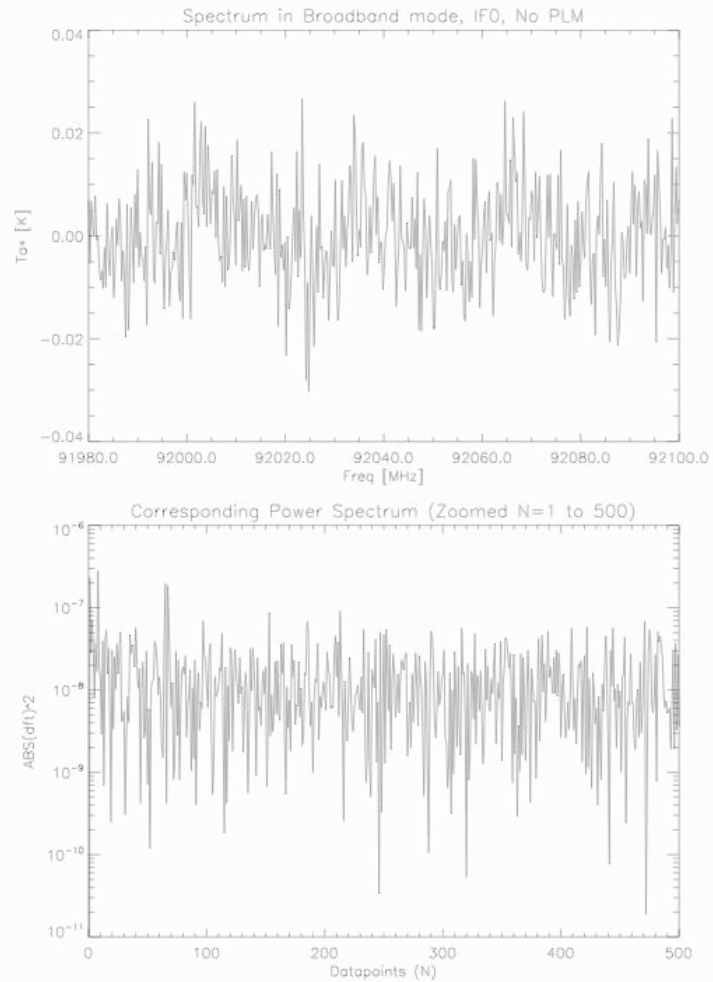
- Proposal for a 48 GHz DFB:
 - MUSE (Mopra Ultrawideband SpEctrometer) e.g. 69 – 117 GHz
 - Requires upgrading of IF to match BW, currently 4-12 GHz
 - 2 – 50 GHz, RoF implications, may require multiple separate downconversions at e.g. 2 - 12 GHz
 - Receiver system and LO(s): As long as within the already designed freq ranges OK
 - DFB implementation in software (harnessing e.g. CUDA/GPU computing)
 - Existing channel resolution must remain, 0.1km/s in zoom modes, 1 km/s in broadband mode
 - Data throughput will require parallel processing
 - Fraction of the HW cost of MOPS
 - Data rates high, but manageable: ~ 8 x the data rate of MOPS if spectral resolution is retained.
- **\sim \$300k project**

Hardware 3: Standing Wave Mitigation

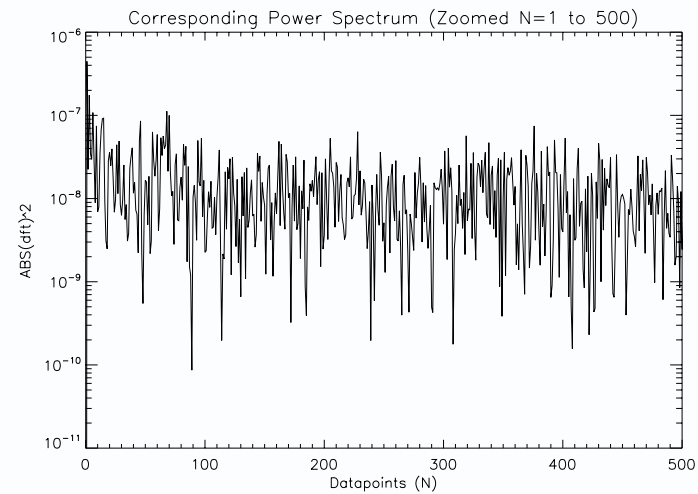
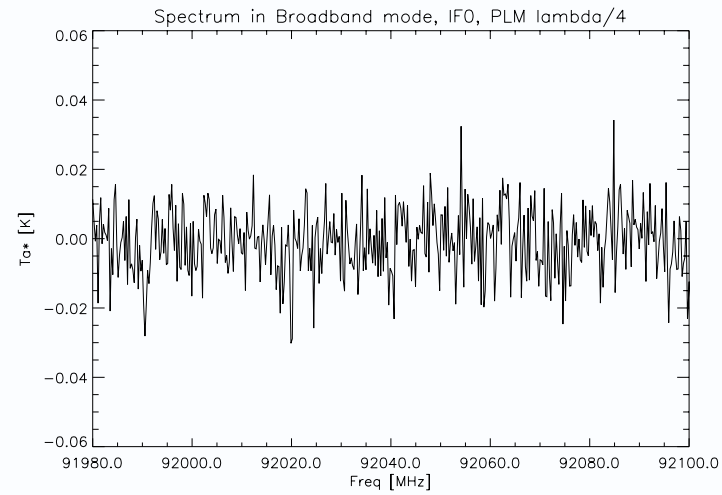
- Currently no deep observations are possible. Extragalactic science not really feasible.
- “Defocusing” experiments using the subreflector (Indermühle 2010)



Hardware 3: Long integration

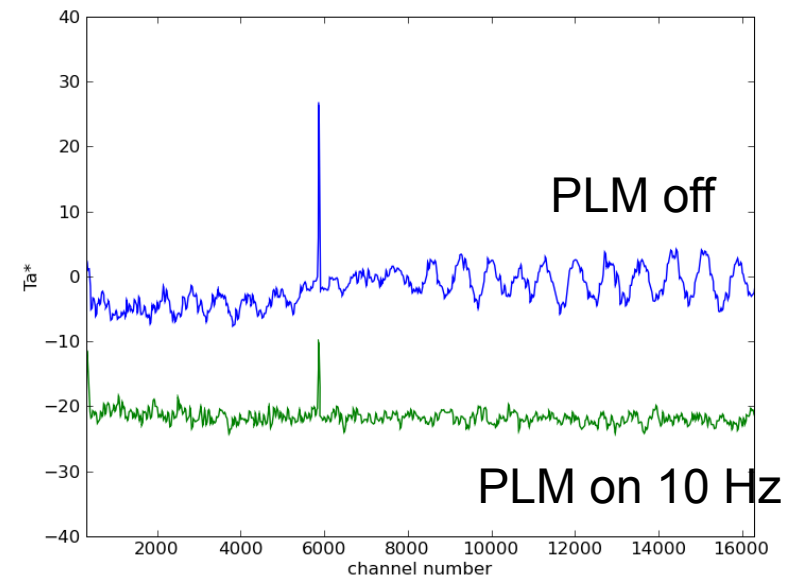
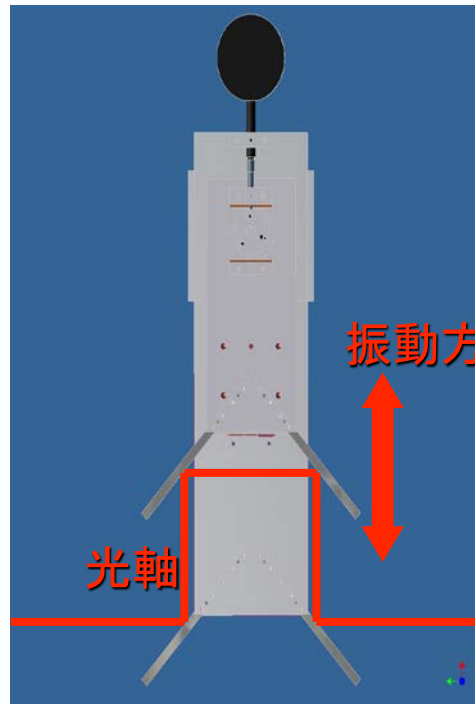


Hardware 3: Long integration



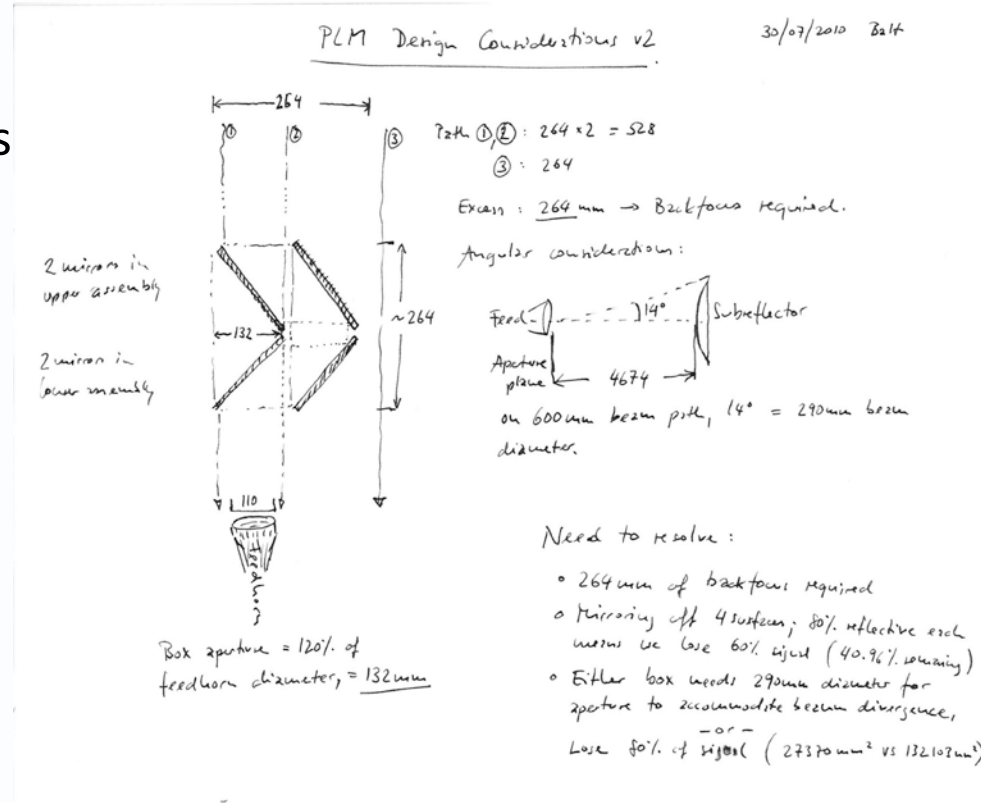
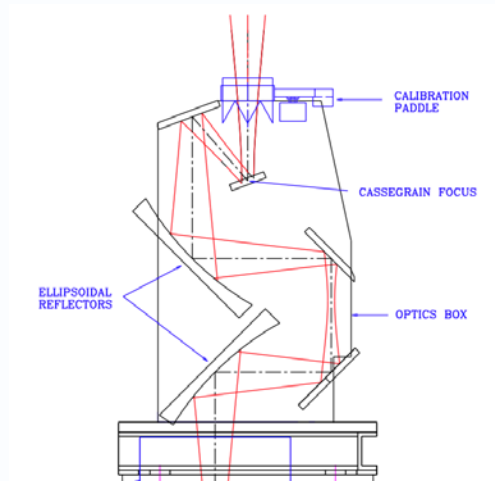
Hardware 3: Standing Wave Mitigation

- Nagoya 1.85m telescope experiment (2009)



Hardware 3: Standing Wave Mitigation

- For Mopra
 - Optical design considerations



- Good student project, collaboration with engineering?
- <\$100k project

A circular fisheye photograph of a night sky. The Milky Way galaxy is visible as a bright, hazy band of light stretching across the sky. In the lower right corner, a large satellite dish antenna is visible, illuminated from below. The sky is filled with numerous stars, and the overall scene is dark with some ambient light from the antenna and surrounding structures.

Thank you

Dr Balthasar Indermühle