

INVERSION FOR MERIDIONAL FLOW IN THE SOLAR INTERIOR

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- Rajaguru & Antia (2015) used 4 years of data from HMI
- The travel-time difference between N–S and S–N along a ray path Γ_0 is given by

$$\delta\tau = -2 \int_{\Gamma_0} \frac{\mathbf{u} \cdot \hat{\mathbf{n}}}{c^2} ds$$

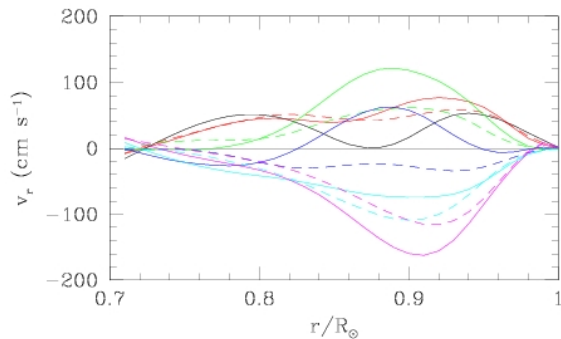
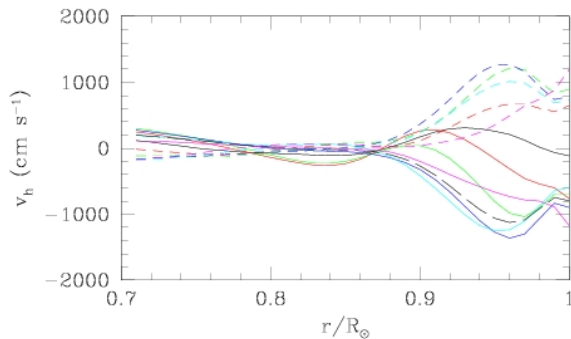
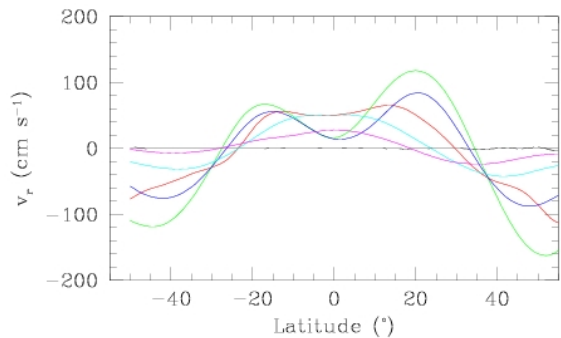
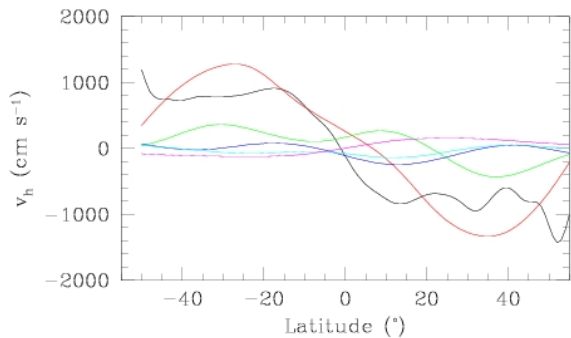
- Stream function was used to represent meridional velocity:

$$\rho u_r = \frac{1}{r} \frac{\partial \psi}{\partial \theta} + \frac{\cos \theta}{r \sin \theta} \psi$$

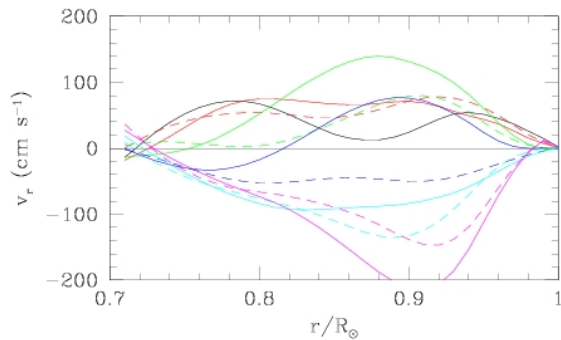
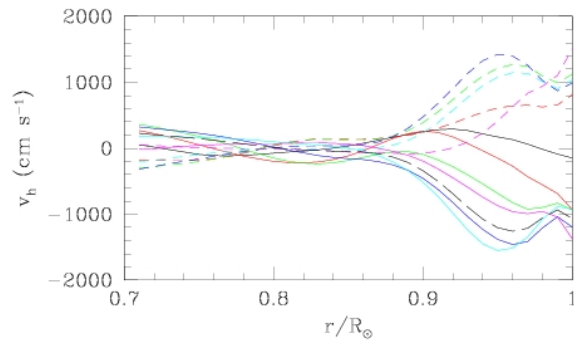
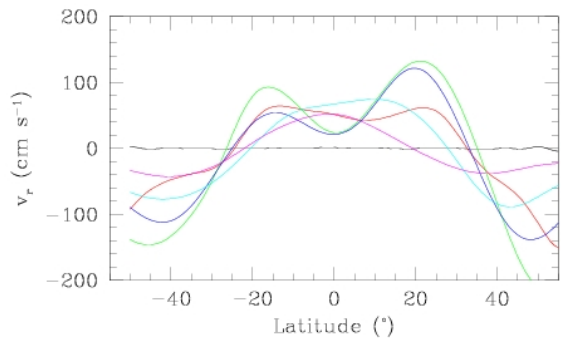
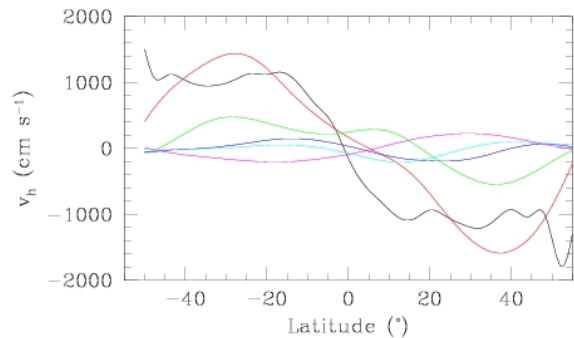
$$\rho u_\theta = -\frac{\partial \psi}{\partial r} - \frac{\psi}{r}$$

which ensures that the solution automatically satisfies the continuity equation and also enables u_r to be determined.

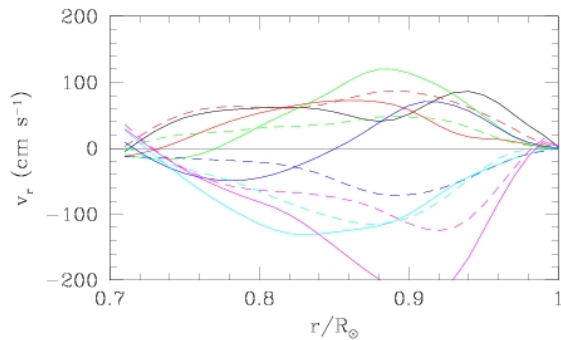
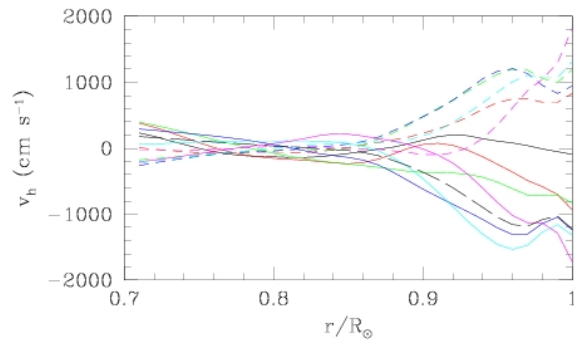
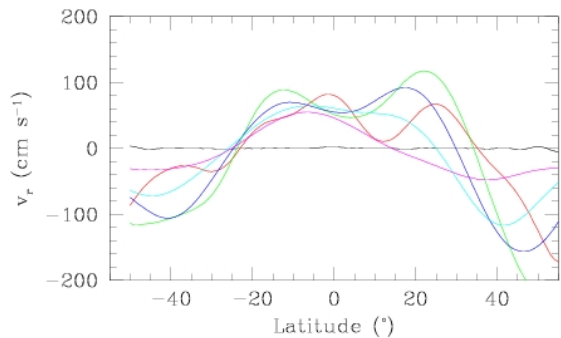
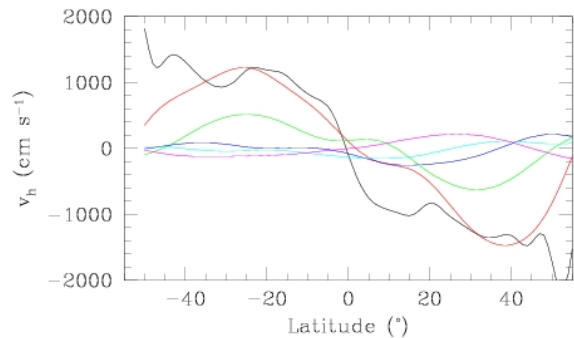
- Inversions are done using the Regularised Least Squares (RLS) technique.
- Boundary condition of $u_r = 0$ at outer boundary is used
- To get more realistic error estimates on inverted velocity we also perturb the smoothing used (apart from travel times) in inversion which gives some estimate of systematic errors in inversions.
- Results using travel times calculated for different frequencies using 6-year HMI data are as follows:



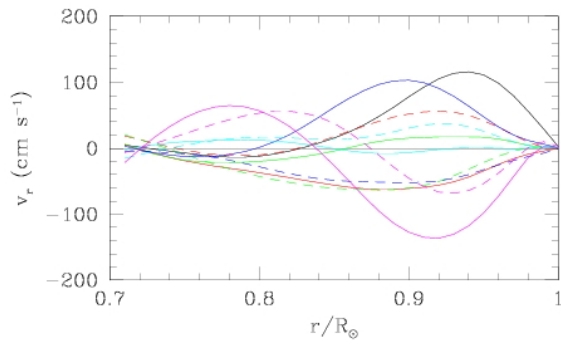
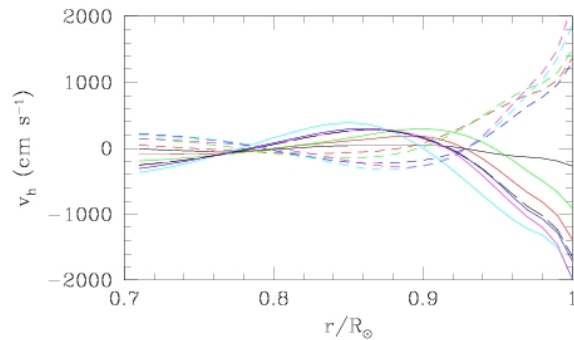
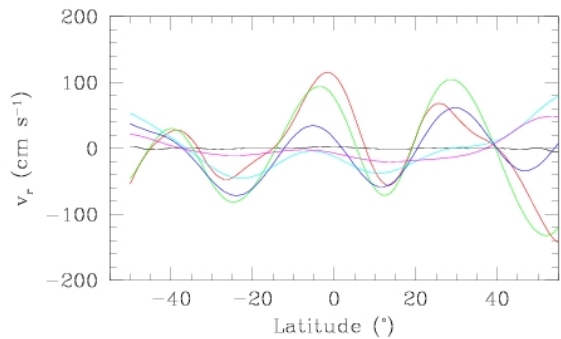
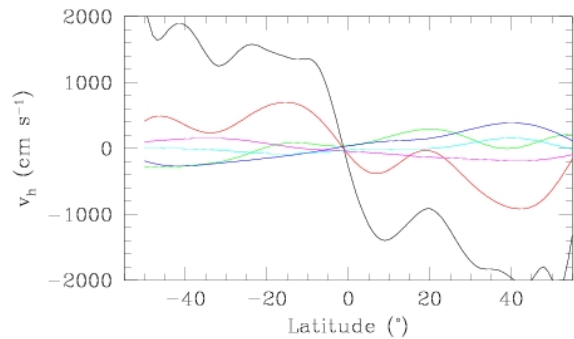
$\nu = 2.7$ mHz



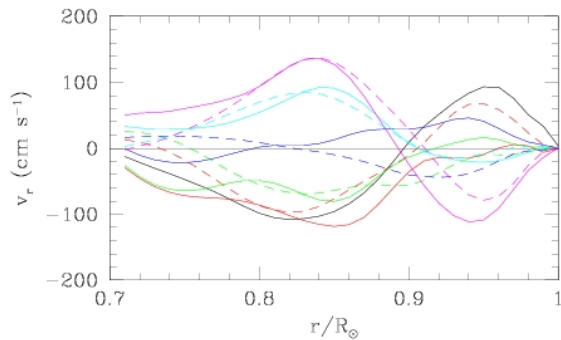
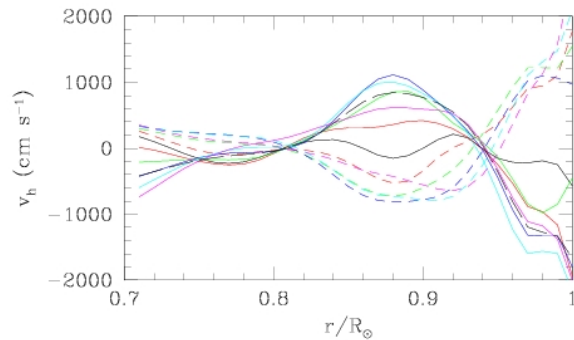
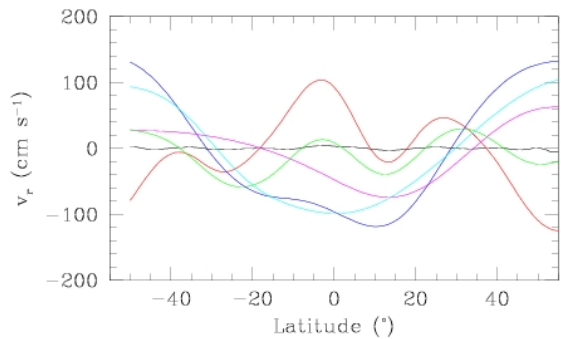
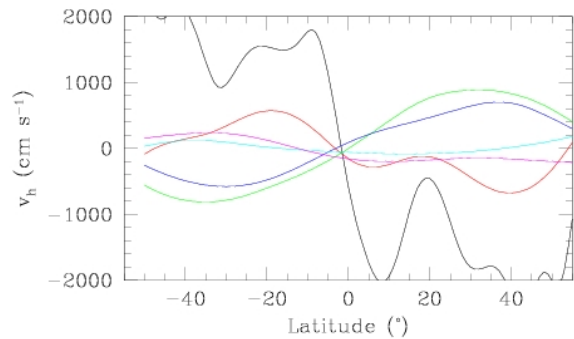
$\nu = 3.0$ mHz



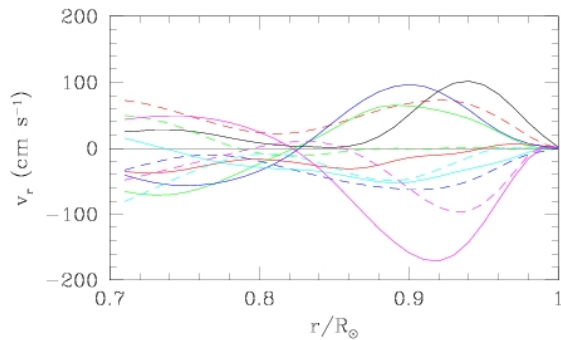
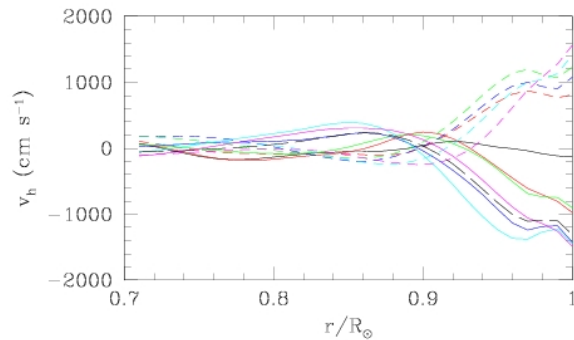
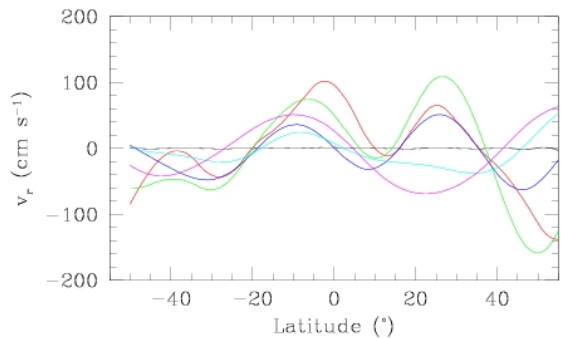
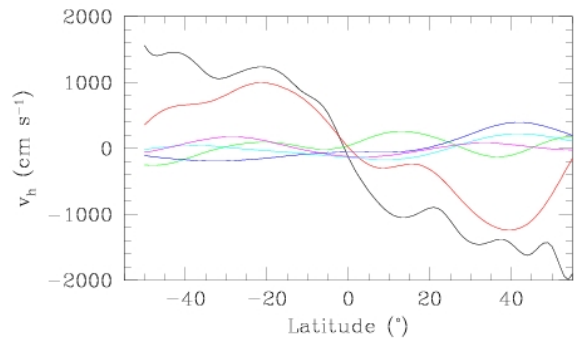
$\nu = 3.3$ mHz



$\nu = 4.0$ mHz



$$\nu = 4.5 \text{ mHz}$$



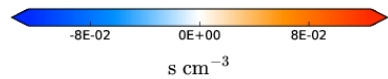
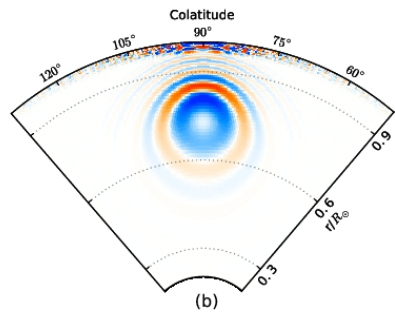
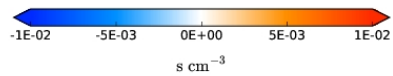
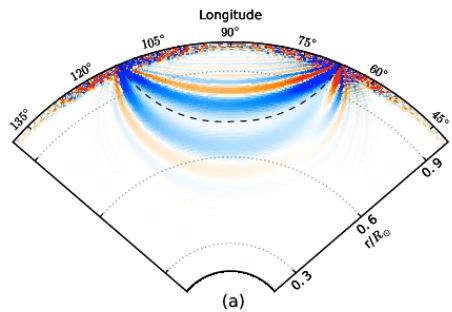
all frequencies combined

Sensitivity Kernels using Born Approximation

- Hanasoge and Mandal (2017) have calculated the sensitivity kernels using Born Approximation. The difference in travel times due to flows is defined by

$$\delta\tau(\mathbf{x}_r, \mathbf{x}_s) = \int_{\odot} d\mathbf{x} \mathbf{K}_v(\mathbf{x}_r, \mathbf{x}_s, \mathbf{x}) \cdot \delta\mathbf{v}$$

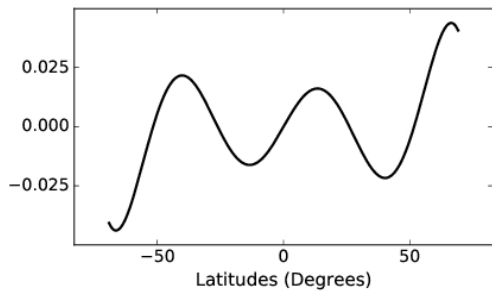
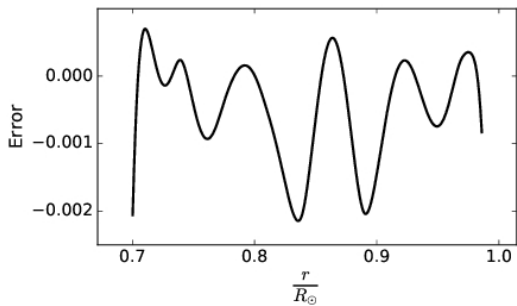
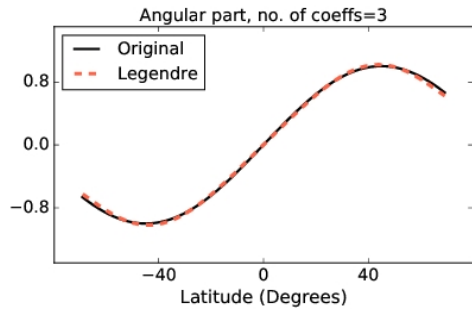
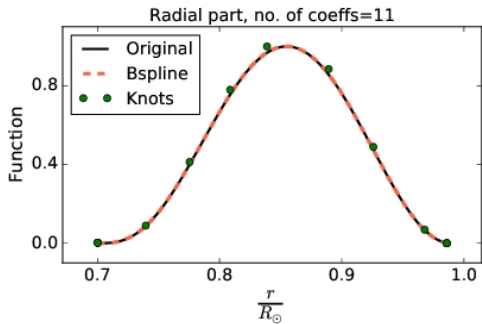
These have been tested for the case of uniform rotation. For the v_ϕ component a sample kernel is

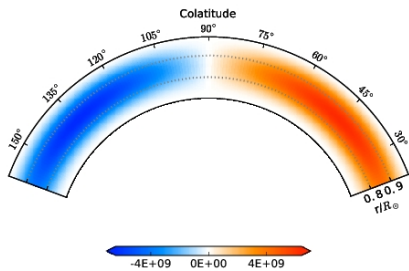


- For meridional flow they use the stream function χ . To invert for the meridional flow the stream function is expressed as

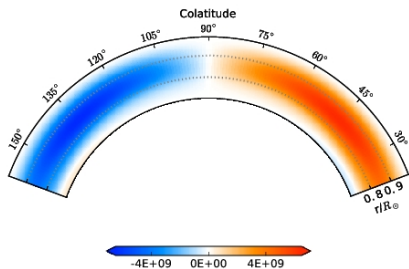
$$\chi(r, \theta) = \sum_{i,j} a_{ij} B_i(r) P_j(\cos \theta)$$

with 11 knots to define B-splines in r and 3 Legendre polynomials. Using a simple 1-cell and 2-cell meridional flow profile the travel times were calculated and these were inverted to test the process. No errors are added in these calculations.

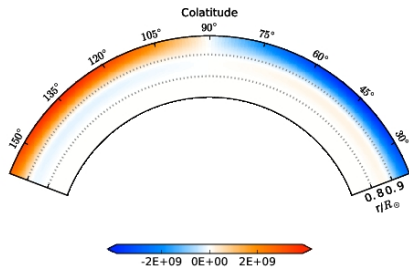




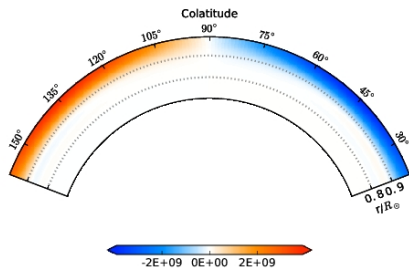
Input profile



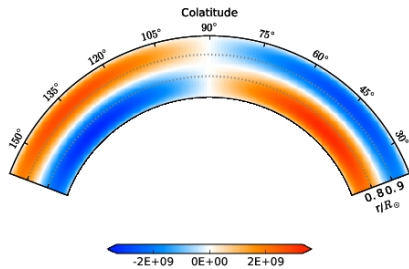
Inverted profile



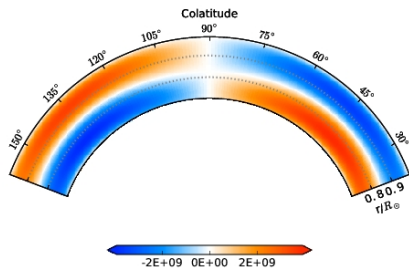
Input profile



Inverted profile



Input profile



Inverted profile