

# **Inverse transfer of Magnetic Energy in a decaying MHD system**

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# ■ Large & small scale B field in space

## • Plasma & Magnetic field in space

Their mutual interactions, ubiquitous phenomena:

→ Evolution of B-field, constraint of Plasma

## • Role of B-field in plasma?

- Lorentz force constrains plasma

- Stabilization or destabilization of a plasma system  
Pinch, Kruskal-Schwarzschild instability...

- Transport of angular momentum (axisymmetric system)  
Braking fast spinning collapse or making accretion continue

- Source of kinetic & thermal energy  
Magnetic reconnection

-...

※ But, B-field is not a prerequisite for the evolution of a plasma system

# ■ Large & small scale B field in space

## • Origin of B-field?

### - Cosmological model (Primordial & Astrophysical model)

#### 1. Primordial model

a. When conformal invariance of EM fields was broken.  
(Inflation, Turner & Widrow 1988)

b. Through cosmic phase transition  
Electro-Weak Phase Transition (EWPT),  
Quantum Chromo Dynamics transition (QCD)  
(Grasso & Rubinstein 2001)

#### 2. Due to Plasma fluctuations (Astrophysical model)

Biermann battery (Biermann 1950), Harrison effect (Harrison 1970)

$$\rightarrow B \sim 10^{-62} G - 10^{-19} G$$

### - Observed B-field?

$$\rightarrow B \sim 10^{-5} G$$



**Dynamo**

# ■ Large & small scale B field in space

- Requirements of LSD & SSD?

- LSD

- Helicity, Differential Rotation, Magneto-Rotational Instability  
→ External forcing source

- SSD

- No instability, no shear → Forced with nonhelical field

- Question

1. Is LSD impossible without these specific forcing sources?

- Dynamo theory: Impossible

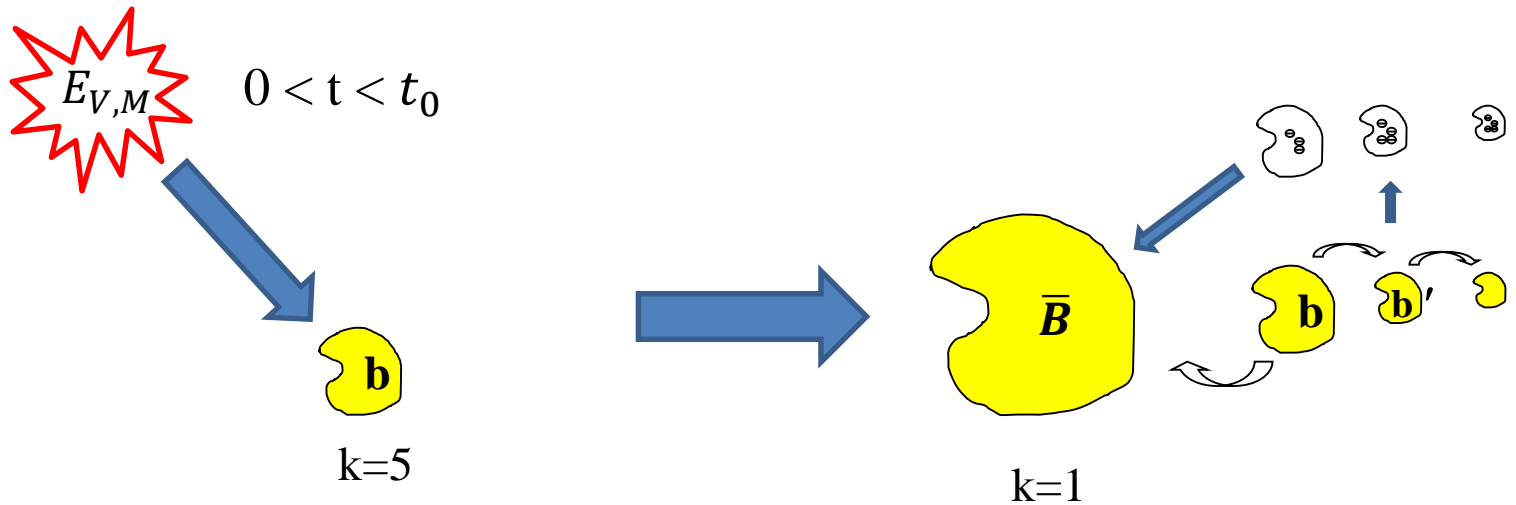
- But some theories and simulations suggested and reported the possibility of LSD in a decaying MHD system.

- Olesen 1997, Ditlevsen et al. 2004, Brandenburg et al. 2015, 2017, J. Zrake 2014, 2016, Park 2017a, 2017b

2. How can we explain LSD?

# ■ Energy Spectrum in a decaying MHD system

- Simulation setting

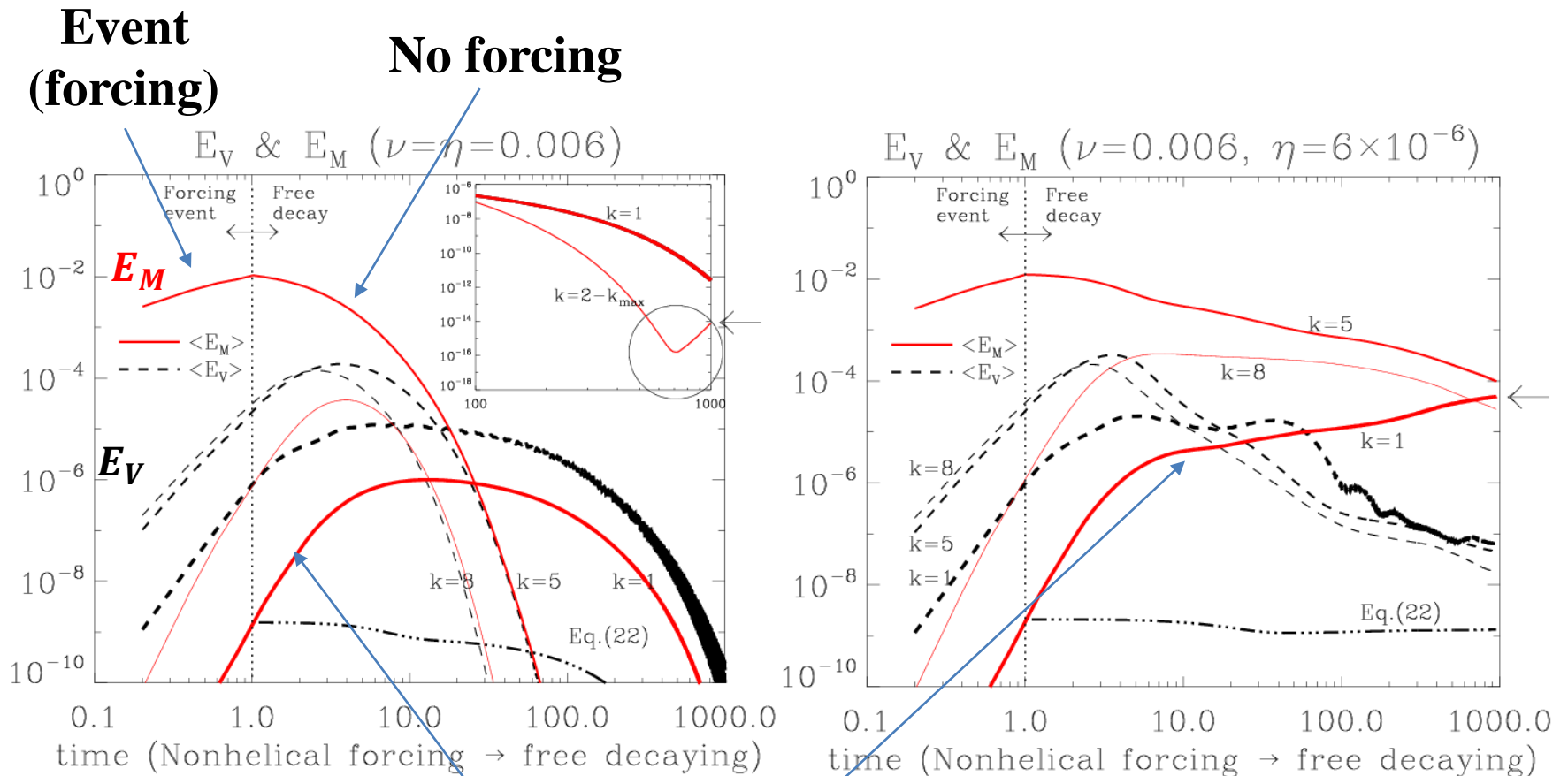


Ephemeral supply of  $E_V$  &  $E_M$

Decaying (e.g. Supernovae etc.)

# ■ Energy Spectrum in a decaying MHD system

- **NonHelical** magnetic energy

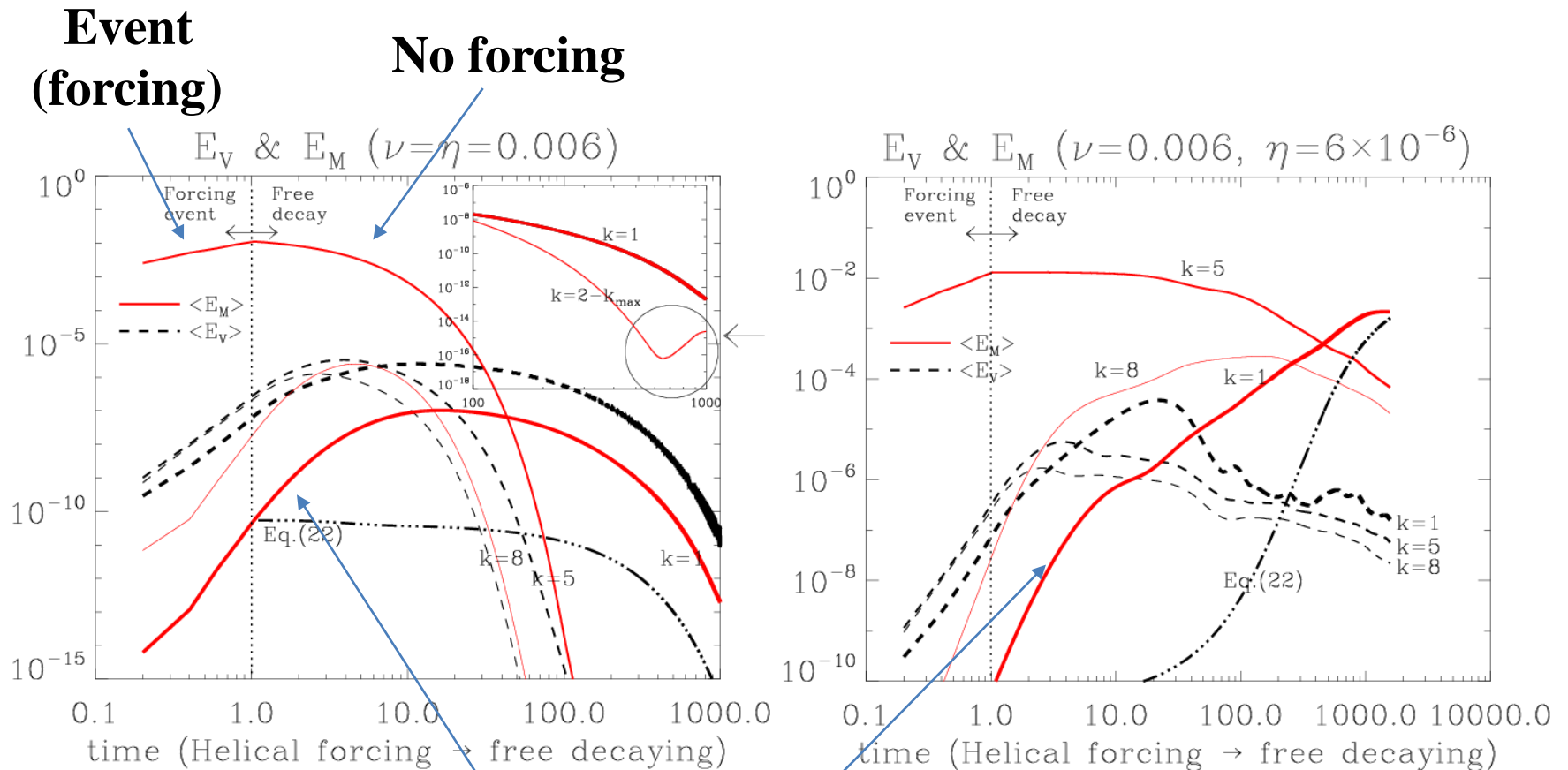


**Park 2017a, 2017b**

**Large Scale**

# ■ Energy Spectrum in a decaying MHD system

- **Helical** magnetic energy



**Large Scale**

- **How can we explain?**

**(1) Helical case**

-  $\alpha$  effect

$$\begin{aligned}\frac{\partial \bar{\mathbf{B}}}{\partial t} &= \nabla \times \langle \mathbf{u} \times \mathbf{b} \rangle + \eta \nabla^2 \bar{\mathbf{B}} \\ &\Rightarrow \nabla \times \alpha \bar{\mathbf{B}} + (\eta + \beta) \nabla^2 \bar{\mathbf{B}}\end{aligned}$$

$$\because \alpha \sim \int \langle \mathbf{j} \cdot \mathbf{b} \rangle - \langle \mathbf{u} \cdot \boldsymbol{\omega} \rangle d\tau \neq 0$$

**(2) NonHelical case?**

- **No  $\alpha$  effect  $\Rightarrow \bar{\mathbf{B}} \rightarrow \mathbf{0}$**

But,  $\bar{\mathbf{B}}(t) \neq 0$ ,



# ■ Analytic solution

## 1. Helical initial $E_{M0}$

$$\bar{A} \cdot \left( \frac{\partial \bar{B}}{\partial t} = \nabla \times \alpha \bar{B} + (\beta + \eta) \nabla^2 \bar{B} \right), \bar{B} \cdot \left( \frac{\partial \bar{B}}{\partial t} = \nabla \times \alpha \bar{B} + (\beta + \eta) \nabla^2 \bar{B} \right)$$

$$\Rightarrow \begin{cases} 2H_{ML} = (2E_{M0} + H_{M0})e^{2 \int (\alpha - \beta - \eta) d\tau} - (2E_{M0} - H_{M0})e^{-2 \int (\alpha + \beta + \eta) d\tau} \\ 4E_{ML} = (2E_{M0} + H_{M0})e^{2 \int (\alpha - \beta - \eta) d\tau} + (2E_{M0} - H_{M0})e^{-2 \int (\alpha + \beta + \eta) d\tau} \end{cases}$$

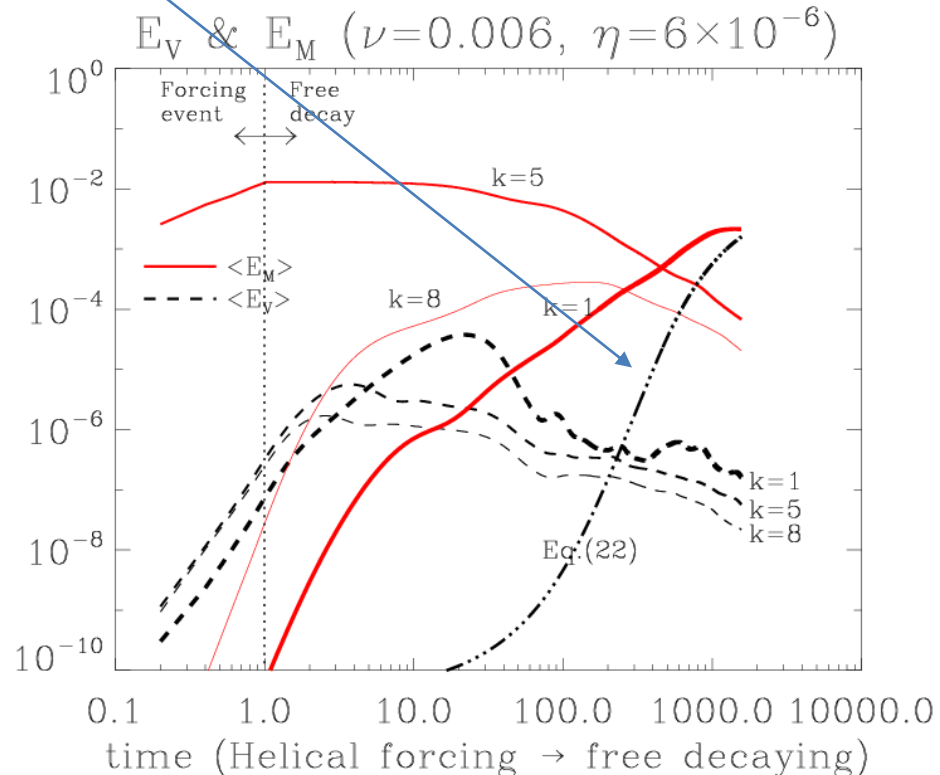
$$\alpha = \frac{1}{3} \int \langle \mathbf{j} \cdot \mathbf{b} \rangle d\tau$$

$$\sim \sum \langle \mathbf{j} \cdot \mathbf{b} \rangle(k) \Delta t$$

$$k = 4 - 6, 7$$

$$\Delta t = t_n - t_{n-1}$$

$$\beta = \frac{1}{3} \int \langle u^2 \rangle d\tau$$



# ■ Analytic solution

## 2. Nonhelical initial $E_{M0}$

- Statistical method for the **ideal** MHD system
  - a. 2D MHD (Fyfe & Montgomery 1976)

① **Conservation**

$E, \langle \mathbf{u} \cdot \mathbf{B} \rangle, \langle A^2 \rangle$  (stationary system)

② **Gibbs distribution function**

$$Z_0 \exp[-(\alpha E + \beta \langle \mathbf{u} \cdot \mathbf{B} \rangle + \frac{\gamma \langle A^2 \rangle}{2})]$$

③  $E_M = \left( \alpha - \frac{\beta^2}{4\alpha} + \frac{\gamma}{k^2} \right)^{-1}$

④ With  $\beta=0$  ( $P=0$ , No cross helicity) and  $\gamma < 0$

$$E_M = \frac{1}{\left( \alpha - \frac{|\gamma|}{k^2} \right)} \rightarrow E_{M, peak} \text{ at } k_{min} \leftarrow \text{Inverse transfer of } E_M.$$

b. 3D MHD (Frisch et al. 1975)

✳ Valid for a decaying MHD system? Assumption is not valid.

- **Scaling invariant method (Olesen 1997)**

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times \langle \mathbf{u} \times \mathbf{B} \rangle + \eta \nabla^2 \mathbf{B}$$

① Scale invariants :  $r \rightarrow lr, t \rightarrow l^{1-h}t, l^h u, l^h b, \dots$

$$\rightarrow E_{V,M}(k/l, l^{1-h}t) = E_{V,M}(k, t)$$

② Assuming  $E_{V,M}(k, t) \rightarrow k^{-1-2h}\psi_{V,M}(k, t)$

$$\rightarrow k \frac{\partial \psi_{V,M}}{\partial k} + (h - 1)t \frac{\partial \psi_{V,M}}{\partial t} = 0$$

③  $E_{V,M}(k, t) \rightarrow k^{-1-2h}\psi_{V,M}(k^{1-h}t)$  where  $k^{1-h}t = \text{const.}$   
If  $1 - h > 0 \rightarrow$  Inverse transfer of  $E_{V,M}(k, t)$ .

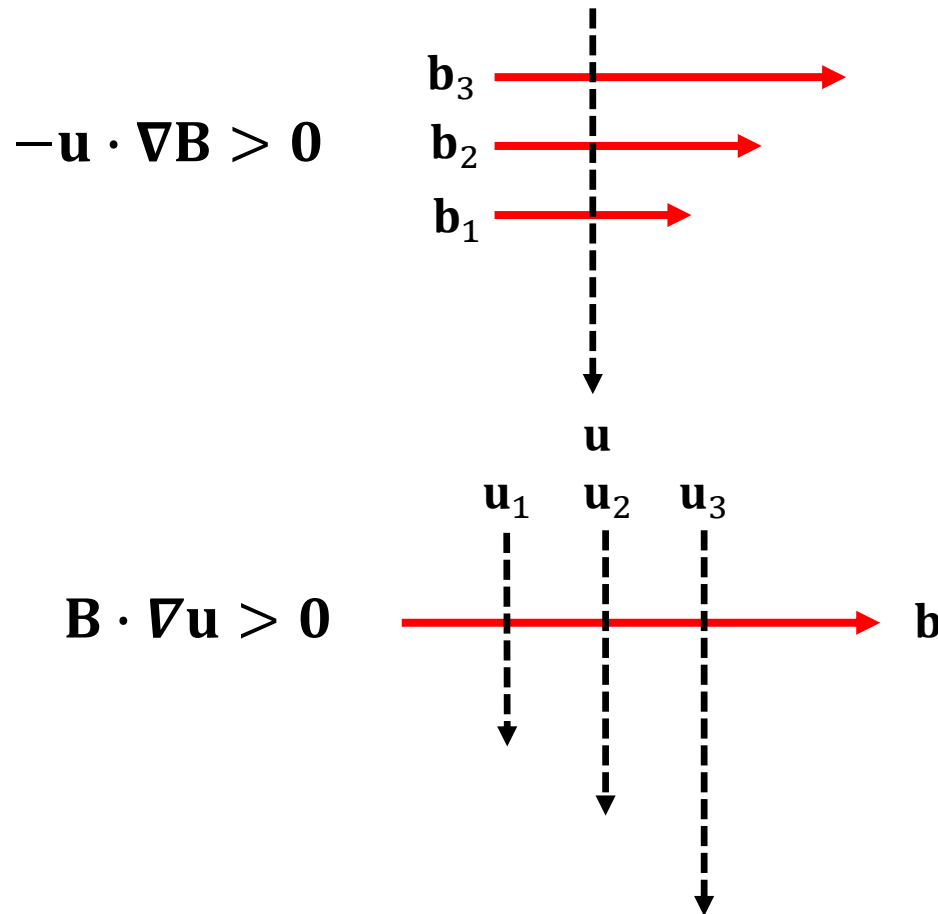
④  $\int E_{V,M}(k) dk = E_{V,M}$ ? Unreliable... (Ditlevsen et al. 2004)

- **Can we use another intuitive or theoretical model to explain the inverse transfer of nonhelical  $E_M$  in a decaying MHD system?**
- **Field structure model,  $\alpha$  theory, Quasi Normal approx.**

- **Field Structure model**

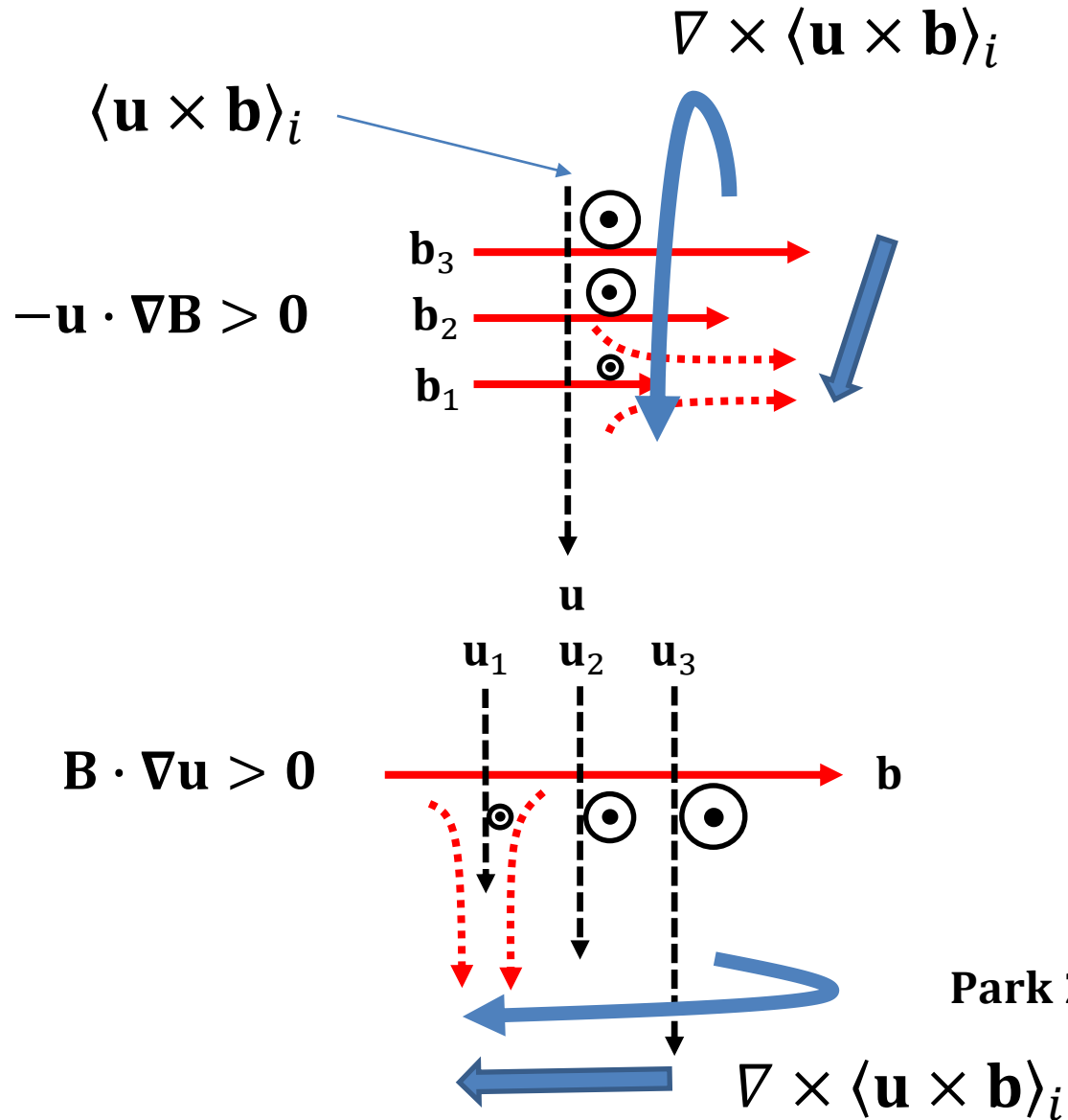
**Nonhelical B field**

$$\frac{\partial \mathbf{B}}{\partial t} = \mathbf{B} \cdot \nabla \mathbf{u} - \mathbf{u} \cdot \nabla \mathbf{B} + \eta \nabla^2 \mathbf{B}$$



• **Inhomogeneous EMF**

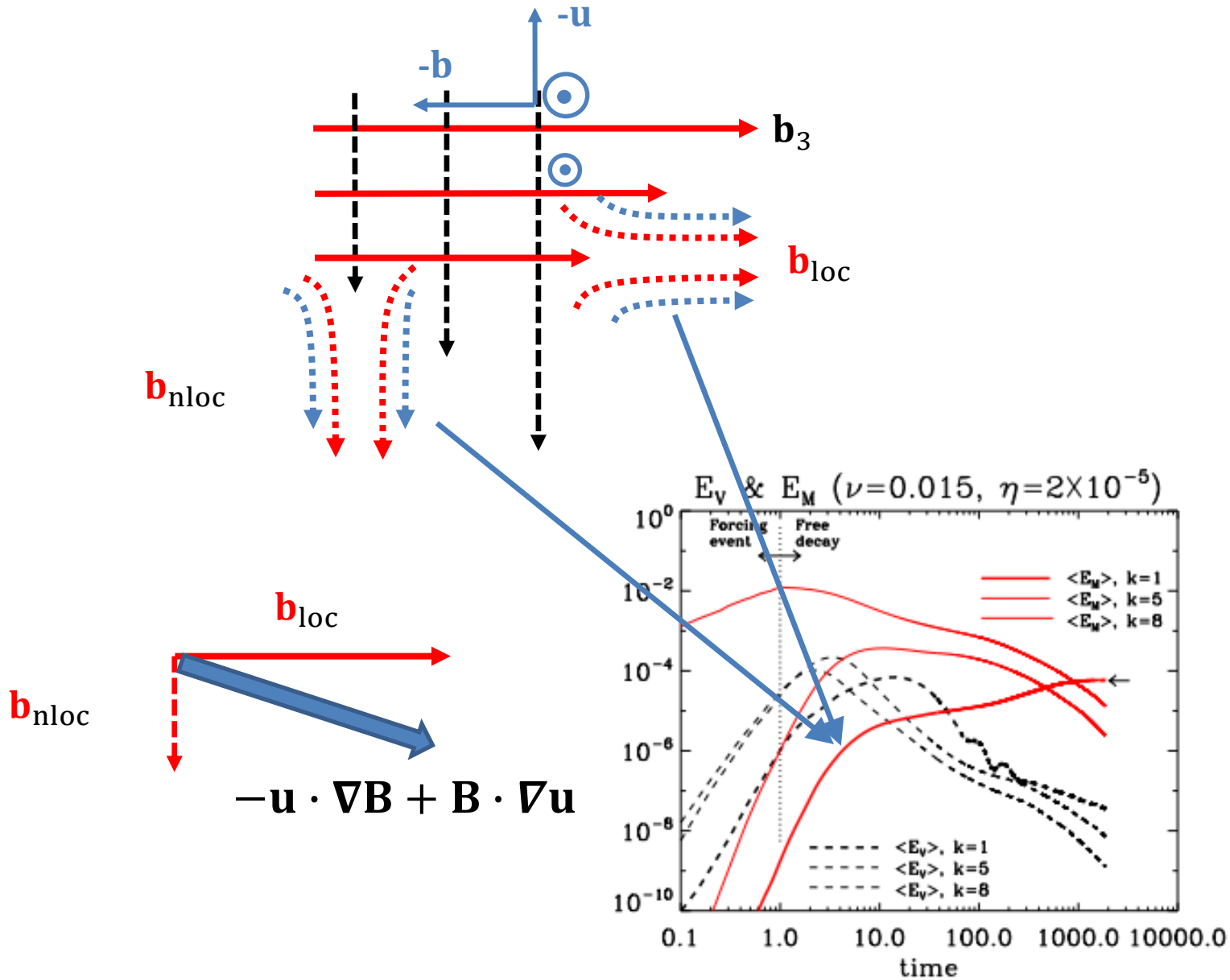
**Nonhelical B field**



Park 2017a, 2017b

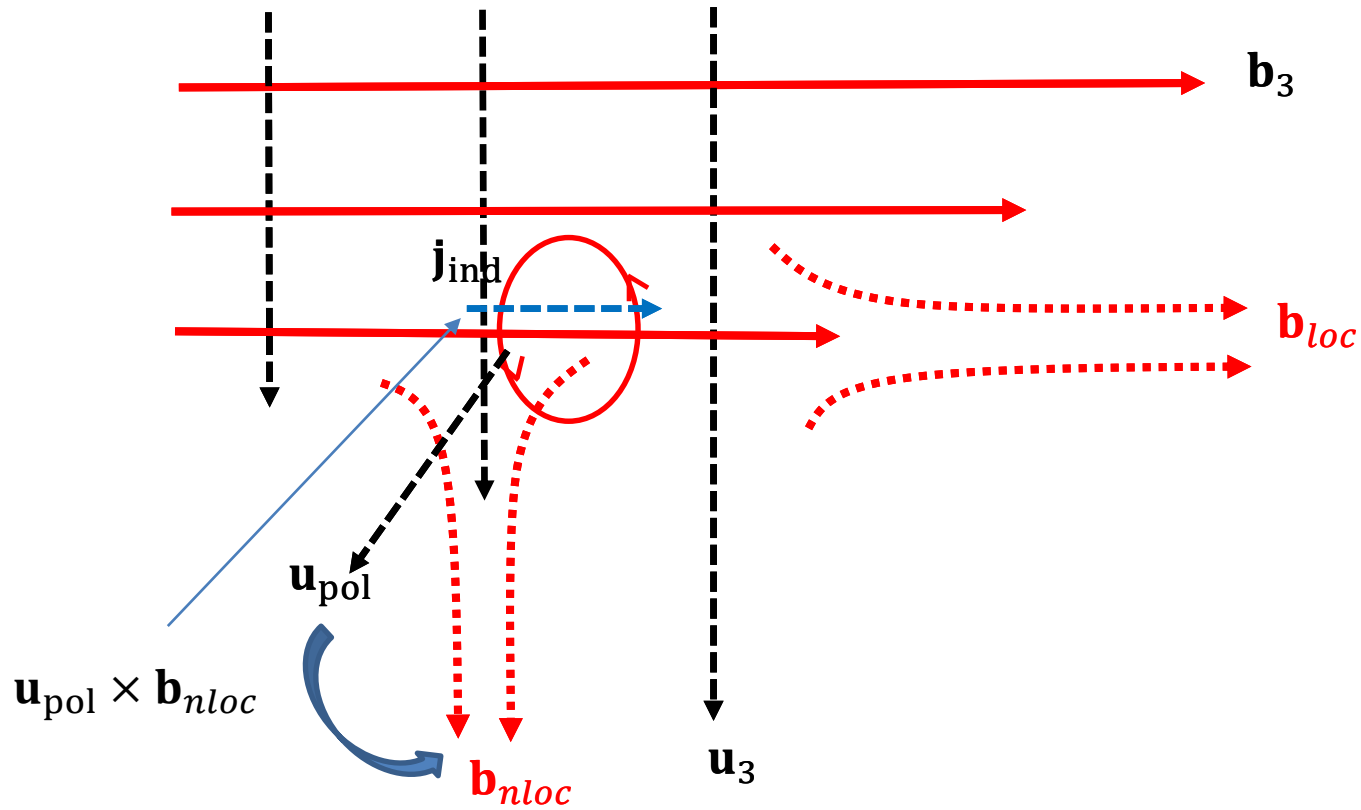
# • Dissipation effect

Nonhelical B field

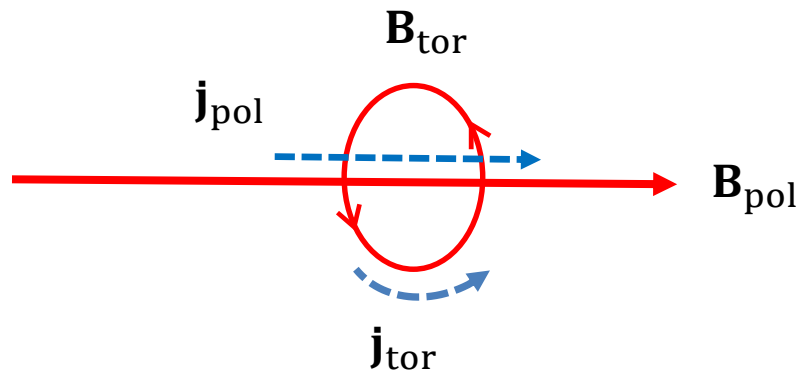


# • Helical Field effect

Helical B field



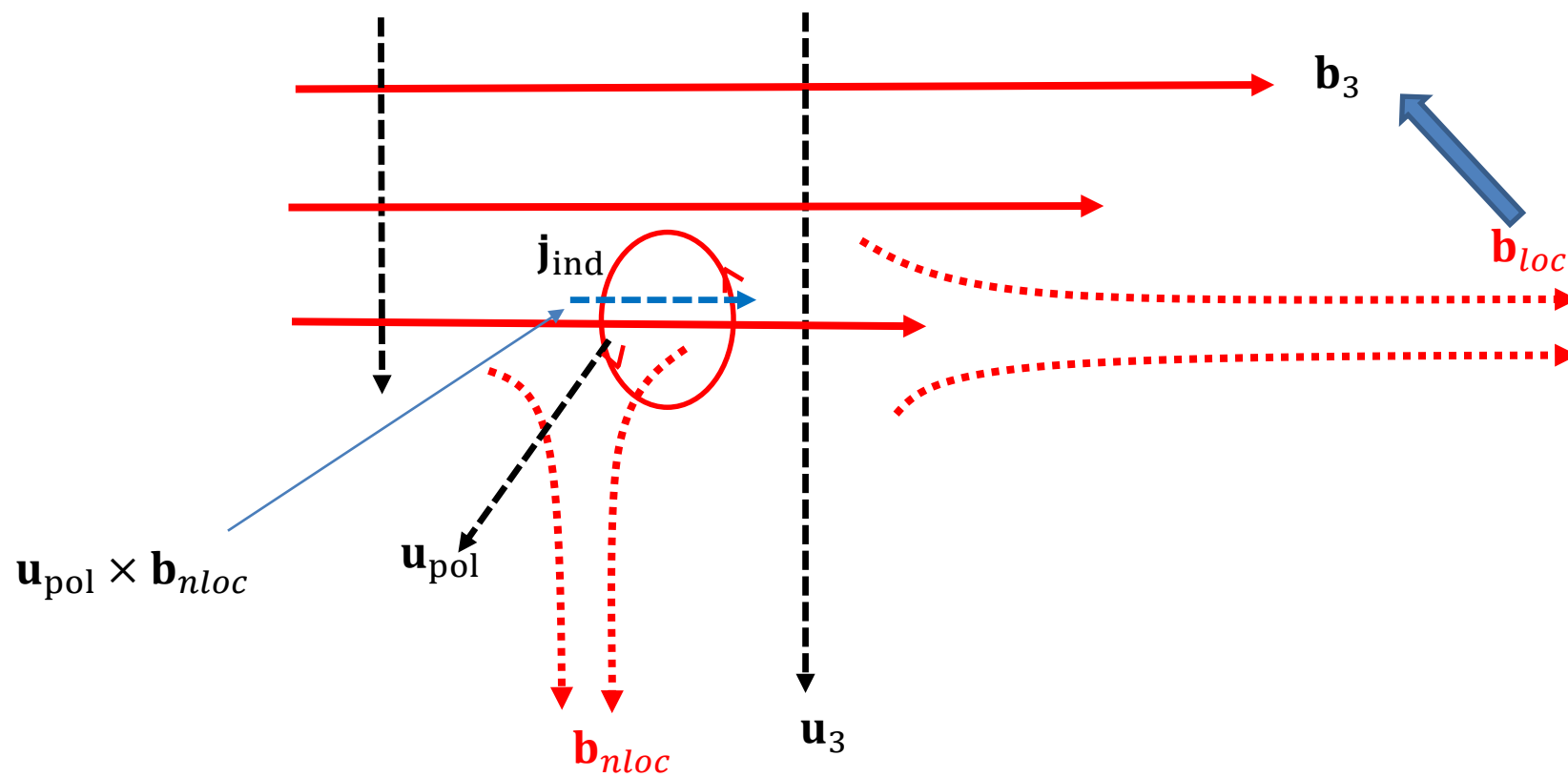
## Helical B field



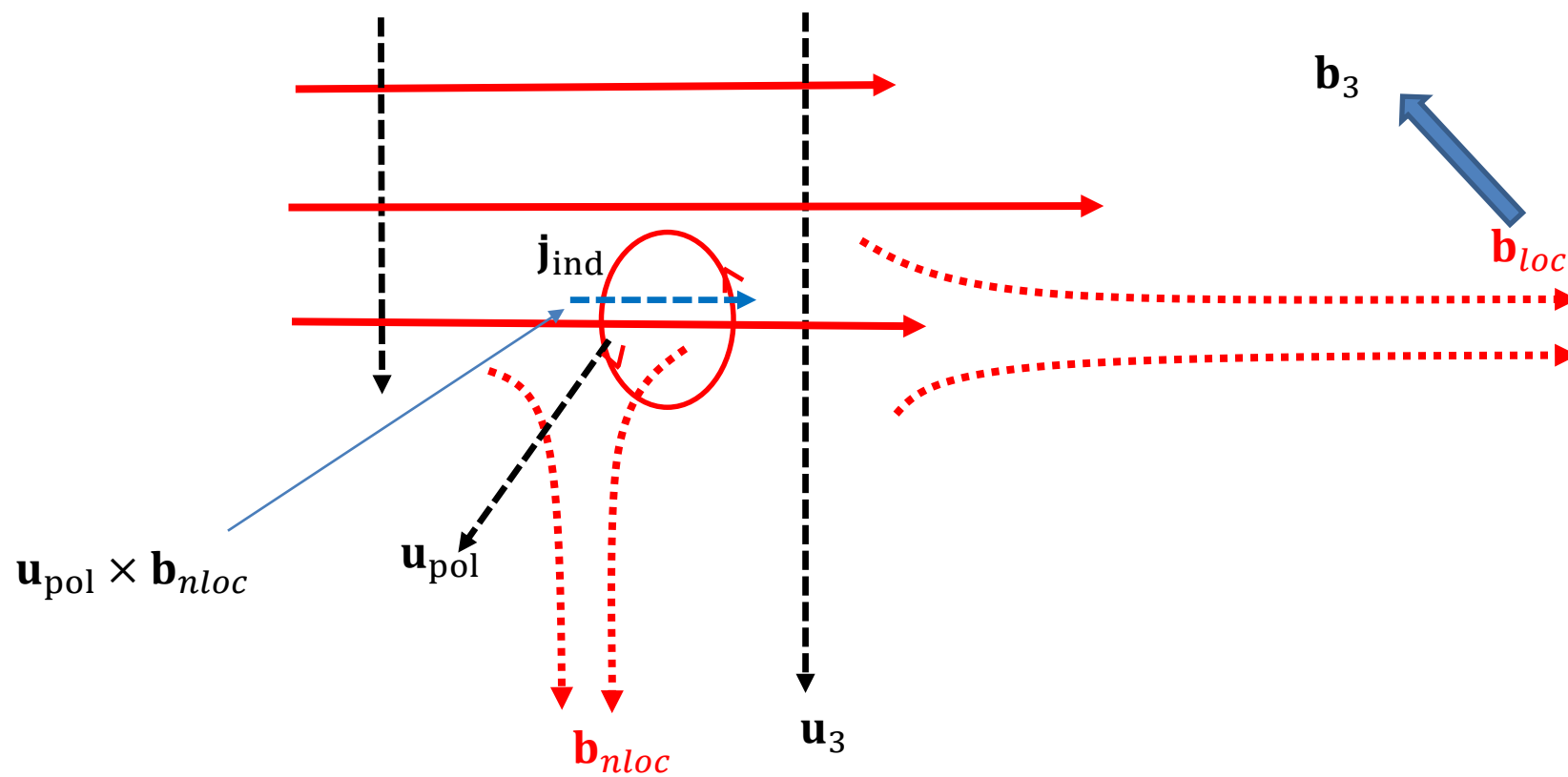
$$\langle \mathbf{u} \cdot \boldsymbol{\omega} \rangle < 0 \rightarrow \langle \mathbf{A} \cdot \mathbf{B} \rangle > 0 \Rightarrow \alpha^2 \text{dynamo}$$



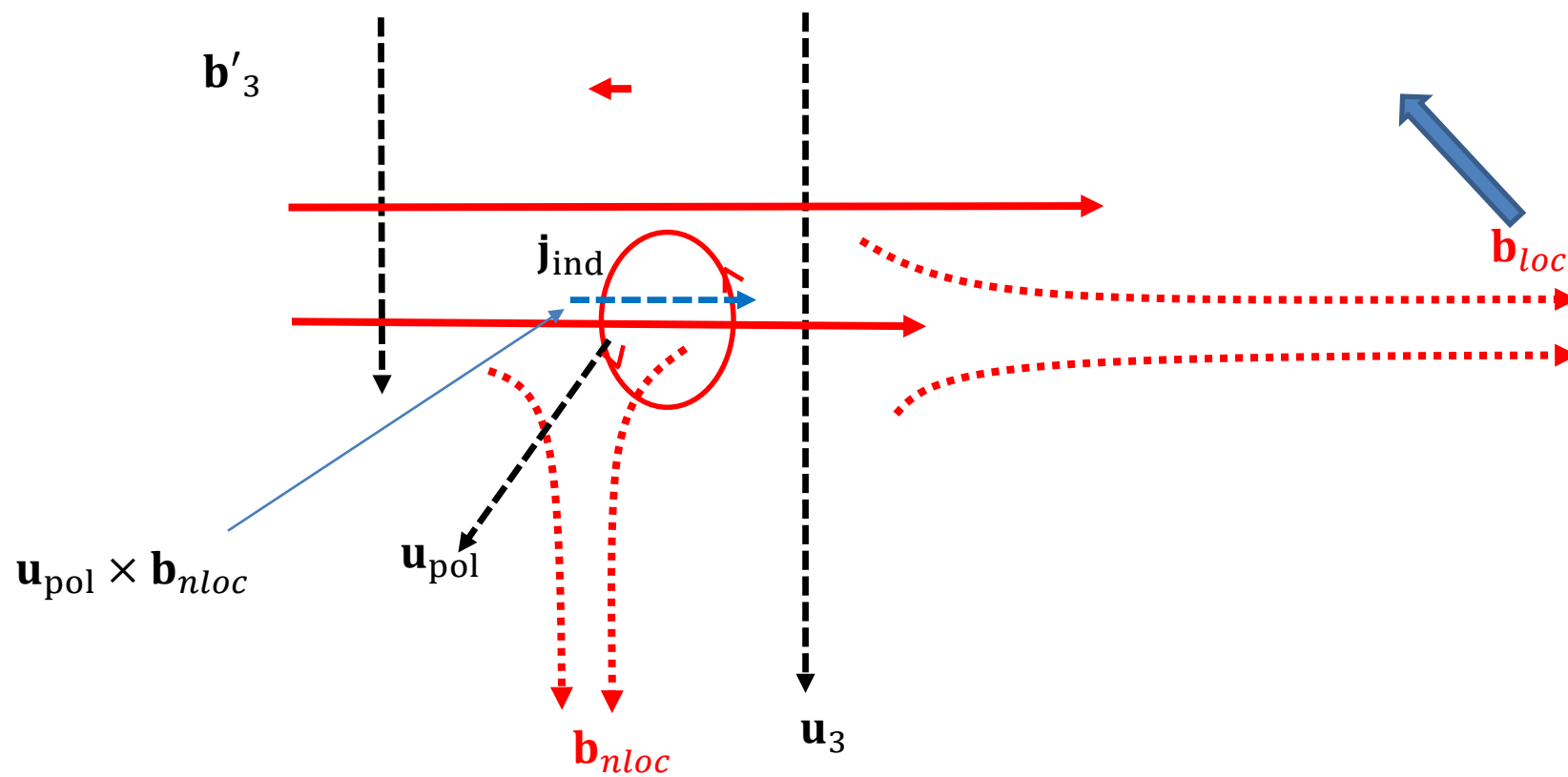
# Helical B field



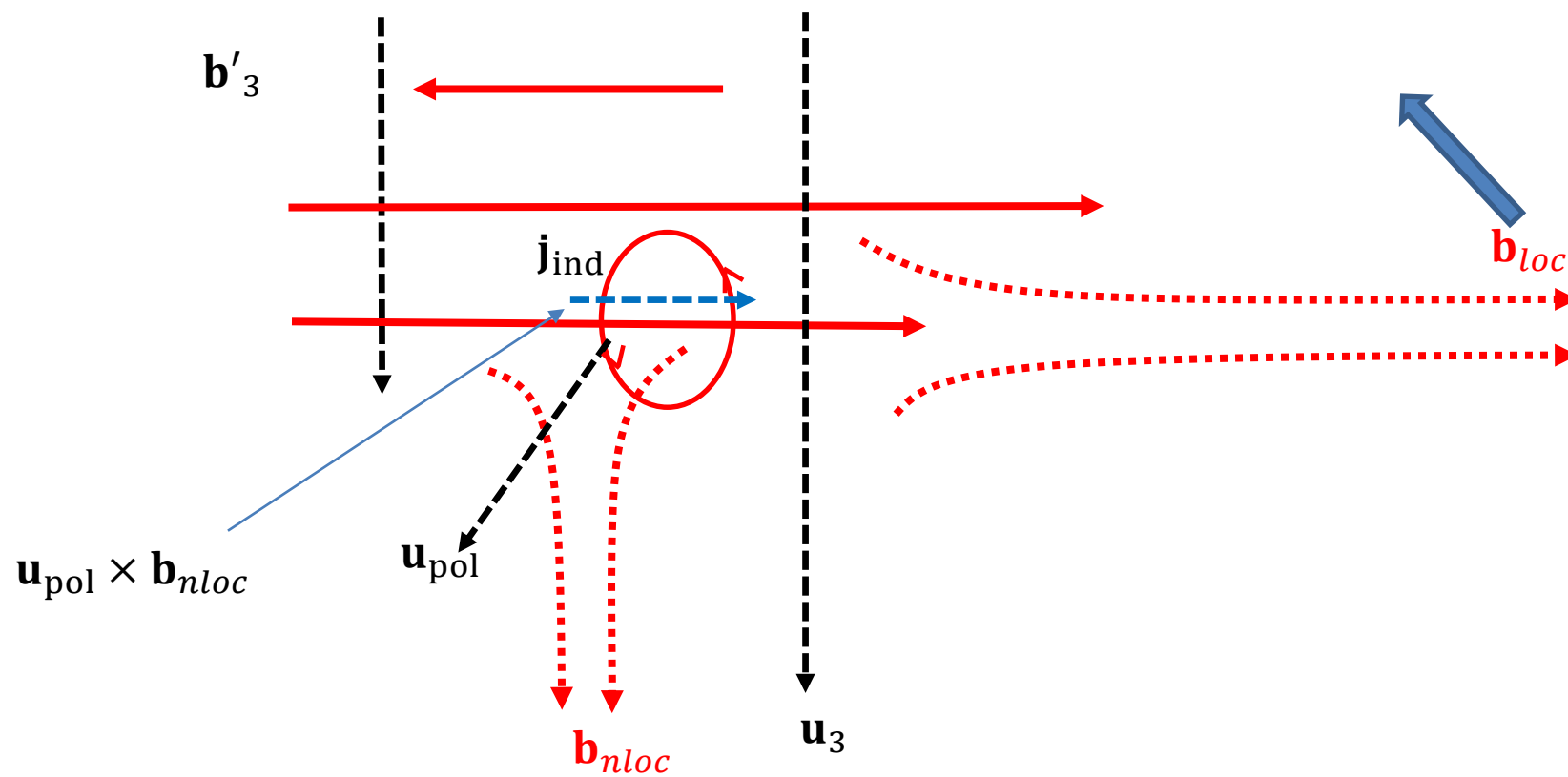
# Helical B field



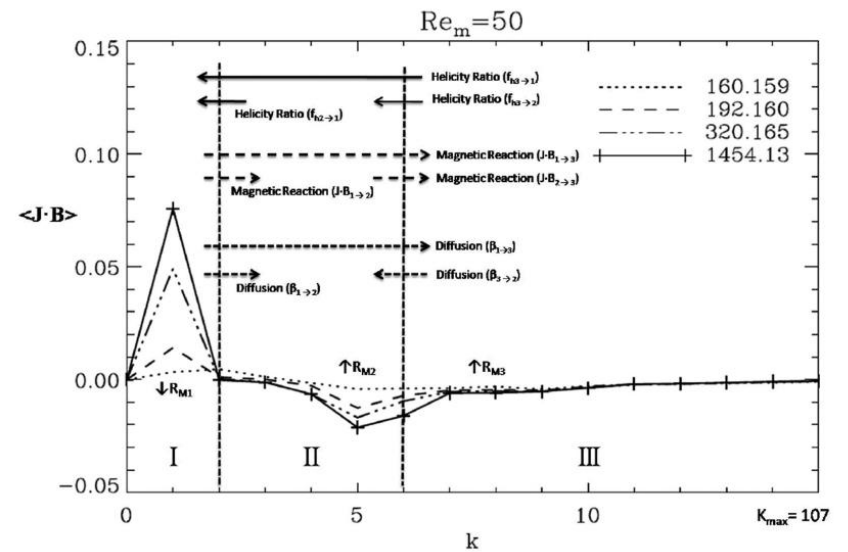
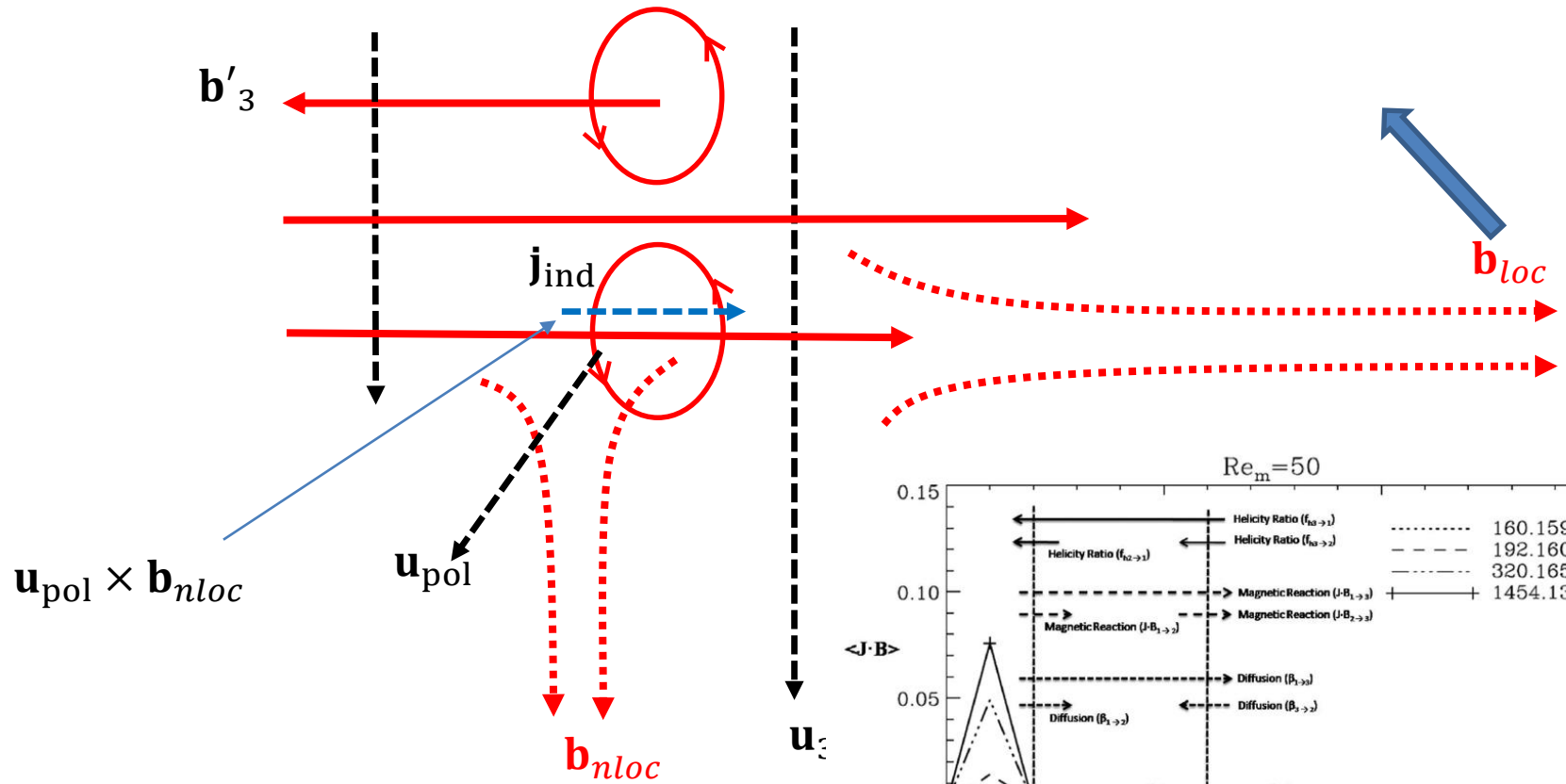
# Helical B field



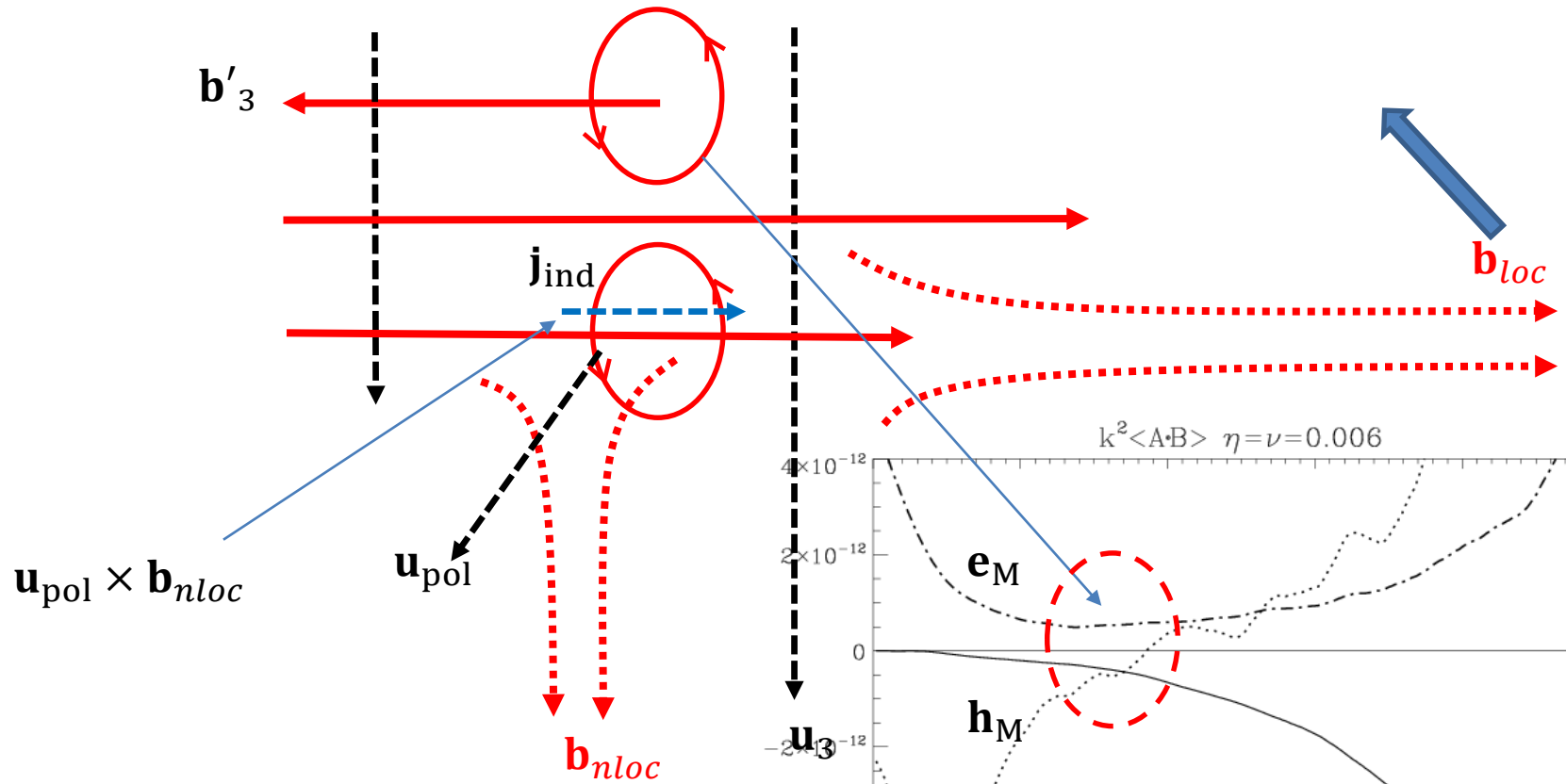
# Helical B field



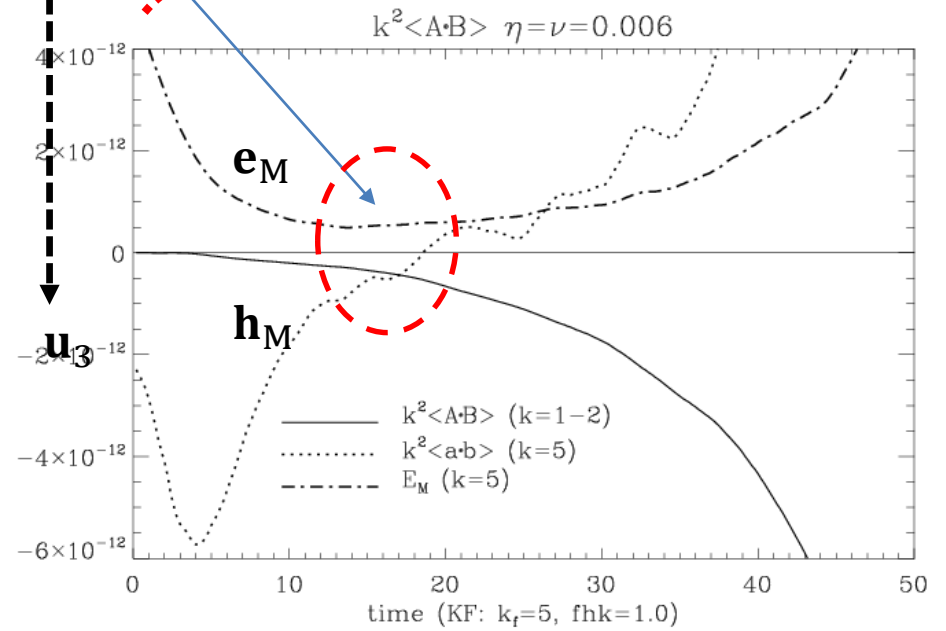
# Helical B field



# Helical B field



Park 2017a, 2017b



# ■ Resuming the analytic solution for the nonhelical

- EDQNM

(Eddy Damped Quasi Normalized Markovian approx.)

$$\frac{\partial \mathbf{v}}{\partial t} = -\mathbf{v} \cdot \nabla \mathbf{v} - \nabla P + \mathbf{B} \cdot \nabla \mathbf{B} + \nu \nabla^2 \mathbf{v}$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B}) + \lambda \nabla^2 \mathbf{B}$$

$$\langle \mathbf{v} \mathbf{v} \rangle = (4\pi k^2)^{-1} E_V$$

$$\langle \mathbf{B} \mathbf{B} \rangle = (4\pi k^2)^{-1} E_M$$

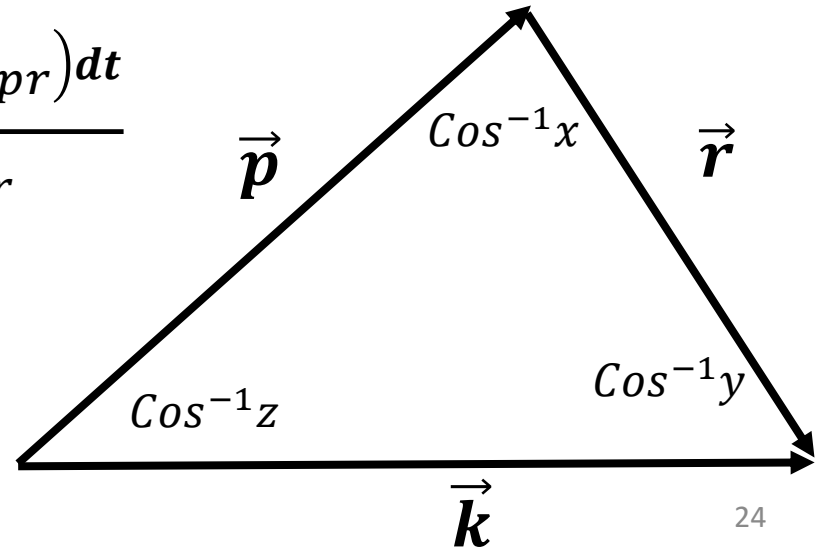
- **Analytic solution II (EDQNM)**

$$\begin{aligned} \frac{\partial E_M(k)}{\partial t} = & - \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{p^2}{r} z(1-x^2) E_M(r) E_M(k) \\ & - \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{r^2}{p} (y+xz) E_V(p) E_M(k) \\ & + \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{k^3}{pr} (1+xyz) E_V(p) E_M(r) \end{aligned}$$

$$\theta_{kpr}^{\nu\eta\eta}(t) = \frac{1 - e^{-\int (\nu p^2 + \eta k^2 + \eta r^2 + \mu_{kpr}) dt}}{\nu p^2 + \eta k^2 + \eta r^2 + \mu_{kpr}}$$

(triad relaxation time)

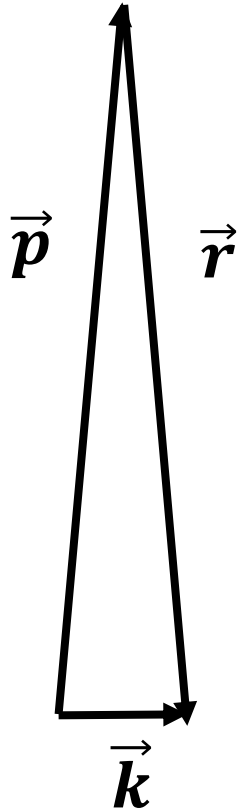
$$\mathbf{k} = \mathbf{p} + \mathbf{r}$$





• Analytic solution II (EDQNM)

$p \sim r \gg k$



$\eta \rightarrow 0, p \sim r \gg k = 1$

$$\theta_{kpr}^{\nu\eta\eta}(t) = \frac{1 - e^{-\int (\nu p^2 + \eta k^2 + \eta r^2 + \mu_{kpr}) dt}}{\nu p^2 + \eta k^2 + \eta r^2 + \mu_{kpr}} \rightarrow \frac{1}{\nu p^2}$$

$$\begin{aligned} \frac{\partial E_M(k, t)}{\partial t} = & - \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{p^2}{r} z(1 - x^2) \dot{E}_M(r) E_M(k) \\ & - \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{r^2}{p} (y + xz) \dot{E}_V(p) E_M(k) \\ & + \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{k^3}{pr} (1 + xyz) E_V(p) E_M(r) \end{aligned}$$

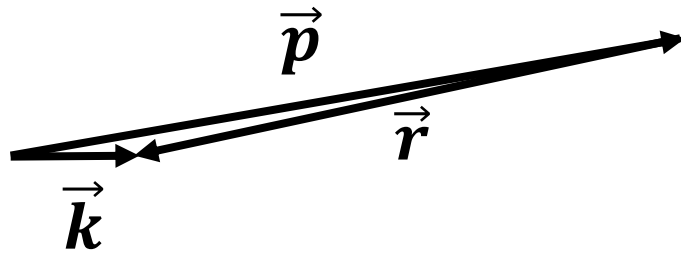
$$\frac{\partial E_M(k, t)}{\partial t} \sim \frac{1}{\nu p^4} E_V(p) E_M(p)$$

$$\rightarrow E_M(k, t) \sim \int \frac{1}{\nu p^4} E_V(p) E_M(r) dt \quad (k=1)$$

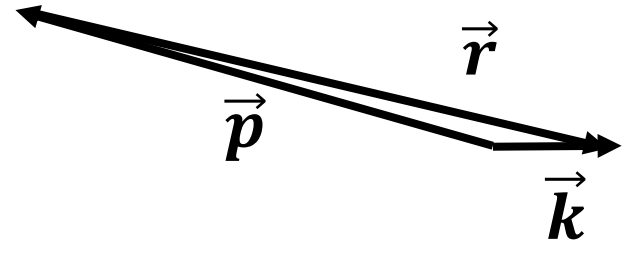
$x \sim 1, y \sim z \sim 0$

- Analytic solution II (EDQNM)

$$p \sim r \gg k$$



$$x \sim z \sim 1, y \sim -1$$

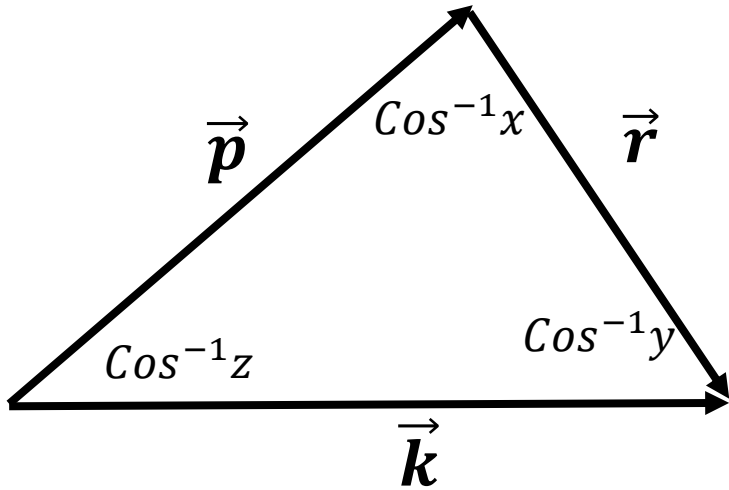


$$x \sim y \sim 1, z \sim -1$$

$$\begin{aligned} \frac{\partial E_M(k, t)}{\partial t} &= - \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{p^2}{r} z (1 - x^2) E_M(r) E_M(k) \\ &\quad - \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{r^2}{p} (y + xz) E_V(p) E_M(k) \\ &\quad + \int dp dr \theta_{kpr}^{\eta\eta\nu}(t) \frac{k^3}{pr} (1 + xyz) E_V(p) E_M(r) \\ &\sim 0 \end{aligned}$$

→ No inverse transfer from  $E_V(p)E_M(r)$

- **Analytic solution II (EDQNM)**



$$E_M(k, t) \sim \frac{Q}{P} - \frac{Q}{P^2} e^{-\int P dt}$$

$$P = f(E_M(p, t), E_V(r, t))$$

$$Q = g(E_M(p, t), E_V(r, t))$$

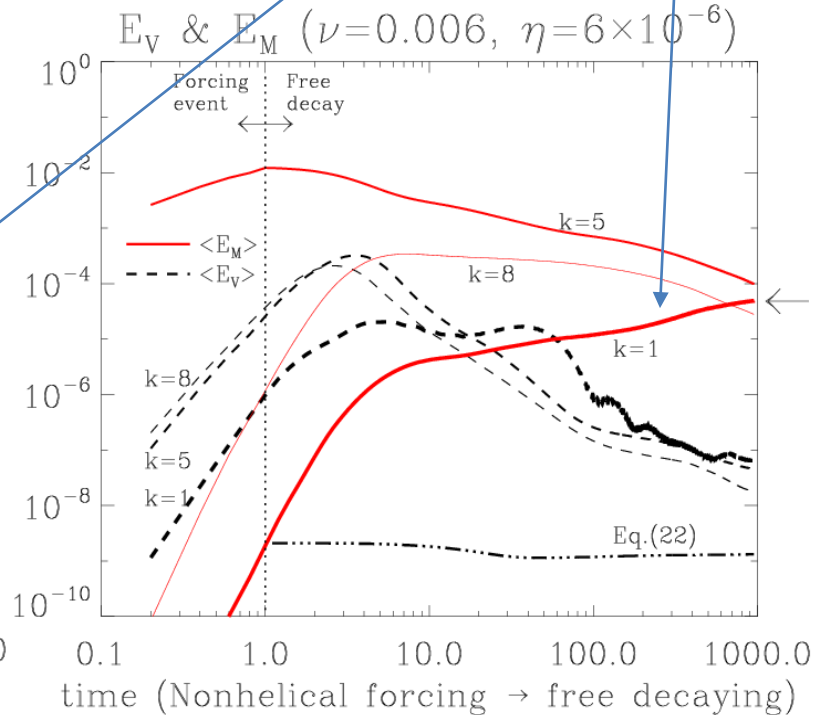
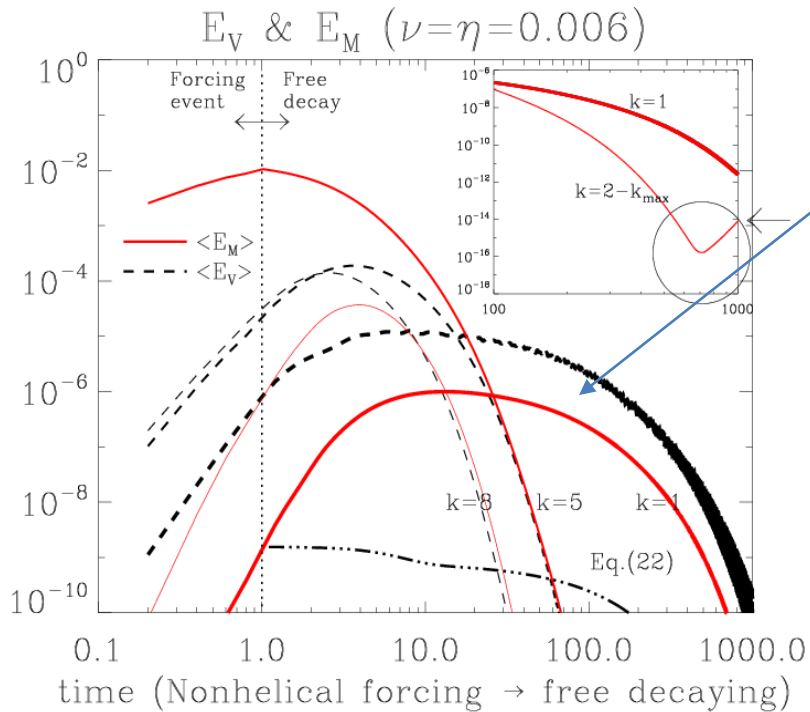
$$p \sim r \sim k, x \sim y \sim z \sim \frac{1}{2}$$

$$E_{M, tot}(k, t) \geq \int \frac{1}{vp^4} E_V(p) E_M(r) dt + \frac{Q}{P} - \frac{Q}{P^2} e^{-\int P dt}$$

# • Analytic solution II (EDQNM)

If  $\nu = \eta$  ( $p_{r_M} = 1$ , Son1999)

$$\theta_{kpr}^{\nu\eta\eta}(t) = \frac{1 - e^{-\int (\nu p^2 + \eta k^2 + \eta r^2 + \mu_{kpr}) dt}}{\nu p^2 + \eta k^2 + \eta r^2 + \mu_{kpr}} \rightarrow \frac{1}{\nu(2p^2 + r^2)} \left( < \frac{1}{\nu p^2} \right)$$



# ■ Summary

1. Inverse transfer of  $E_M$  in a decaying MHD system

2. Helical  $E_M$ :  $\alpha$  effect

3. Nonhelical  $E_M$ :

- Statistical method: conserved system variables

- Scaling invariant method:

$$E_M(k, t) = k^{-1-2h} \psi_{V,M}(k^{1-h}t), \quad k^{1-h}t = \text{const.}$$

4. Field structure from magnetic induction equation

$$-\mathbf{u} \cdot \nabla \mathbf{B}, \quad \mathbf{B} \cdot \nabla \mathbf{u}$$

5. EDQNM

$$p, r \rightarrow k \quad (p + r = k), \quad \theta_{kpr}^{v\eta\eta}(t), \quad x, y, z \Rightarrow E_M(k, t)$$