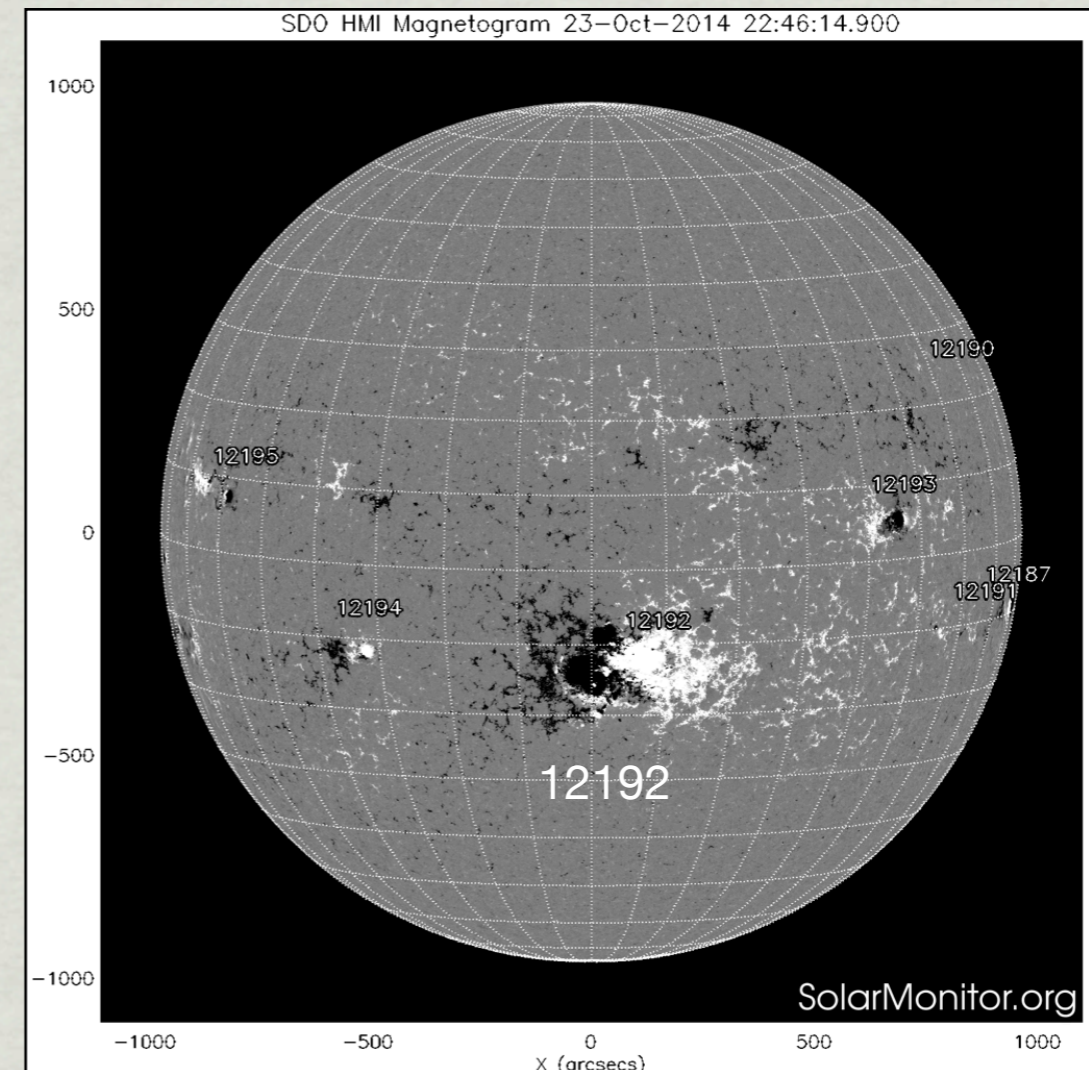
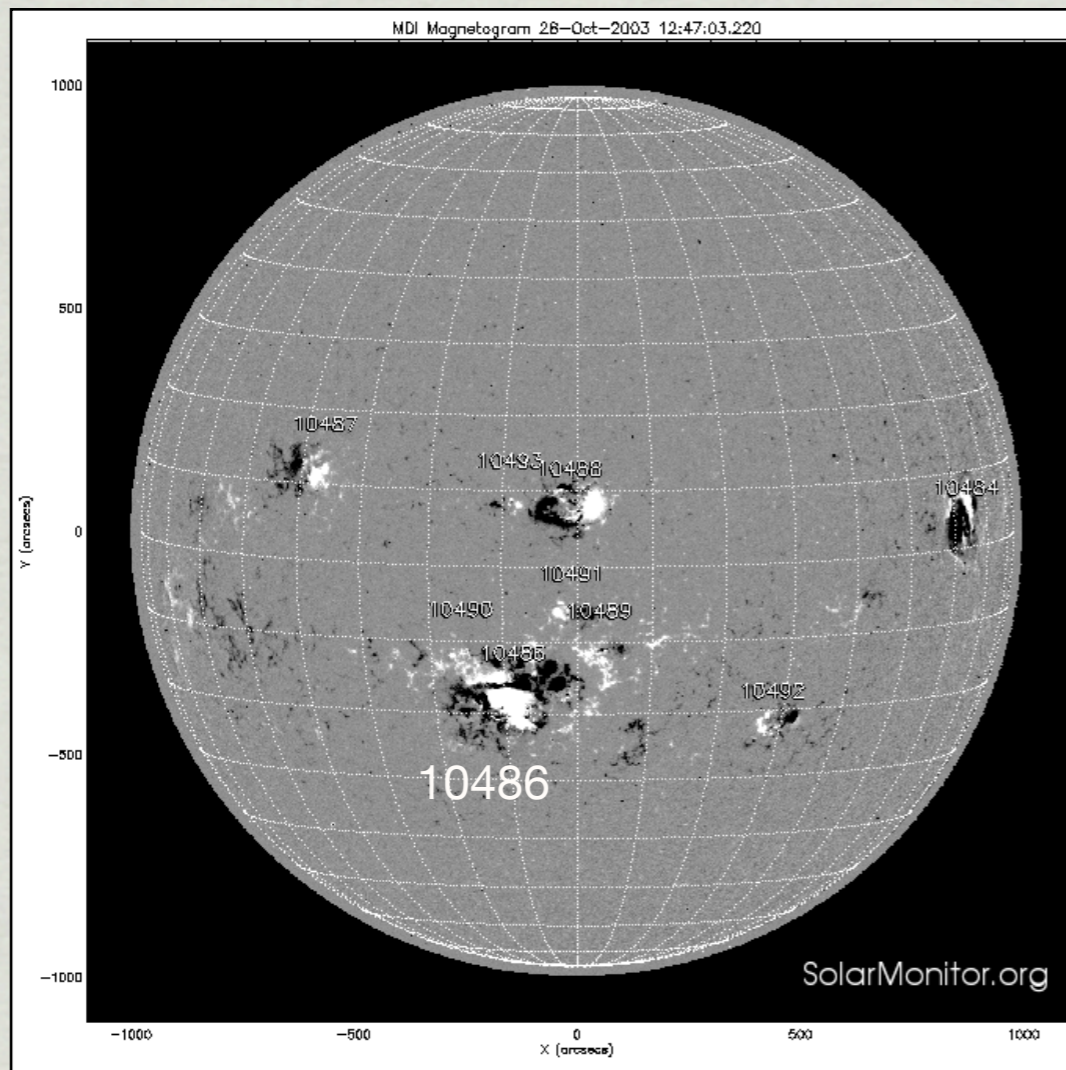


# **Subsurface Helicity of Active Regions 12192 & 10486**

**R. Komm, S. Tripathy, R. Howe, F. Hill**

# Introduction



Two active regions separated by 11 years.  
AR 10486 (left): large and complex, many flares;  
AR 12192 (right): larger but simpler, not as many flares.

# Introduction

- \* How different / similar are their subsurface flows?  
Especially their kinetic helicity?
- \* Is it obvious from the subsurface flows which region produced more eruptive events?
- \* Subsurface flows from ring-diagram analysis (0-16 Mm)
  - \* SOHO/MDI (AR 10486)
  - \* GONG (AR 10486, 12192)
  - \* SDO/HMI (AR 12192).

# Kinetic helicity (density)

Kinetic helicity density is the scalar product of the vorticity and velocity vector.

$$h_k = \boldsymbol{\omega} \cdot \boldsymbol{v}$$

Helicity: volume integral of helicity density

$$\boldsymbol{\omega} = \nabla \times \boldsymbol{v}$$

$\boldsymbol{\omega}$  : vorticity;  
 $\boldsymbol{v}$  : velocity vector

Kinetic helicity density (vertical contribution)

Div Curl of horizontal velocities is a proxy of the vertical contribution.

$$\langle h_z \rangle = \langle \boldsymbol{\omega}_z \cdot \boldsymbol{v}_z \rangle = \left\langle \left( \frac{\partial v_y}{\partial x} - \frac{\partial v_x}{\partial y} \right) \cdot v_z \right\rangle$$

# Current helicity (density)

Current helicity density (vertical component)

$$\langle h_c \rangle = \langle J_z \cdot B_z \rangle = \left\langle \left( \frac{\partial B_y}{\partial x} - \frac{\partial B_x}{\partial y} \right) \cdot B_z \right\rangle$$

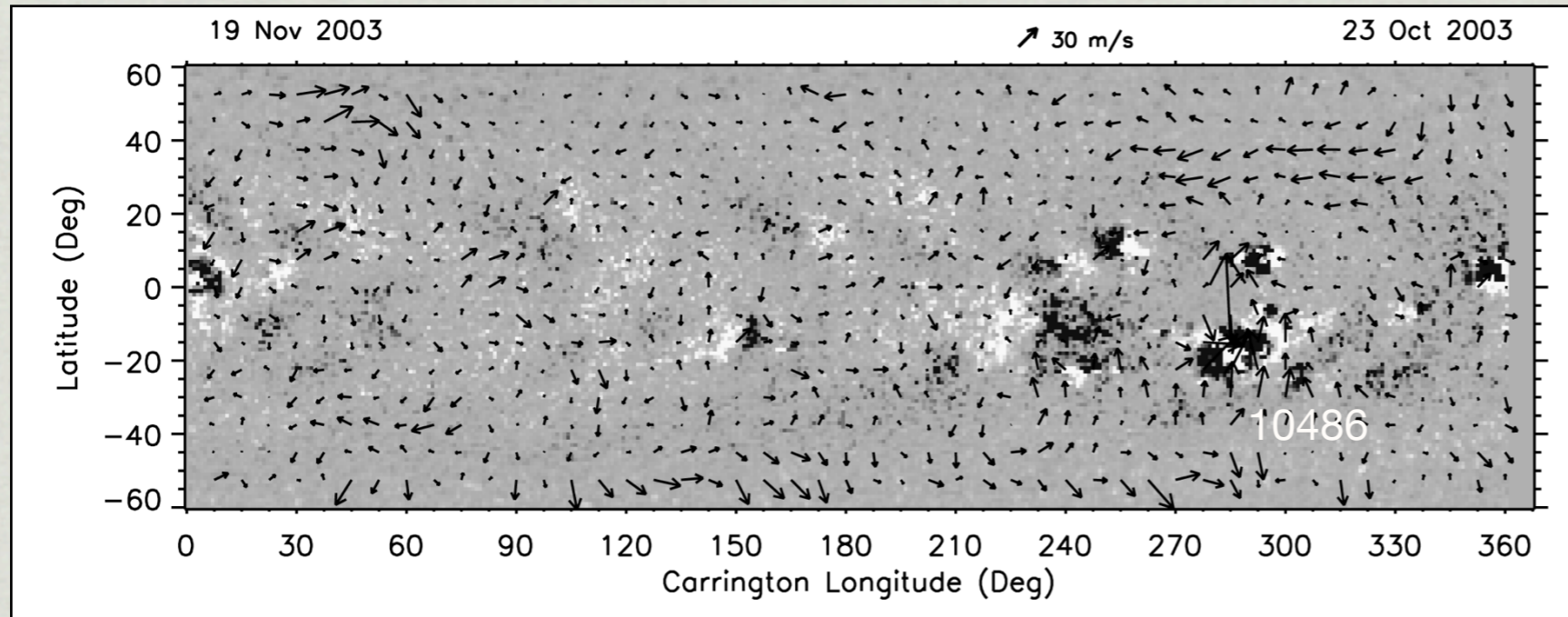
$J_z$  : vertical current density

$(B_x, B_y, B_z)$  : magnetic field components

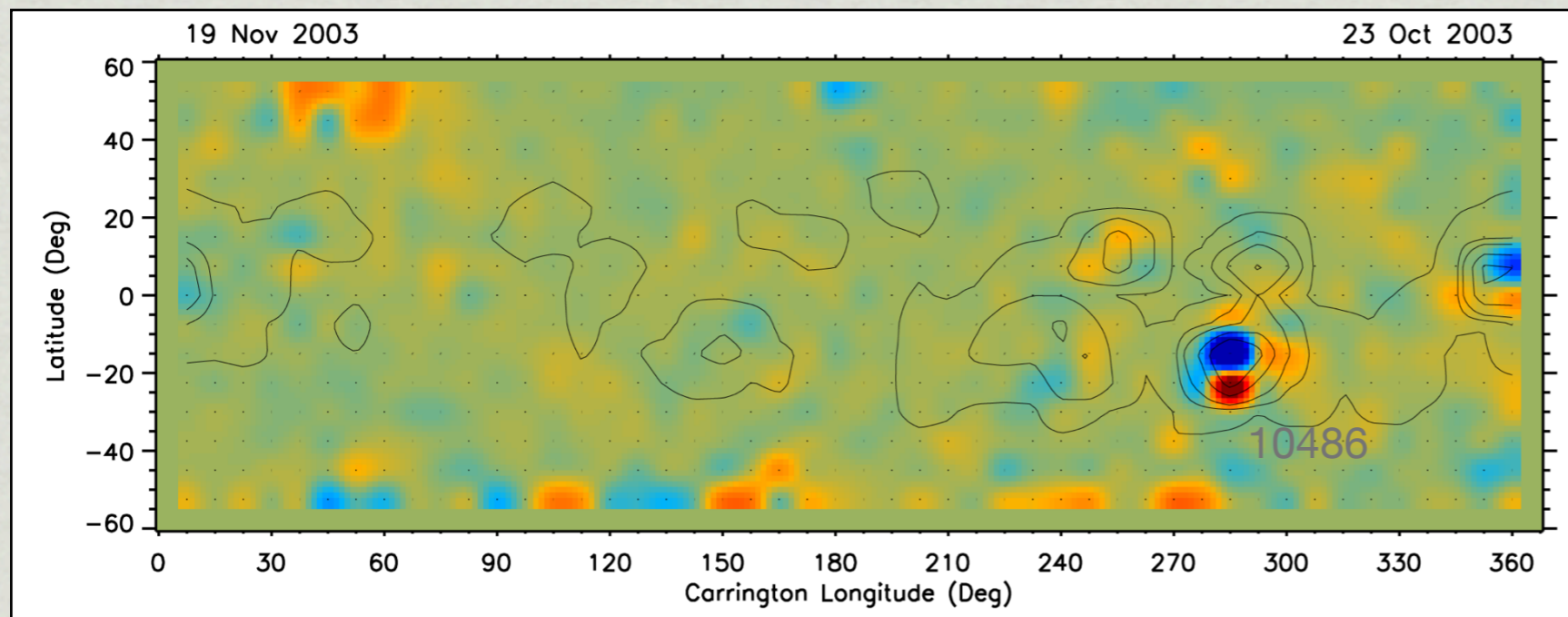
$$\langle \alpha_z \rangle = \langle J_z / B_z \rangle$$

Mean twist parameter

# Flow and helicity of AR 10486



Complex flows near active regions in synoptic maps.



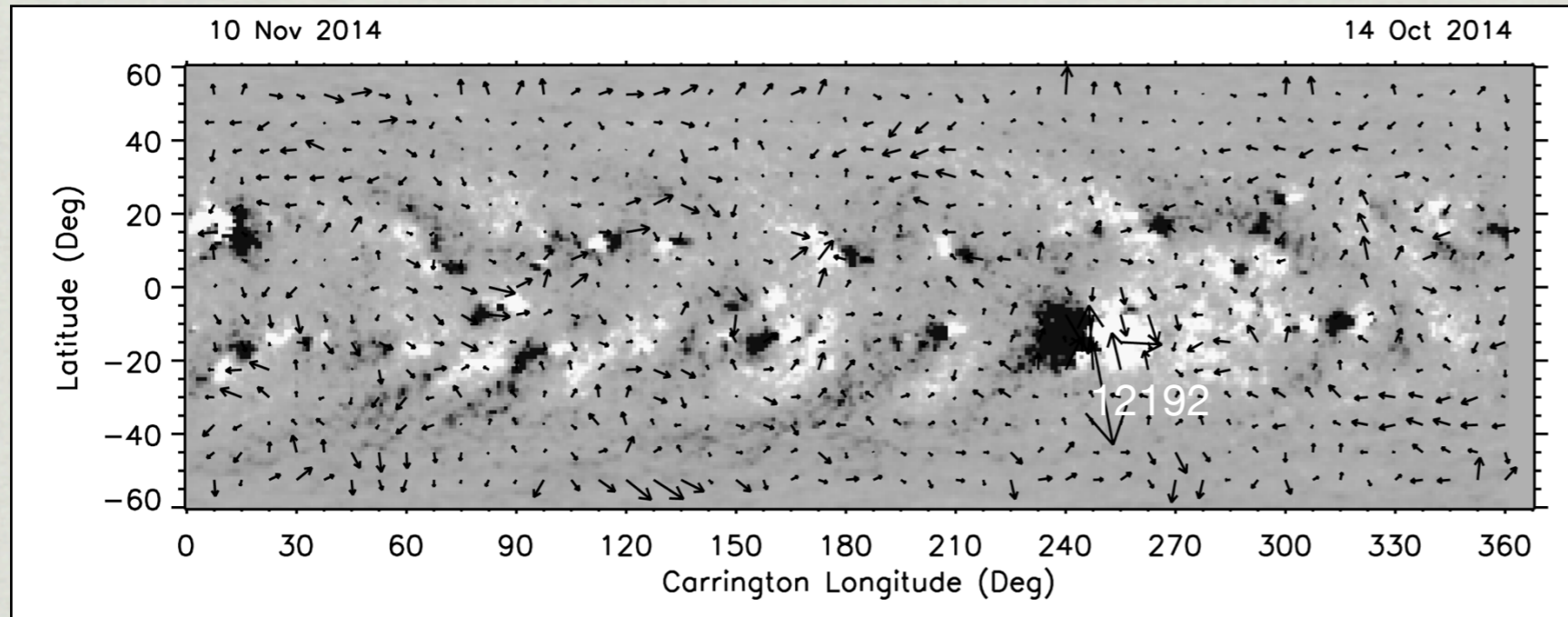
Flare-productive regions show “dipolar” pattern in kinetic helicity density.

CR 2009

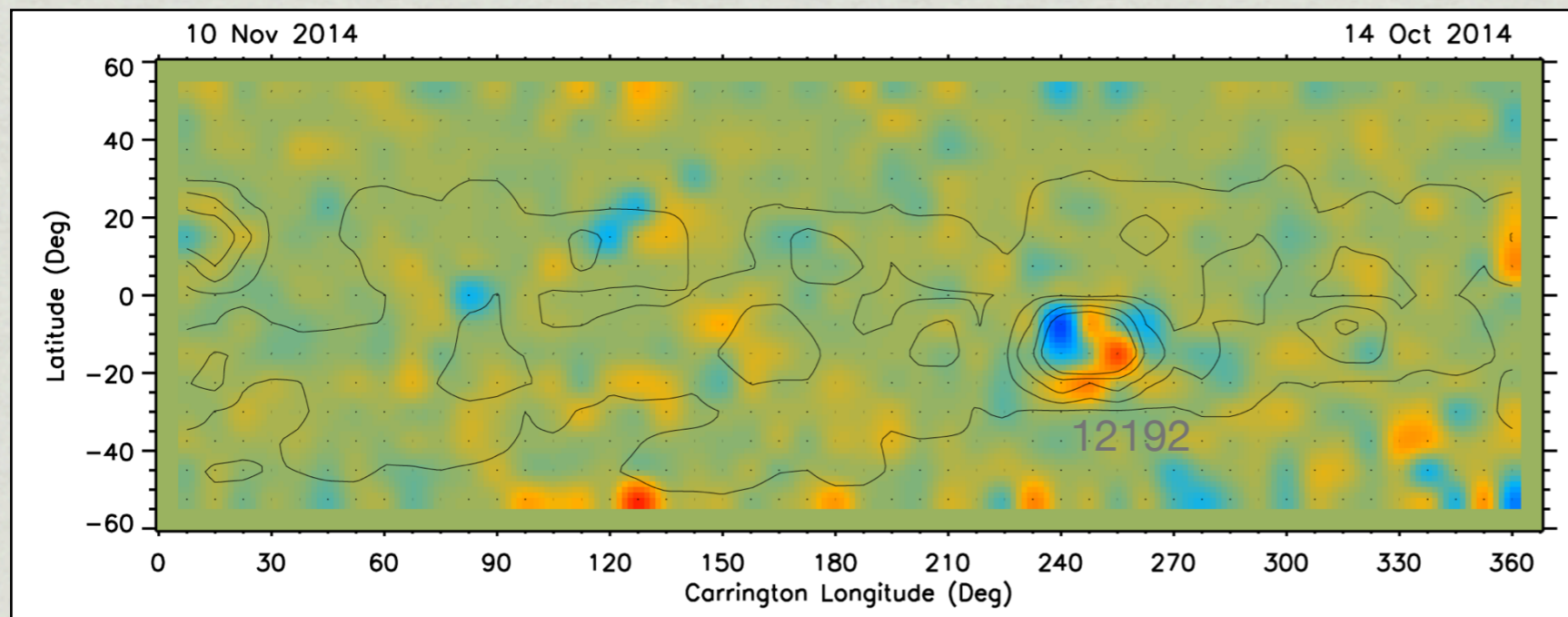
7.1 Mm

GONG

# Flow and helicity of AR 12192



Residual flows:  
average rotation  
and meridional  
flow has been  
subtracted.



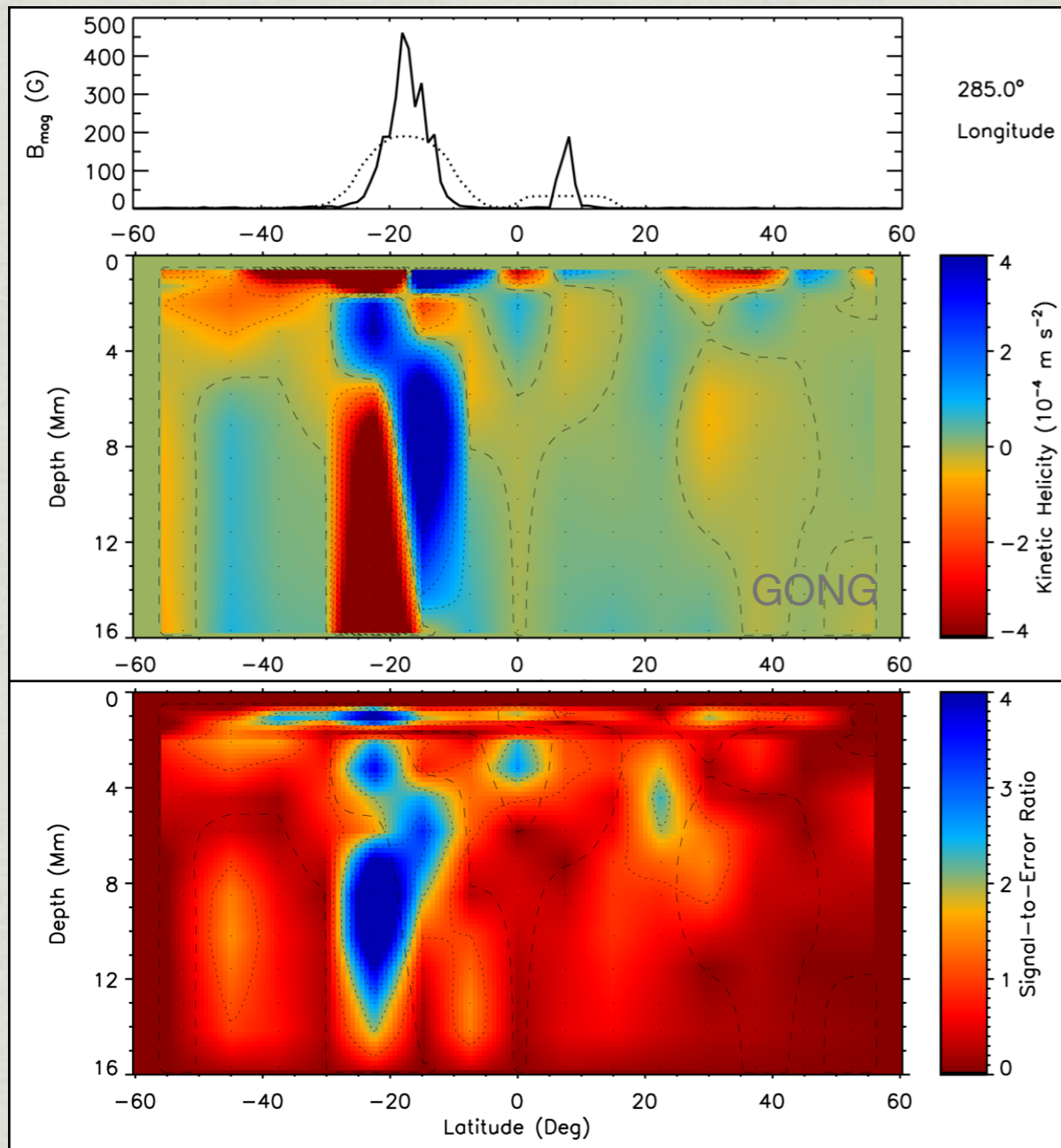
Magnetic polarity  
separates kinetic  
helicity pattern.

CR 2156

7.1 Mm

GONG

# Helicity slice of 10486



Top: magnetic flux  
Middle: kinetic helicity density  
Bottom: signal-to-noise ratio.

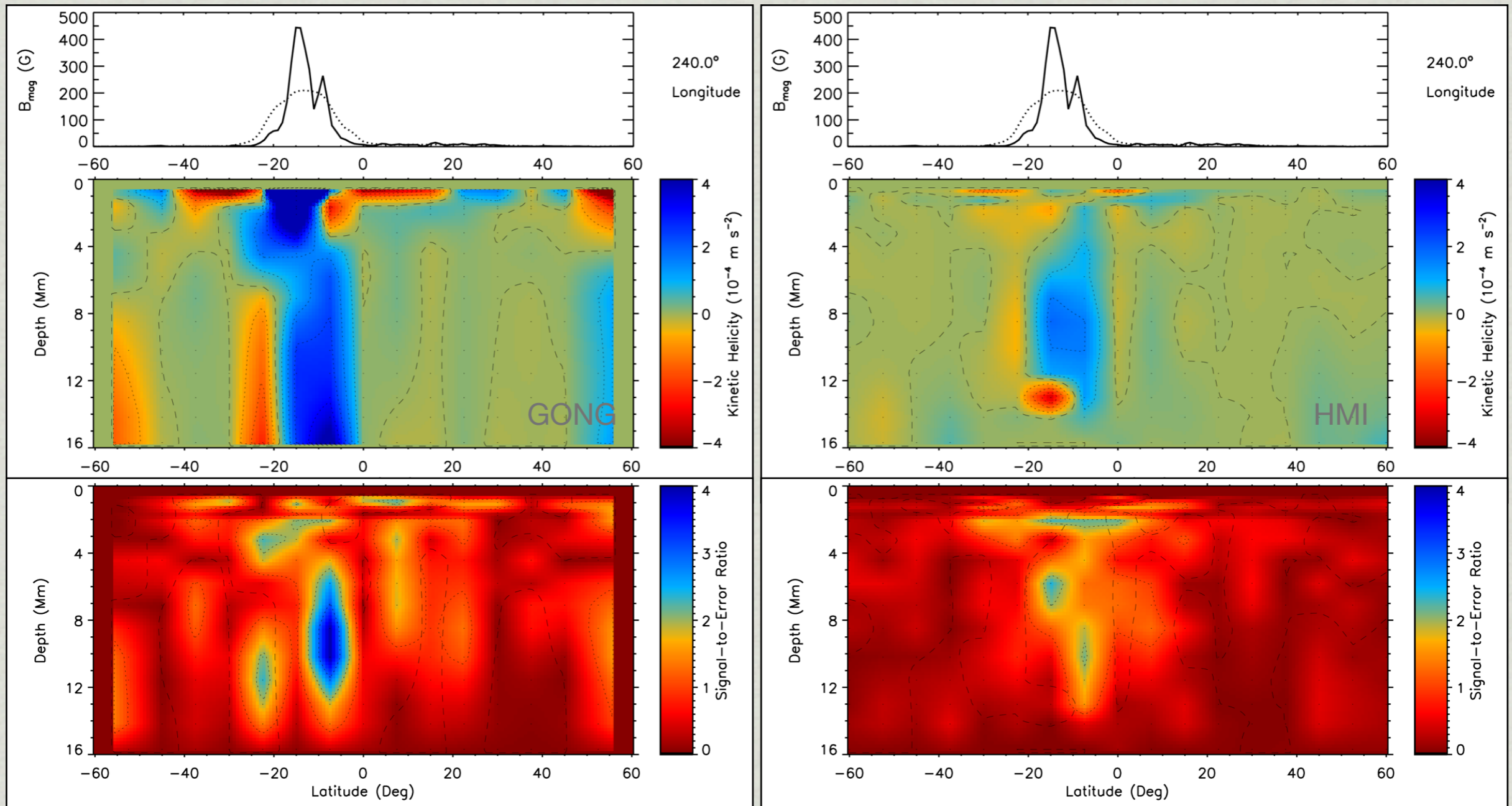
Strong “dipolar” pattern in kinetic helicity density with a sign change near 2 and 6 Mm.

This is a characteristic of flare-productive active regions.

Interpreted as stacked vortex rings.

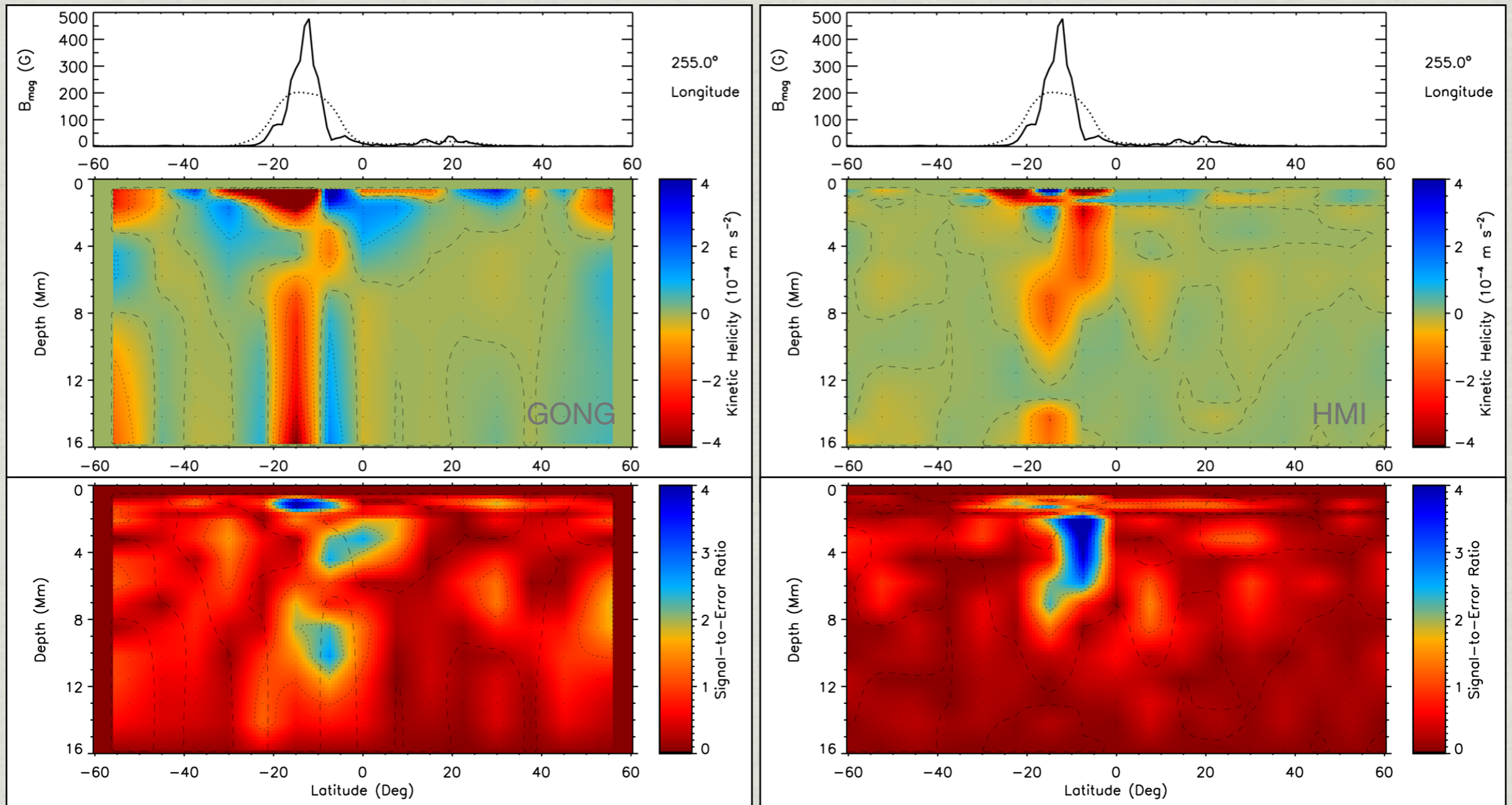


# Helicity slice of 12192 (N)



Dipolar pattern is rather weak. S/N ratio is smaller.  
GONG and HMI are reasonably similar for large S/N.

# Helicity slice of 12192 (P)

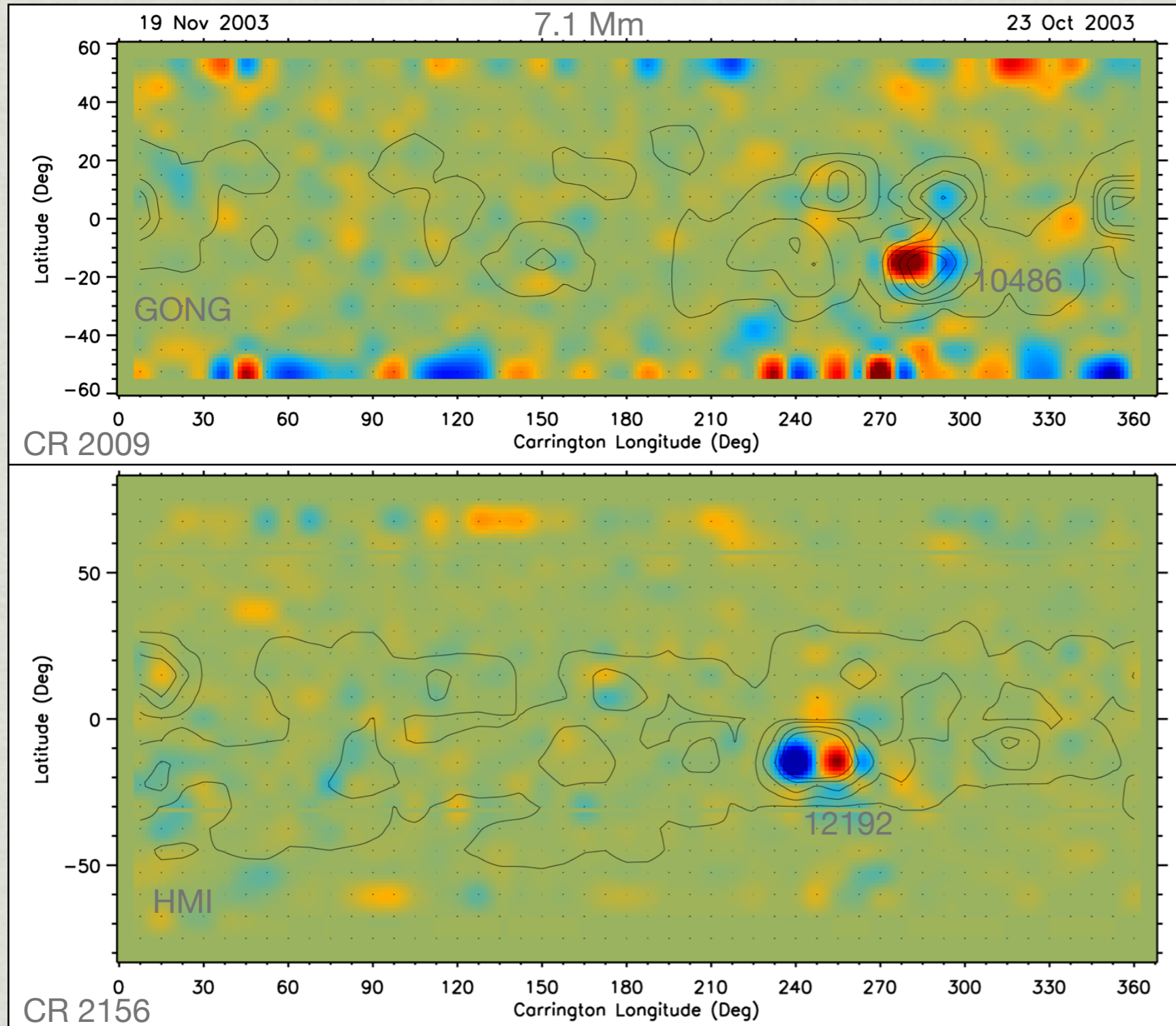


Dipolar pattern is rather weak. S/N ratio is smaller.  
GONG and HMI are reasonably similar for large S/N.

# Summary (1)

- \* The subsurface flows of AR 10486 show large values of kinetic helicity density and a strong dipolar pattern.
- \* The kinetic helicity values of AR 12192 are much smaller and the dipolar pattern is weaker.
- \* Conclusion: AR 10486 had greater potential of producing eruptive events.

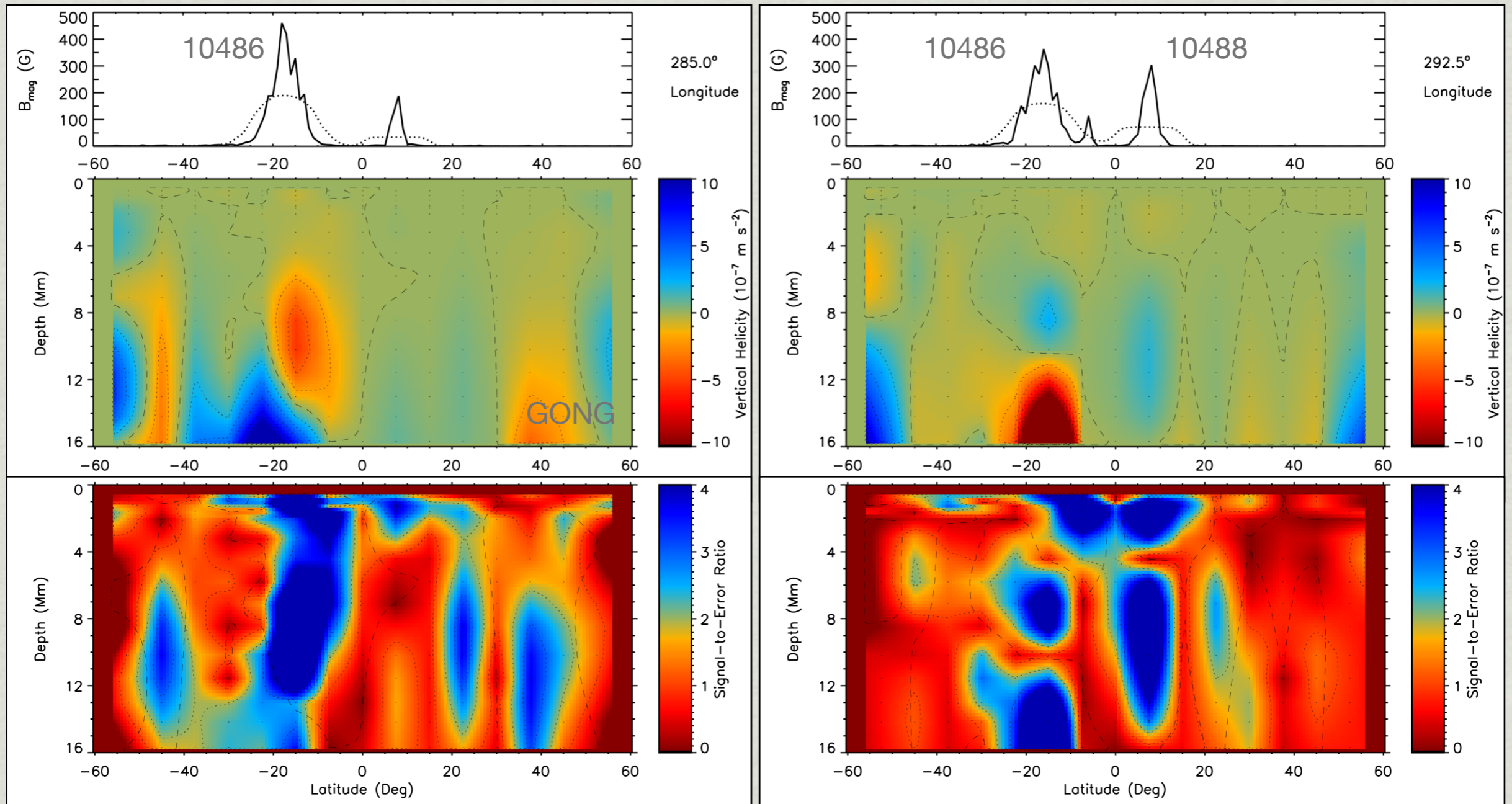
# Vertical helicity of 10486 & 12192



What about the vertical component of kinetic helicity density?

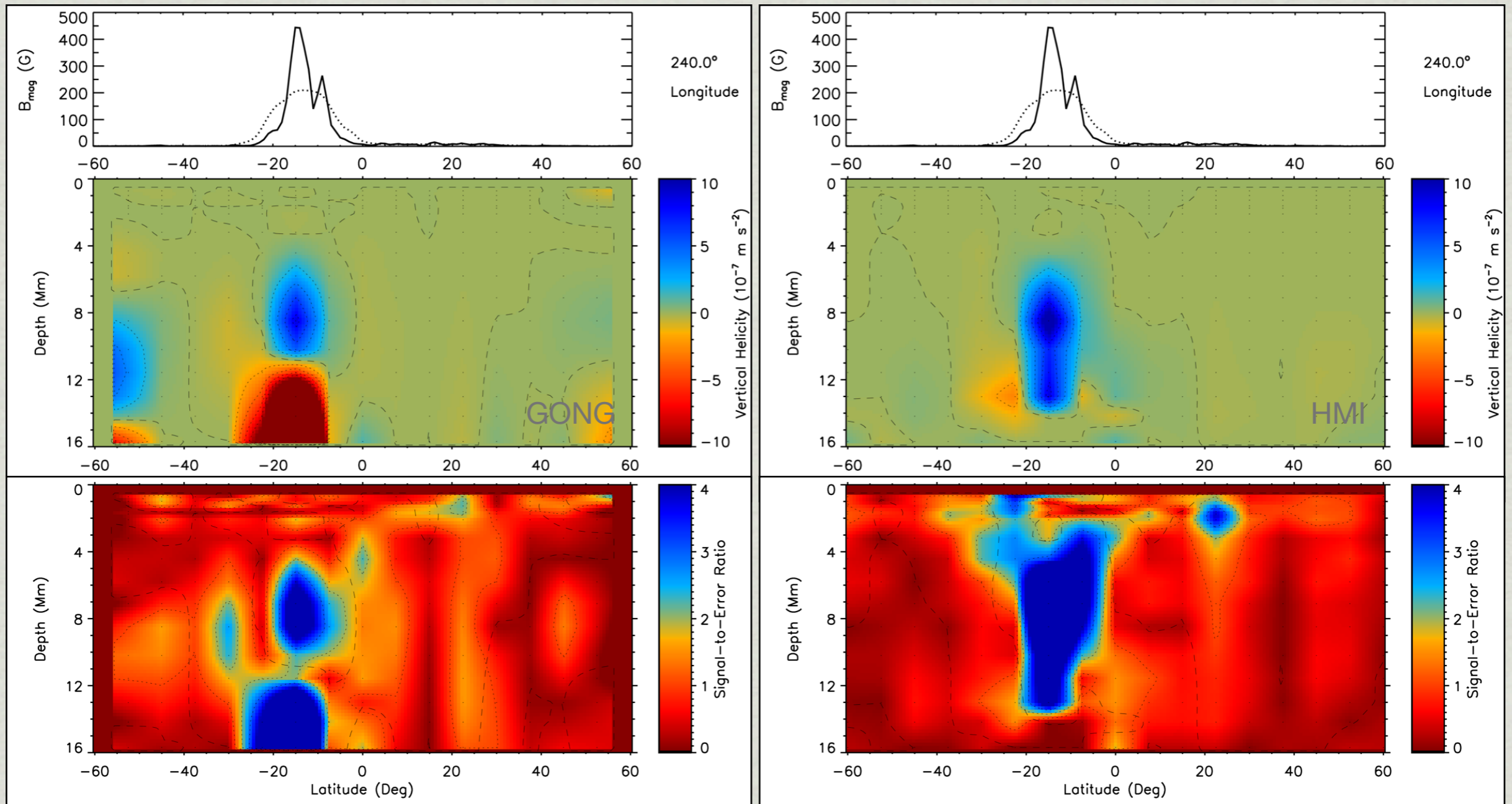
The equivalent of current helicity density.

# Vertical helicity of 10486



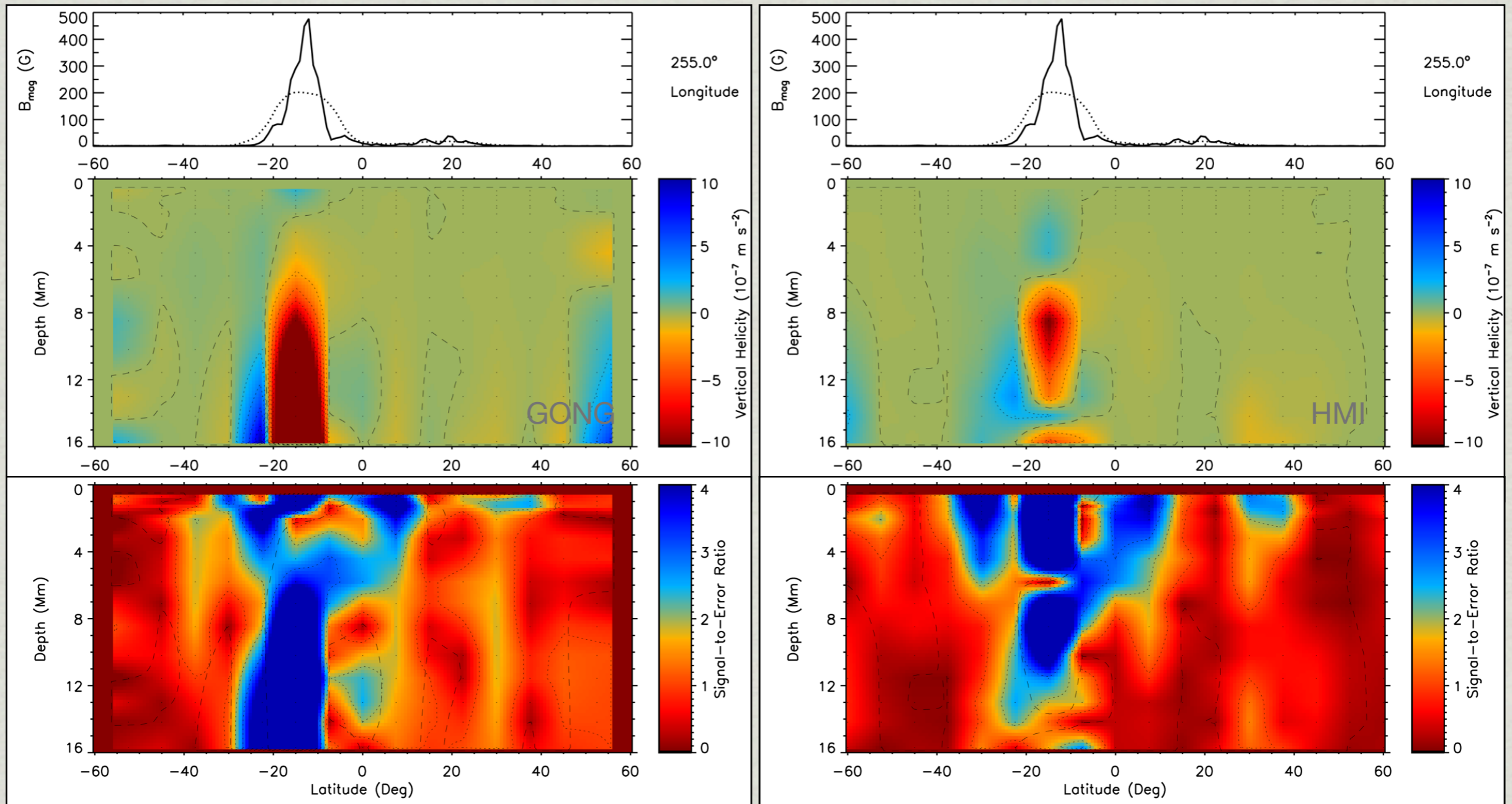
Large negative values of vertical helicity near 10486.  
Are AR 10486 and AR 10488 connected?

# Vertical helicity of 12192 (N)



Large positive values of vertical helicity near 6 -11 Mm.  
GONG and HMI are reasonably similar for large S/N.

# Vertical helicity of 12192 (P)



Large negative values of vertical helicity near 6 -11 Mm.  
GONG and HMI are reasonably similar for large S/N.

# Summary (2)

- ✱ The vertical kinetic helicity density of AR 10486 is negative (opposite of hemispheric helicity rule).
- ✱ The vertical kinetic helicity density of AR 12192 is large and its sign is different for different magnetic polarities.
- ✱ HMI and GONG helicity values are reasonably similar when the S/N ratio is high.