

Dynamics and magnetism of selected phenomena in the solar atmosphere

Martin Benko

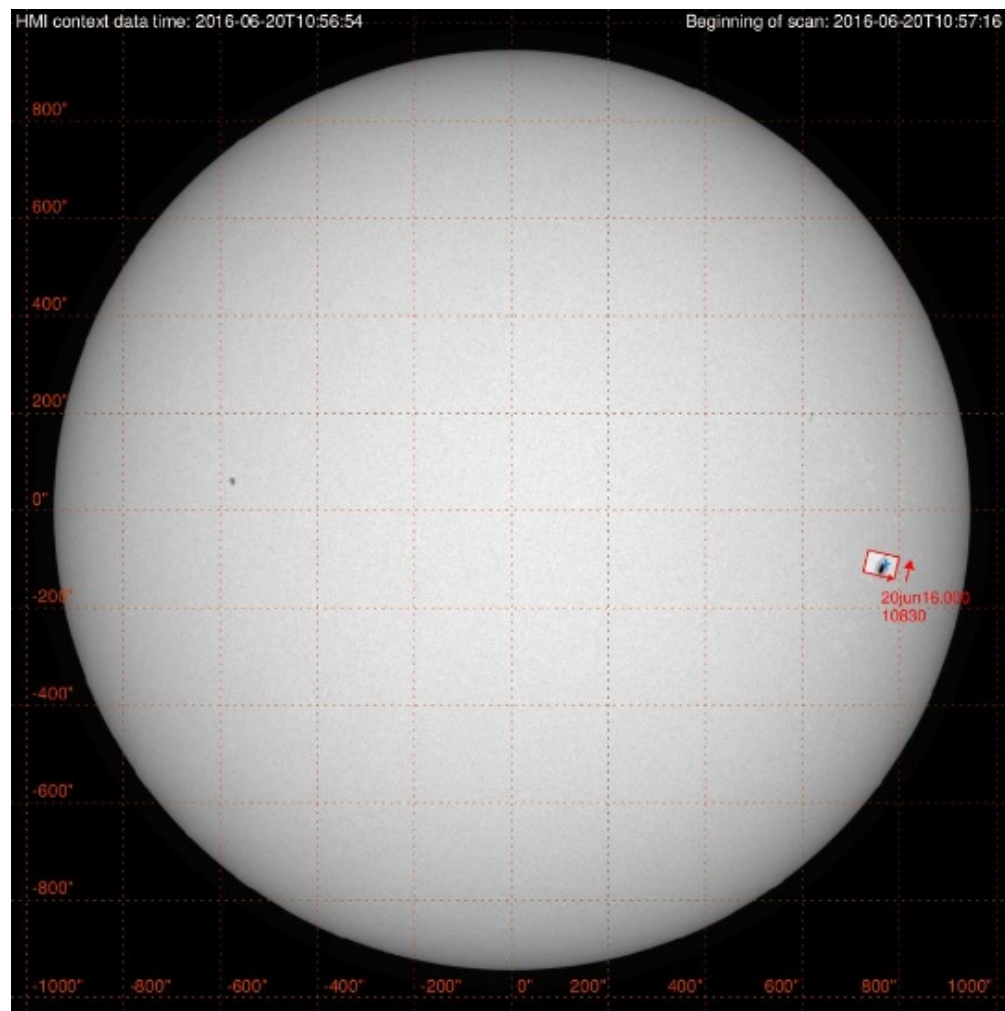
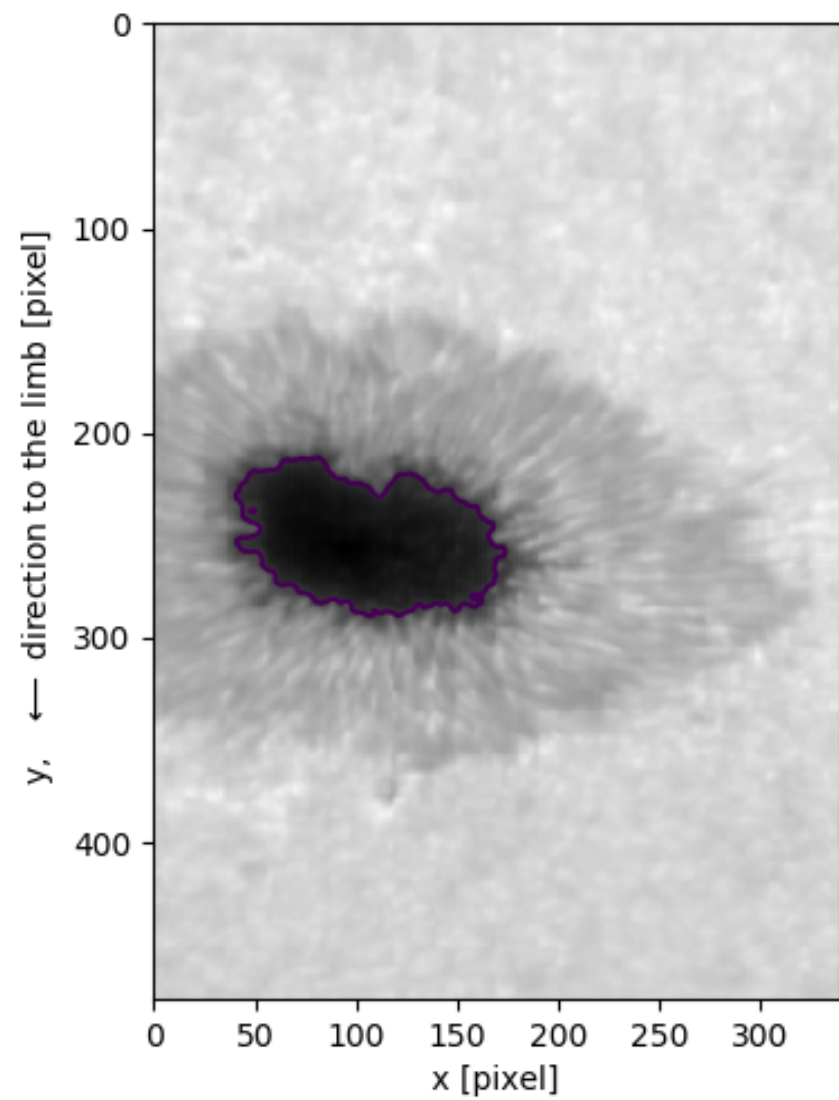
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About me



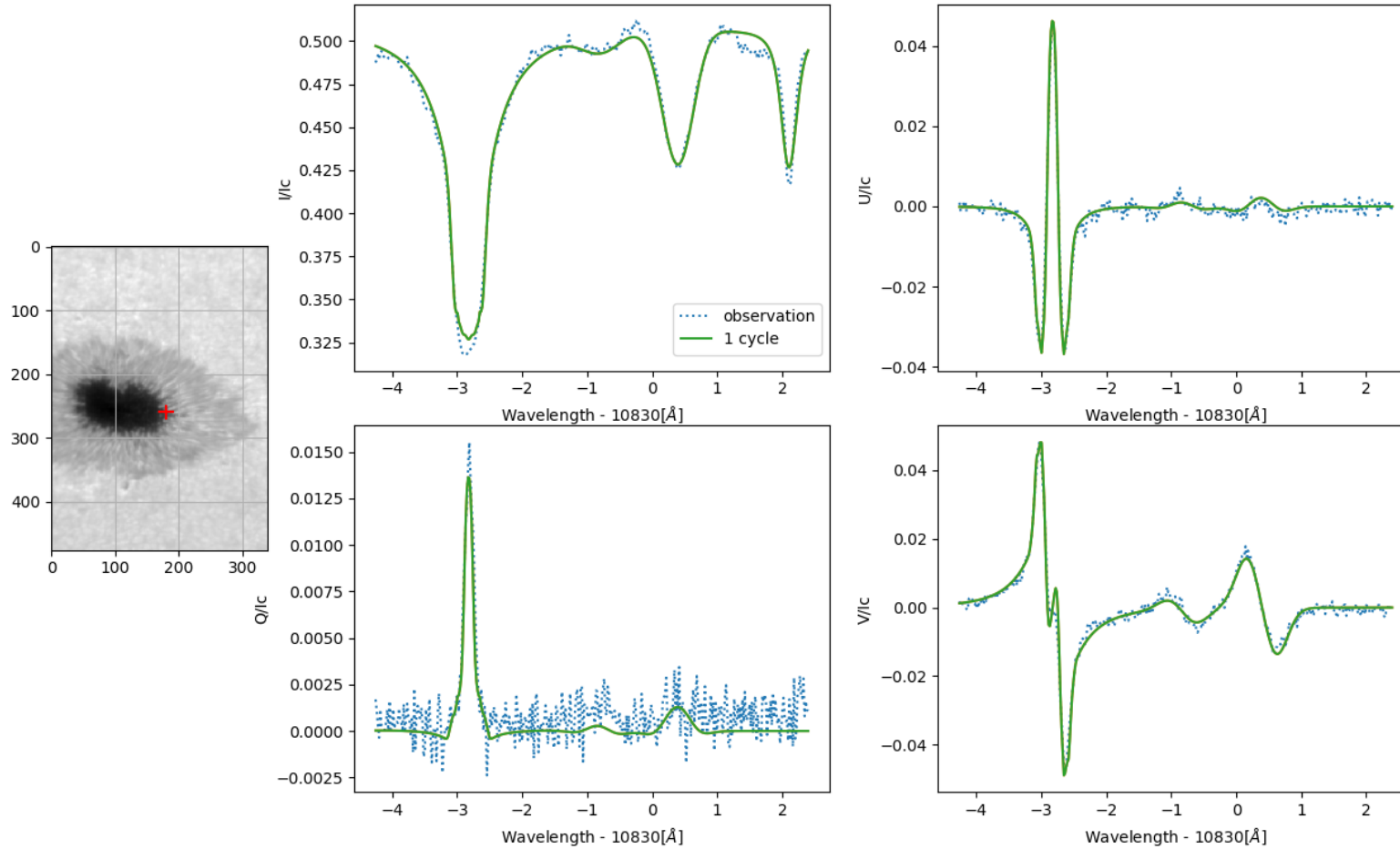
Observed data

- spectro-polarimetric data obtained with GRIS (GREGOR)
- observed area: AR 12553
- observed date: 20.6.2016
- the observed spectral region in infrared at 1 micron
- the photospheric spectral lines - Si 10827 Å and Ca 10839 Å
- the chromospheric He I 10830 Å triplet



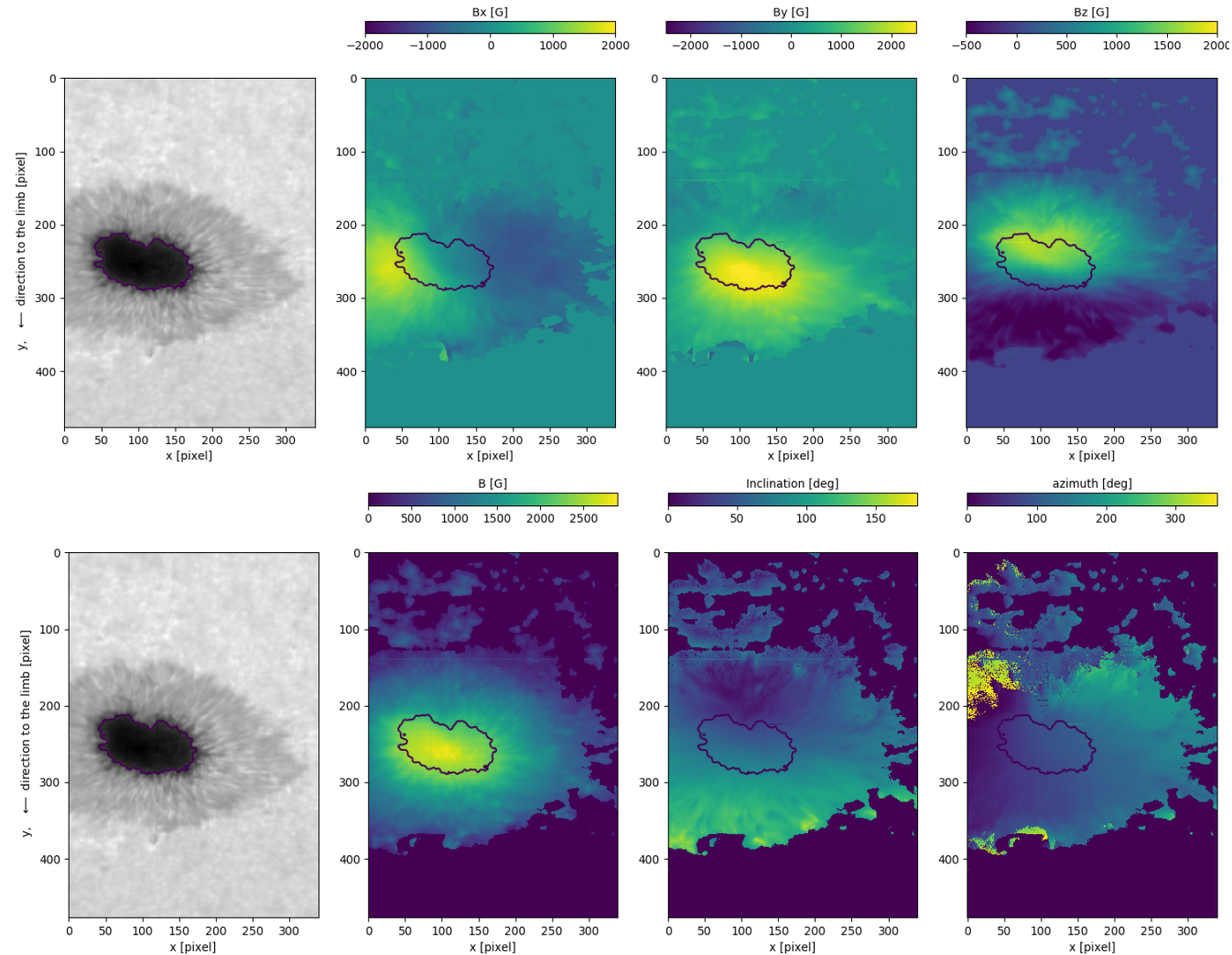
Inversion codes

- SIR – Si I 10827 Å
- HAZEL (2.0) – He I 10830 Å

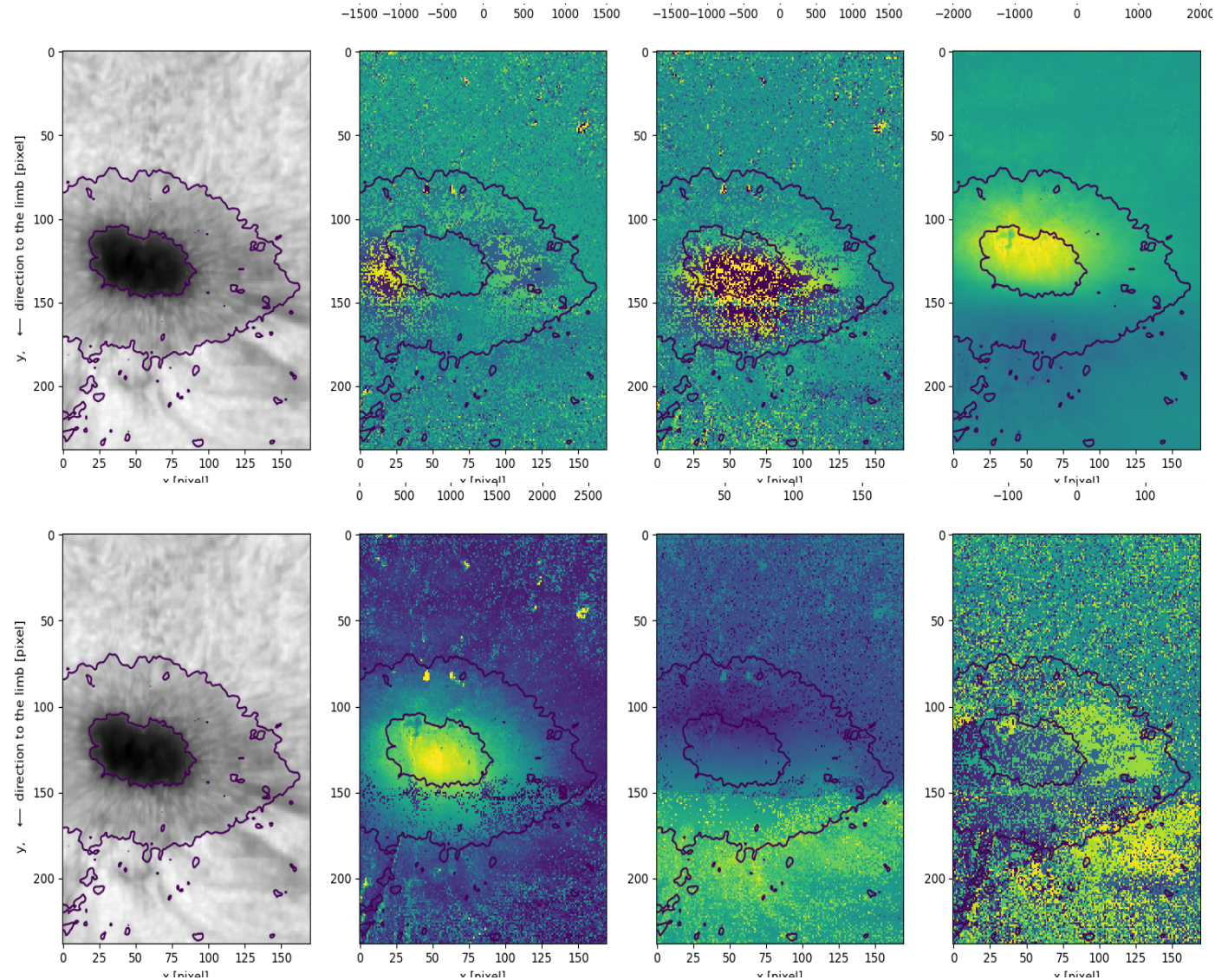


- the inferred magnetic field for chromosphere are $B_x=1120$ G, $B_y=1490$ G, $B_z=900$ G, the inferred thermal velocity is $\Delta v=7$ km s^{-1} , the doppler velocity $v=1.5$ km/s and the optical depth $\tau=0.7$
- the inferred physical parameters for photosphere are $B_x=1200$ G, $B_y=-1600$ G, $B_z=680$ G and doppler velocity is $v=2.2$ km s^{-1}

Vector of magnetic field in photosphere



Vector of magnetic field in chromosphere



Physical parameters in the photosphere and chromosphere

